# A STUDY ON THE ROTATIONAL b-FAMILY OF EQUATIONS 

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Presented to the Faculty of the Graduate School of The University of Texas at Arlington in Partial Fulfillment of the Requirements for the Degree of DOCTOR OF PHILOSOPHY

THE UNIVERSITY OF TEXAS AT ARLINGTON
May 2018

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To my parents Nezihe and Yaşar,
And my sisters Ebru, Eda, Ece

## ACKNOWLEDGEMENTS

First, I would like to express my sincere gratitude to my advisor Prof. Yue Liu for all the support and the motivation he provided, his patience and academic guidance throughout my Ph.D. study. His guidance will keep leading me the rest of my academic studies. It is him who taught me how to approach mathematical problems and to be a great mentor for students. Being his student is an honor for me.

Besides my advisor, I would like to thank Dr. Tuncay Aktosun, Dr. Ren-Cang Li, and Dr. Guojun Liao for their interest in my research and for taking time to serve in my comprehensive committee and dissertation committee.

I would like to thank the Department of Mathematics, University of Texas at Arlington for supporting me in my studies. I had a chance to work with Ting Luo and Jun Wei Sun. I am so thankful to them for the time and the effort they have invested in my studies as well.

I feel that I've made some wonderful friends so far. I am so thankful to my friends Burcu, Gül, Imelda, Mayowa, and Neslihan for their encouragement, support and time with them far away from my country during my studies in the U.S..

In addition, I would like to express my deepest thanks to my master's thesis advisor Ülkü Dinlemez for her endless inspiration and encouragement. I also thank Guilong Gui for his help in my research during his visit to UTA.

Last but not least, I would like to thank my family: my parents Nezihe and Yaşar, and my sisters Ece, Ebru, Eda, for their love, patience and support throughout
my education life and my precious cousin Ahmet who helped me survive through any kinds of language barriers I have faced.

April 10, 2018

# ABSTRACT <br> <br> A STUDY ON THE ROTATIONAL $b$-FAMILY OF EQUATIONS 

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In this thesis, we study a mathematical model of long-crested water waves propagating in one direction with the effect of Earth's rotation near the equator by following the formal asymptotic procedures. Firstly, we derive a new model equation called the rotational $b$-family of equations by using the Camassa-Holm approximation of the two-dimensional incompressible and irrotational Euler equations. Secondly, we establish that the local well-posedness of the Cauchy problem for the rotational $b$-family of equations on the Sobolev space $H^{s}$, for $s>3 / 2$. In addition, we study the effects of the Coriolis force and nonlocal higher nonlinearities on blow-up criteria and wave-breaking phenomena.

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## CHAPTER 1

## INTRODUCTION

In this thesis, we study the modelling and analysis of the rotational $b$-family of equations, such as local well-posedness, blow-up criteria, and wave breaking criteria. The early studies about shallow water equations, $b$-family of equations and rotational $b$-family of equations are reviewed in this chapter.

### 1.1 Early Studies on Shallow Water Equations

The studies on shallow water waves have been studied in the areas of geophysical fluid dynamics, physics and applied mathematics to interpret the wave behavior on the geophysical flows. Firstly, in the history of shallow water wave theory, John Scott Russell (1808-1882) explored a phenomena that he called as the wave of translation and explained in his words [39, 40]:
"I was observing the motion of a boat which was rapidly drawn along a narrow channel by a pair of horses, when the boat suddenly stopped-not so the mass of water in the channel which it had put in motion; it accumulated round the prow of the vessel in a state of violent agitation, then suddenly leaving it behind, rolled forward with great velocity, assuming the form of a large solitary elevation, a rounded, smooth and well-defined heap of water, which continued its course along the channel apparently without change of form or diminution of speed. I followed it on horseback, and overtook it still rolling on at a rate of some eight or nine miles an hour, preserving its original figure some thirty feet long and a foot to a foot and a half in height. Its height gradually diminished, and after a chase of one or two miles I lost it in the
windings of the channel. Such, in the month of August 1834, was my first chance interview with that singular and beautiful phenomenon which I have called the Wave of Translation."

After his observation, he also did an experiment to replicate his observation by building a tank and trying to repeat the solitary waves in the tank. He received successful results at the end of the experiment. However, the importance of his phenomena was not understood until the studies of D. J. Korteweg and G. de Vries in 1895. They analyzed the following equation

$$
u_{t}+u_{x}+\frac{3}{2} u u_{x}+\frac{1}{6} u_{x x x}=0
$$

which is called as the Korteweg-de Vries (KdV) equation [32] . For the KdV equation, they also derived the traveling wave solution, which confirmed "the Wave of Translation" of Russell. Later, it is called as one soliton solution.

The KdV equation and many other shallow water models known as approximations to the full Euler dynamics are only valid in the weakly nonlinear regime [32]. On the other hand, this is not enough to satisfy some physical phenomena such as wave breaking, which means that the wave remains bounded while its slope becomes unbounded in finite time [43], waves of maximum height [1, 42]. They need a transition to full nonlinearity. This leads researchers to investigate new models for nonlinear shallow water waves [43].

To describe the long-wave regime, the positive parameters $\varepsilon$ and $\mu$ are defined by respectively the amplitude parameter and the shallowness parameter as

$$
\varepsilon=\frac{a}{h_{0}}, \quad \mu=\frac{h_{0}^{2}}{\lambda^{2}}
$$

where long wavelength is $\lambda$, small amplitude is $a$ and mean level of water surface is $h_{0}$. Considering the Boussinesq regime $\varepsilon=O(\mu)$ as $\mu \ll 1$, asymptotic approximations for
the unidirectional solutions of the irrotational two-dimensional water waves problem is obtained by the KdV model $[3,13]$. However, for more accurate asymptotic approximations for these types waves which have more nonlinear behavior than dispersive, larger values of $\varepsilon$ are considered, which is called the Camassa-Holm ( CH ) scaling [11], $\varepsilon=O(\sqrt{\mu})$ as $\mu \ll 1$. Stronger nonlinear effects are obtained with CH scaling. That means the presence of breaking waves could be investigated with a higher nonlinearity [11].

Camassa and Holm derived a new completely integrable dispersive shallow water equation [4]

$$
u_{t}-u_{t x x}+\kappa u_{x}+3 u u_{x}=2 u_{x} u_{x x x}+u u_{x x x}
$$

where $u$ is the fluid velocity and $\kappa$ is a constant related to the critical shallow water wave speed. They derived the equation by using Hamiltonian methods in [4]. In addition, by using the asymptotic expansion at linear order for unidirectional shallow water waves, the integrable third-order KdV equation was obtained in [18]. On the other hand, a family of shallow water wave equations was derived at quadratic order in this asymptotic expansion and these equations are asymptotically equivalent to each other under a group of nonlinear, non-local, normal-form transformations introduced by Kodama in combination with the application of the Helmholtz-operator [18].

The $b$-family of equations is described by the following family of $1+1$ evolutionary equations which is one dimensional nonlinear waves in fluids [29]

$$
\begin{equation*}
m_{t}+\underbrace{u m_{x}}_{\text {convection }}+\underbrace{b u_{x} m}_{\text {stretching }}=0, \quad \text { with } \quad u=g * m \tag{1.1}
\end{equation*}
$$

where $u(t, x)$ is denoted as fluid velocity on the real line and vanishing at spatial infinity and the $*$ denotes the convolution, yielding

$$
u(x)=\int_{-\infty}^{\infty} g(x-y) m(y) d y
$$

which relates the velocity $u$ to the momentum density $m$ by integration against the kernel $g(x)$ over the real line. The kernel function $g(x)$ is chosen as $g(x)=\frac{1}{2} e^{-|x|}$ which implies that $m=u-u_{x x}$. The kernel $g$ and the dimensionless constant $b$ which is the ratio of stretching to convection represents the family of equations (1.1). The traveling wave shape and length scale for (1.1) are established by the function $g(x)$, while a balance or bifurcation parameter for the nonlinear solution behavior is given by the constant b.

In the equation (1.1), the quadratic terms show the balance, in fluid convection between nonlinear steeping and amplification owing to $b$-dimensional stretching. For $b \neq-1$ the equation, (1.1) can be derived as the family of asymptotically analogous shallow water wave equations that appears at quadratic-order accuracy by a proper Kodama transformation [16, 18].

In addition, the local well-posedness of $b$-family of equations is established by using Kato's semi group theory in [30]. Also, Escher and Yin obtained the local well-posedness of $b$-family of equations to get a precise blow-up scenario, to prove that the equation has strong and finite-time blow-up solutions[21]. Degasperis, Holm, and Hone[17] showed that the $b$-family of equations have the peakon solutions

$$
\begin{equation*}
u(x, t)=c e^{-|x-c t|}, c>0, \tag{1.2}
\end{equation*}
$$

besides, they have multipeakon solutions

$$
u(x, t)=\sum_{j=1}^{N} p_{k}(t) e^{-\left|x-q_{j}\right|}
$$

For any $b$, the quantities $p_{j}$ and $q_{j}$ are not canonical variables but satisfy the dynamical system

$$
\dot{p_{j}}=-(b-1) \frac{\partial G_{N}}{\partial q_{j}}, \quad \dot{q_{j}}=\frac{\partial G_{N}}{\partial p_{j}},
$$

where the overdot denotes the t-derivative and the generating function $G_{N}$ is given by

$$
G_{N}=\frac{1}{2} \sum_{j, k=1}^{N} p_{j} p_{k}(t) e^{-\left|q_{j}-q_{k}\right|}
$$

For each $b \neq 0,(1.1)$ has at least three conserved quantities as follows:

$$
\begin{aligned}
E_{1}(u) & =\int_{\mathbb{R}} m d x \\
E_{2}(u) & =\int_{\mathbb{R}} m^{1 / b} d x \\
E_{3}(u) & =\int_{\mathbb{R}} m^{-1 / b}\left(\frac{m_{x}^{2}}{b^{2} m^{2}}+1\right) d x .
\end{aligned}
$$

On the other hand, the equation (1.1) is completely integrable if $b=2$ or $b=3$. The integrability is proved by using the method of asymptotic integrability [16].

When $b=2$, the equation (1.1) is called the Camassa-Holm equation in the form

$$
\begin{equation*}
u_{t}-u_{t x x}+3 u u_{x}=2 u_{x} u_{x x x}+u u_{x x x}, \quad t>0, \quad x \in \mathbb{R}, \tag{1.3}
\end{equation*}
$$

which is the model of the unidirectional propagation of shallow water waves over a flat bottom $[4,19]$. Here $u(t, x)$ is the fluid velocity at time t in the spatial $x$ direction. (1.3) has bi-Hamiltonian structure [23, 41] and is completely integrable [4, 10]. It has also solitons as the KdV equation, while the CH equation yields permanent and breaking waves, and has peaked solitons of the form (1.2) [4, 5, 8, 36]. The Cauchy problem for (1.3) is studied thoroughly. Local well-posedness for the Camassa-Holm equation with the initial data $u_{0} \in H^{s}(\mathbb{R}), s>3 / 2$ is shown in [7, 33, 38].

When $b=3$, (1.1) is called the Degasperis-Procesi (DP) equation given by

$$
\begin{equation*}
u_{t}-u_{t x x}+4 u u_{x}=3 u_{x} u_{x x x}+u u_{x x x}, \quad t>0, \quad x \in \mathbb{R} . \tag{1.4}
\end{equation*}
$$

The formal integrability of the Degasperis-Procesi equation (1.4) is proved in [17] by setting up a Lax pair. In addition, bi-Hamiltonian structure and an infinite number
of conserved quantities for (1.4) are shown in [17]. On the other hand, similar to the Camassa- Holm equation, local well-posedness of the Degasperis-Procesi equation is proved in [44], for the initial data $u_{0} \in H^{s}(\mathbb{R}), s>3 / 2$.

The Degasperis-Procesi equation has many properties which are similar to the Camassa-Holm equation such as global strong solutions shown in $[9,7]$ for the CH equation and in [20] for the DP equation, finite time blow-up solutions, peakon solutions $u_{c}(t, x)=c e^{-|x-c t|}$ for $c>0$ [17]. In addition, the Degasperis-Procesi equation has shock peakons [34].

In the next section, some details of the new model equation are given.

### 1.2 Rotational b-Family of Equations

The rotational $b$-family (R-b-family) of equations is a model equation with the Coriolis effect from the incompressible and irrotational two-dimensional shallow water in the equatorial region in the following form

$$
\begin{align*}
u_{t}-\beta \mu u_{x x t}+c u_{x}+(b+1) \alpha \varepsilon u u_{x}-\beta_{0} \mu u_{x x x} & +\omega_{1} \varepsilon^{2} u^{2} u_{x}+\omega_{2} \varepsilon^{3} u^{3} u_{x}  \tag{1.5}\\
= & \alpha \beta \varepsilon \mu\left(b u_{x} u_{x x}+u u_{x x x}\right),
\end{align*}
$$

where the constant rotational frequency due to the Coriolis effect is defined as the parameter $\Omega$ and the wave speed is $c:=\sqrt{1+\Omega^{2}}-\Omega$. The other constants in (1.5) are defined by

$$
\begin{aligned}
& \alpha:=\frac{3 c^{2}}{(b+1)\left(1+c^{2}\right)}, \beta_{0}:=\frac{c\left(3 c^{4}+(5 b+8) c^{2}-(b+1)\right)}{9 b\left(c^{2}+1\right)^{2}}, \beta:=\frac{(b+1)\left(3 c^{4}+8 c^{2}-1\right)}{9 b\left(c^{2}+1\right)^{2}} \\
& \omega_{1}:=\frac{-3 c\left(c^{2}-1\right)\left(c^{2}-2\right)}{2\left(1+c^{2}\right)^{3}}, \text { and } \omega_{2}:=\frac{\left(c^{2}-2\right)\left(c^{2}-1\right)^{2}\left(8 c^{2}-1\right)}{2\left(1+c^{2}\right)^{5}} .
\end{aligned}
$$

The horizontal velocity field at height $z_{0}$ is represented by the solution $u$ of (1.5), and after the re-scaling, it is required that $0 \leq z_{0} \leq 1$, where

$$
\begin{equation*}
z_{0}=\sqrt{\frac{b-1}{b}-\frac{2}{3} \frac{1}{\left(c^{2}+1\right)}+\frac{(b-2) c^{2}+2 b+4}{3 b\left(c^{2}+1\right)^{2}}} \tag{1.6}
\end{equation*}
$$

As the constant $\beta$ has to be greater than 0 , it must be the case

$$
0 \leq \Omega<\sqrt{\frac{1}{6}(1+2 \sqrt{19})} \approx 1.273
$$

Let $b \geq \frac{10}{11}$ and $c_{0}=\sqrt{\frac{\sqrt{19}-4}{3}}$,

$$
\sqrt{\frac{11 b-10}{12 b}}=z_{0}(b, 1)=\inf _{c_{0}<c \leq 1} z_{0}(c) \leq z_{0} \leq \sup _{c_{0}<c \leq 1} z_{0}(c)<0.984
$$

There are two special cases for R-b-family equation, which corresponds to some wellknown equations for some different $b$ values. The first case, $b=2$ is corresponding to Rotational Camassa-Holm (R-CH) equation:

$$
\left\{\begin{array}{l}
c=1, \beta=\frac{5}{12}, \beta_{0}=\frac{1}{4}  \tag{1.7}\\
\omega_{1}=0, \omega_{2}=0, \alpha=\frac{1}{2}
\end{array} ; \Omega=0 .\right.
$$

Also, if $\Omega=0, z_{0}=\frac{1}{\sqrt{2}}$ which corresponds to the height of the case of the classical CH equation.

The second case, $b=3$ is corresponding to the Rotational Degasperis-Procesi(RDP) equation:

$$
\left\{\begin{array}{l}
c=1, \beta=\frac{10}{27}, \beta_{0}=\frac{11}{54}  \tag{1.8}\\
\omega_{1}=0, \omega_{2}=0, \quad \alpha=\frac{3}{8}
\end{array} ; \Omega=0 .\right.
$$

In addition, if $\Omega=0, z_{0}=\frac{\sqrt{23}}{6}$ which corresponds to the height of the case of classical DP equation

To derive the R-b-family of equations model in (1.5), we refer the reader to the paper [27] where the classical CH equation was derived. The R-b-family equation in (1.5) is established by showing that after a double asymptotic expansion with respect
to $\varepsilon$ and $\mu$, the free surface $\eta=\eta(\tau, \xi)$ under the field variable $(\eta, \xi)$ defined in (2.2) in 2D Euler's dynamics (2.3) (see Section 2), is governed by the equation

$$
\begin{array}{r}
2(\Omega+c) \eta_{\tau}+3 c^{2} \eta \eta_{\xi}+\frac{c^{2}}{3} \mu \eta_{\xi \xi \xi}+A_{1} \varepsilon \eta^{2} \eta_{\xi}+A_{2} \varepsilon^{2} \eta^{3} \eta_{\xi}+A_{5} \varepsilon^{3} \eta^{4} \eta_{\xi} \\
=\varepsilon \mu\left[A_{3} \eta_{\xi} \eta_{\xi \xi}+A_{4} \eta \eta_{\xi \xi \xi}\right]+O\left(\varepsilon^{4}, \mu^{2}\right)
\end{array}
$$

where the constants

$$
\begin{gathered}
A_{1}:=\frac{3 c^{2}\left(c^{2}-2\right)}{\left(c^{2}+1\right)^{2}}, \quad A_{2}:=-\frac{c^{2}\left(2-c^{2}\right)\left(c^{6}-7 c^{4}+5 c^{2}-5\right)}{\left(c^{2}+1\right)^{4}}, \\
A_{3}:=\frac{-c^{2}\left(9 c^{4}+16 c^{2}-2\right)}{3\left(c^{2}+1\right)^{2}}, \quad A_{4}:=\frac{-c^{2}\left(3 c^{4}+8 c^{2}-1\right)}{3\left(c^{2}+1\right)^{2}} \\
A_{5}:=\frac{c^{2}\left(c^{2}-2\right)\left(3 c^{10}+228 c^{8}-540 c^{6}-180 c^{4}-13 c^{2}+42\right)}{12\left(c^{2}+1\right)^{6}} .
\end{gathered}
$$

The free surface $\eta$ with respect to the horizontal component of the velocity $u$ at $z=z_{0}$ under the CH regime $\varepsilon=O(\sqrt{\mu})$ as $\mu \rightarrow 0$ is also given by

$$
\eta=\frac{1}{c} u+\gamma_{1} \varepsilon u^{2}+\gamma_{2} \varepsilon^{2} u^{3}+\gamma_{3} \varepsilon^{3} u^{4}+\gamma_{4} \varepsilon \mu u_{\xi \xi}+O\left(\varepsilon^{4}, \mu^{2}\right),
$$

where the constants in the expression are given by

$$
\begin{gathered}
\gamma_{1}=\frac{2-c^{2}}{2 c^{2}\left(c^{2}+1\right)}, \quad \gamma_{2}=\frac{\left(c^{2}-1\right)\left(c^{2}-2\right)\left(2 c^{2}+1\right)}{2 c^{3}\left(c^{2}+1\right)^{3}} \\
\gamma_{3}=-\frac{\left(c^{2}-1\right)^{2}\left(c^{2}-2\right)\left(21 c^{4}+16 c^{2}+4\right)}{8 c^{4}\left(c^{2}+1\right)^{5}}, \quad \gamma_{4}=\frac{z_{0}^{2}}{2 c}-\frac{3 c^{2}+1}{6 c\left(c^{2}+1\right)}=\frac{-\left(3 c^{4}+6 c^{2}-5\right)}{12 c\left(c^{2}+1\right)^{2}}
\end{gathered}
$$

(here the height parameter $z_{0}$ is determined by (1.6)).
The equation (1.5) can be rewritten by defining $m:=\left(1-\beta \mu \partial_{x}^{2}\right) u$, in terms of the evolution of the momentum density $m$, namely,

$$
\begin{equation*}
\partial_{t} m+\alpha \varepsilon\left(u m_{x}+b m u_{x}\right)+c u_{x}-\beta_{0} \mu u_{x x x}+\omega_{1} \varepsilon^{2} u^{2} u_{x}+\omega_{2} \varepsilon^{3} u^{3} u_{x}=0 . \tag{1.9}
\end{equation*}
$$

In the case that the Coriolis effect vanishes $(\Omega=0)$ for $b=2$, the coefficients in the higher-power nonlinearities $\omega_{1}=0$ and $\omega_{2}=0$. Using the scaling transformation
$u(t, x) \mapsto \alpha \varepsilon u(\sqrt{\beta \mu} t, \sqrt{\beta \mu} x)$ and then the Galilean transformation $u(t, x) \mapsto u(t, x-$ $\left.\frac{3}{4} t\right)+\frac{1}{4}$, the R-CH equation (1.9) is then reduced to the classical CH equation (1.3). Note that the R-b-family equation (1.9) has the following conserved quantity

$$
\begin{equation*}
I(u):=\int_{\mathbb{R}} u d x \tag{1.10}
\end{equation*}
$$

As a special case $b=2$, the equation has three conserved quantities as follows

$$
I(u)=\int_{\mathbb{R}} u d x, \quad E(u)=\frac{1}{2} \int_{\mathbb{R}}\left(u^{2}+u_{x}^{2}\right) d x
$$

and

$$
F(u)=\frac{1}{2} \int_{\mathbb{R}}\left(c u^{2}+u^{3}+\frac{\beta_{0}}{\beta} u_{x}^{2}+\frac{\omega_{1}}{6 \alpha^{2}} u^{4}+\frac{\omega_{2}}{10 \alpha^{3}} u^{5}+u u_{x}^{2}\right) d x
$$

In this thesis, due to the Coriolis effect, we obtain higher power nonlinear terms from the derivation of the rotational $b$-family of equations given in next chapter. Then, some intriguing inferences can be made for the fluid motion, especially breaking waves and the permanent waves. In addition, we analyze the effects of the Coriolis force with the Earth rotation on the appearance of the wave-breaking phenomena. Local well-posedness, blow-up and wave breaking criteria are proved for this equation in chapter 3. In chapter 4, wave breaking phenomena is investigated for the rotational Camassa-Holm equation.

## CHAPTER 2

## THE ROTATIONAL $b$-FAMILY OF EQUATIONS

### 2.1 Derivation of the Rotational b-Family of Equations Model

The derivation of $b$-Family of equations model with the Coriolis effect is given in this section. To establish this model equation, incompressible and inviscid with a constant density $\rho$ and no surface tension is considered for water flow. Also the interface between the air and the water is a free surface. Then such a motion of water flow occupying a domain $D_{t}$ in $\mathbb{R}^{3}$ under the influence of the gravity $g$ and the Coriolis force due to the Earth's rotation can be described by the Euler equations [24], viz.,[37]

$$
\left\{\begin{array}{l}
\vec{u}_{t}+(\vec{u} \cdot \nabla) \vec{u}+2 \vec{\Omega} \times \vec{u}=-\frac{1}{\rho} \nabla P+\vec{g}, \quad x \in D_{t}, \\
\nabla \cdot \vec{u}=0, \quad x \in D_{t} \\
\left.\vec{u}\right|_{t=0}=\overrightarrow{u_{0}}, \quad x \in D_{0},
\end{array}\right.
$$

where $\vec{u}=(u, v, w)^{T}$ is the fluid velocity, $P(t, x, y, z)$ is the pressure in the fluid, $\vec{g}=(0,0,-g)^{T}$ with $g \approx 9.8 \mathrm{~m} / \mathrm{s}^{2}$ the constant gravitational acceleration at the Earth's surface, and $\vec{\Omega}=\left(0, \Omega_{0} \cos \phi, \Omega_{0} \sin \phi\right)^{T}$, with the rotational frequency $\Omega_{0} \approx 73 \cdot 10^{-6} \mathrm{rad} / \mathrm{s}$ and the local latitude $\phi$, is the angular velocity vector which is directed along the axis of rotation of the rotating reference frame.

The origin of the rotating reference frame is adopted at a point on the Earth's surface with the $x$-axis, the $y$-axis and $z$-axis respectively chosen horizontally eastward, northward and upward. Also $D_{t}=\left\{(x, y, z): 0<z<h_{0}+\eta(t, x, y)\right\}$ is defined where $h_{0}$ is the typical depth of the water and $\eta(t, x, y)$ measures the deviation from the
average level. Under the $f$-plane approximation $(\sin \phi \approx 0, \phi \ll 1)$, the motion of inviscid irrotational fluid near the Equator in the region $0<z<h_{0}+\eta(t, x, y)$ with a constant density $\rho$ is described by the Euler equations $[12,24]$ in the form

$$
\left\{\begin{array}{l}
u_{t}+u u_{x}+v u_{y}+w u_{z}+2 \Omega_{0} w=-\frac{1}{\rho} P_{x}, \\
v_{t}+u v_{x}+v v_{y}+w v_{z}=-\frac{1}{\rho} P_{y} \\
w_{t}+u w_{x}+v w_{y}+w w_{z}-2 \Omega_{0} u=-\frac{1}{\rho} P_{z}-g
\end{array}\right.
$$

the incompressibility of the fluid

$$
u_{x}+v_{y}+w_{z}=0,
$$

and the irrotational condition

$$
\left(w_{y}-v_{z}, u_{z}-w_{x}, v_{x}-u_{y}\right)^{T}=(0,0,0)^{T} .
$$

The pressure is written as

$$
P(t, x, z)=P_{a}+\rho g\left(h_{0}-z\right)+p(t, x, y, z),
$$

where $P_{a}$ is the constant atmosphere pressure, and $p$ is a pressure variable measuring the hydrostatic pressure distribution.

When $P=P_{a}$ is the dynamic condition on the surface $z=h_{0}+\eta$, pressure $p$ is

$$
p=\rho g \eta .
$$

Besides, the kinematic condition on the surface $z=h_{0}+\eta(t, x, y)$ is given by

$$
w=\eta_{t}+u \eta_{x}+v \eta_{y}
$$

Lastly, considering "no-flow" condition at the flat bottom $z=0$, that is,

$$
\left.w\right|_{z=0}=0
$$

The two-dimensional flow is considered which is moving in the east-west direction along the Equator. It means that $v \equiv 0$ over the flow which is independent of the $y$ coordinate. Then the irrotational condition can be restated as $u_{z}-w_{x}=0$. Moreover, the dimensionless quantities are given

$$
x \mapsto \lambda x, \quad z \mapsto h_{0} z, \quad \eta \mapsto a \eta, \quad t \mapsto \frac{\lambda}{\sqrt{g h_{0}}} t
$$

which implies

$$
u \mapsto \sqrt{g h_{0}} u, \quad w \mapsto \sqrt{\mu g h_{0}} w, \quad p \mapsto \rho g h_{0} p .
$$

Also considering the effect of the Earth rotation, we describe

$$
\Omega=\sqrt{\frac{h_{0}}{g}} \Omega_{0}
$$

additionally, as $\varepsilon \rightarrow 0$,

$$
u \mapsto 0, \quad w \mapsto 0, \quad p \mapsto 0
$$

that is, $u, w$ and $p$ are proportional to the wave amplitude so that a scaling is required as

$$
u \mapsto \varepsilon u, \quad w \mapsto \varepsilon w, \quad p \mapsto \varepsilon p .
$$

As a result the governing equations turn into

$$
\begin{cases}u_{t}+\varepsilon\left(u u_{x}+w u_{z}\right)+2 \Omega w=-p_{x} & \text { in } 0<z<1+\varepsilon \eta(t, x)  \tag{2.1}\\ \mu\left\{w_{t}+\varepsilon\left(u w_{x}+w w_{z}\right)\right\}-2 \Omega u=-p_{z} & \text { in } 0<z<1+\varepsilon \eta(t, x) \\ u_{x}+w_{z}=0 & \text { in } 0<z<1+\varepsilon \eta(t, x) \\ u_{z}-\mu w_{x}=0 & \text { in } 0<z<1+\varepsilon \eta(t, x) \\ p=\eta & \text { on } z=1+\varepsilon \eta(t, x) \\ w=\eta_{t}+\varepsilon u \eta_{x} & \text { on } z=1+\varepsilon \eta(t, x) \\ w=0 & \text { on } z=0\end{cases}
$$

We begin with setting up a proper scale and a double asymptotic expansion to obtain equations in groups with respect to $\varepsilon$ and $\mu$ independent on each other, where $\varepsilon, \mu \ll 1$, to derive the rotational $b$-family of equations for shallow water waves. The proper far field variable is defined in terms of $\varepsilon$

$$
\begin{equation*}
\xi=\varepsilon^{1 / 2}(x-c t), \quad \tau=\varepsilon^{3 / 2} t, \quad w=\sqrt{\varepsilon} W \tag{2.2}
\end{equation*}
$$

where $c$ is the group speed of water waves [27, 28].
Applying this transformations to the governing equations (2.1), we get

$$
\begin{cases}-c u_{\xi}+\varepsilon\left(u_{\tau}+u u_{\xi}+W u_{z}\right)+2 \Omega W=-p_{\xi} & \text { in } 0<z<1+\varepsilon \eta,  \tag{2.3}\\ \varepsilon \mu\left\{-c W_{\xi}+\varepsilon\left(W_{\tau}+u W_{\xi}+W W_{z}\right)\right\}-2 \Omega u=-p_{z} & \text { in } 0<z<1+\varepsilon \eta, \\ u_{\xi}+W_{z}=0 & \text { in } 0<z<1+\varepsilon \eta, \\ u_{z}-\varepsilon \mu W_{\xi}=0 & \text { in } 0<z<1+\varepsilon \eta, \\ p=\eta & \text { on } z=1+\varepsilon \eta, \\ W=-c \eta_{\xi}+\varepsilon\left(\eta_{\tau}+u \eta_{\xi}\right) & \text { on } z=1+\varepsilon \eta, \\ W=0 & \text { on } z=0 .\end{cases}
$$

To find a solution for the system (2.3), a double asymptotic expansion is given as

$$
q \sim \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \varepsilon^{n} \mu^{m} q_{n m}
$$

as $\varepsilon \rightarrow 0, \mu \rightarrow 0$, where the each function $q_{n m}$ provides the far field conditions $q_{n m} \rightarrow 0$ as $|\xi| \rightarrow \infty$ for every $n, m=0,1,2,3, \ldots$. This expansion is considered for the scale functions $u, W, p$, and $\eta$.

Now, every coefficients of the order $O\left(\varepsilon^{i}, \mu^{j}\right)(i, j=0,1,2,3, \ldots)$ are analyzed by applying the asymptotic expansions of $u, W, p, \eta$ into (2.3).

Considering the order $O\left(\varepsilon^{0}, \mu^{0}\right)$ terms of (2.3),

$$
\begin{equation*}
f(z)=f(1)+\sum_{n=1}^{\infty} \frac{(z-1)^{n}}{n!} f^{(n)}(1) \tag{2.4}
\end{equation*}
$$

is derived from the Taylor expension that

$$
\begin{cases}-c u_{00, \xi}+2 \Omega W_{00}=-p_{00, \xi} & \text { in } \quad 0<z<1  \tag{2.5}\\ 2 \Omega u_{00}=p_{00, z} & \text { in } \quad 0<z<1, \\ u_{00, \xi}+W_{00, z}=0 & \text { in } \quad 0<z<1, \\ u_{00, z}=0 & \text { in } 0<z<1, \\ p_{00}=\eta_{00} & \text { on } z=1, \\ W_{00}=-c \eta_{00, \xi} & \text { on } z=1, \\ W_{00}=0 & \text { on } z=0\end{cases}
$$

where the expression $u_{00, \xi}$ is the derivation of $u_{00}$ with respect to $\xi$.
First of all, from the fourth equation of the (2.5), we see that $u_{00}$ is independent of $z$, which yields $u_{00}=u_{00}(\tau, \xi)$.

Nextly, from the third equation in (2.5) and the boundary condition of $W$ on $z=0$, we obtain

$$
\begin{equation*}
W_{00}=\left.W_{00}\right|_{z=0}+\int_{0}^{z} W_{00, z^{\prime}} d z^{\prime}=-\int_{0}^{z} u_{00, \xi} d z^{\prime}=-z u_{00, \xi} \tag{2.6}
\end{equation*}
$$

Additionally, considering the boundary condition of $W$ on $z=1$, it gives us the following

$$
\begin{equation*}
u_{00, \xi}(\tau, \xi)=c \eta_{00, \xi}(\tau, \xi) \tag{2.7}
\end{equation*}
$$

As a result, the following equations are obtained from (2.5)

$$
\begin{equation*}
u_{00}(\tau, \xi)=c \eta_{00}(\tau, \xi), \quad W_{00}=-c z \eta_{00, \xi}, \tag{2.8}
\end{equation*}
$$

where use has been made of the far field conditions $u_{00}, \eta_{00} \rightarrow 0$ as $|\xi| \rightarrow \infty$.
In addition, integrating the second equation of (2.5), we get

$$
\begin{equation*}
p_{00}=\left.p_{00}\right|_{z=1}+\int_{1}^{z} p_{00, z^{\prime}} d z^{\prime}=\eta_{00}+2 \Omega \int_{1}^{z} u_{00} d z^{\prime}=\eta_{00}+2 \Omega(z-1) u_{00} \tag{2.9}
\end{equation*}
$$

Then from (2.9) and (2.7), we can have

$$
\begin{equation*}
p_{00, \xi}=\left(\frac{1}{c}+2 \Omega(z-1)\right) u_{00, \xi} \tag{2.10}
\end{equation*}
$$

To conclude, the first equation of $(2.5),(2.10)$ and (2.6) provide us the following

$$
\begin{equation*}
\left(c^{2}+2 \Omega c-1\right) u_{00, \xi}=0 \tag{2.11}
\end{equation*}
$$

that means

$$
\begin{equation*}
c^{2}+2 \Omega c-1=0 \tag{2.12}
\end{equation*}
$$

if $u_{00}$ is assumed a non-trivial velocity. Consequently, if we suppose the waves flow to the right direction $(c>0)$,

$$
\begin{equation*}
c=\sqrt{1+\Omega^{2}}-\Omega \tag{2.13}
\end{equation*}
$$

is obtained.
Following the similar way to get results for the order $O\left(\varepsilon^{0}, \mu^{1}\right), O\left(\varepsilon^{2}, \mu^{0}\right)$, $O\left(\varepsilon^{1}, \mu^{1}\right), O\left(\varepsilon^{3}, \mu^{0}\right), O\left(\varepsilon^{4}, \mu^{0}\right)$, and $O\left(\varepsilon^{2}, \mu^{1}\right)$ terms of (2.3) respectively, we have

$$
\begin{gather*}
u_{01}=c \eta_{01}=c \eta_{01}(\tau, \xi)  \tag{2.14}\\
u_{20}=u_{20}(\tau, \xi)=c \eta_{20}-2\left(c+c_{1}\right) \eta_{00} \eta_{10}-\frac{2 c_{1}-3 \Omega}{3(c+\Omega)}\left(c+c_{1}\right) \eta_{00}^{3},  \tag{2.15}\\
u_{11}=u_{11}(\tau, \xi)=\left(\frac{c}{6}-\frac{2 c_{1}}{9}-\frac{c}{2} z^{2}\right) \eta_{00, \xi \xi}+c \eta_{11}-2\left(c+c_{1}\right) \eta_{00} \eta_{01}, \\
u_{30}=u_{30}(\tau, \xi)=c \eta_{30}-2\left(c+c_{1}\right)\left(\eta_{00} \eta_{20}\right)-\left(c+c_{1}\right)\left(\eta_{10}^{2}\right)-\frac{2 c_{1}-3 \Omega}{\Omega+c}\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{10}\right) \\
-\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{24(c+\Omega)^{2}}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right), \tag{2.16}
\end{gather*}
$$

$$
2(c+\Omega) \eta_{30, \tau}+3 c^{2}\left(\eta_{00} \eta_{30}+\eta_{10} \eta_{20}\right)_{\xi}-2\left(3 c+2 c_{1}\right)\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{20}+\eta_{00} \eta_{10}^{2}\right)_{\xi}
$$

$$
\begin{equation*}
-\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}-15\right)}{3(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{3} \eta_{10}\right)_{\xi}-B_{2}\left(\eta_{00}^{5}\right)_{\xi}=0 \tag{2.17}
\end{equation*}
$$

$$
\begin{align*}
& 2(\Omega+c) \eta_{11, \tau}+3 c^{2}\left(\eta_{00} \eta_{11}+\eta_{10} \eta_{01}\right)_{\xi}-2\left(c+c_{1}\right)\left(3 c+2 c_{1}\right)\left(\eta_{00}^{2} \eta_{01}\right)_{\xi}+\frac{c^{2}}{3} \eta_{10, \xi \xi \xi} \\
& -\left(\frac{c^{2}}{6}+\frac{10 c c_{1}}{9}+\frac{2 c_{1}^{2}}{9}\right)\left(\eta_{00, \xi}^{2}\right)_{\xi}-\left(\frac{c^{2}}{3}+\frac{20 c c_{1}}{9}+\frac{8 c_{1}^{2}}{9}\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi}=0 \tag{2.18}
\end{align*}
$$

with

$$
\begin{gather*}
c_{1}:=-\frac{3 c^{2}}{4(\Omega+c)}=-\frac{3 c^{3}}{2\left(c^{2}+1\right)},  \tag{2.19}\\
B_{1}:=\frac{\left(c+c_{1}\right)^{2}\left(82 c c_{1}+36 c_{1}^{2}+45 c^{2}-18 \Omega c_{1}-27 \Omega c-15\right)}{3(\Omega+c)^{2}} \\
+\frac{c_{1}\left(c+c_{1}\right)\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{3(\Omega+c)^{2}}, \\
B_{2}:=\frac{1}{5} B_{1}-\frac{\left(c+c_{1}\right)^{2}\left(2 c_{1}-3 \Omega\right)}{3(\Omega+c)}+\frac{2 c\left(c+c_{1}\right)\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{12(\Omega+c)^{2}} \\
=\frac{c^{2}\left(2-c^{2}\right)\left(3 c^{10}+228 c^{8}-540 c^{6}-180 c^{4}-13 c^{2}+42\right)}{60\left(c^{2}+1\right)^{6}} .
\end{gather*}
$$

Details for the each order can be found in next chapter. After analyzing all orders, we consider $\eta$ as the following

$$
\begin{equation*}
\eta:=\eta_{00}+\varepsilon \eta_{10}+\varepsilon^{2} \eta_{20}+\varepsilon^{3} \eta_{30}+\mu \eta_{01}+\varepsilon \mu \eta_{11}+O\left(\varepsilon^{4}, \mu^{2}\right) . \tag{2.20}
\end{equation*}
$$

Multiplying the equations (2.38), (2.50), (2.61), (2.71), (2.80), and (2.90) by $1, \varepsilon, \mu$, $\varepsilon^{2}, \varepsilon^{3}$, and $\varepsilon \mu$, respectively, and considering (2.20), we obtain the equation of $\eta$ up to the order $O\left(\varepsilon^{4}, \mu^{2}\right)$ that

$$
\begin{align*}
& 2(\Omega+c) \eta_{\tau}+3 c^{2} \eta \eta_{\xi}+\frac{c^{2}}{3} \mu \eta_{\xi \xi \xi}+\varepsilon A_{1} \eta^{2} \eta_{\xi}+\varepsilon^{2} A_{2} \eta^{3} \eta_{\xi}+A_{0} \varepsilon^{3} \eta^{4} \eta_{\xi} \\
& =\varepsilon \mu\left(A_{3} \eta_{\xi} \eta_{\xi \xi}+A_{4} \eta \eta_{\xi \xi \xi}\right)+O\left(\varepsilon^{4}, \mu^{2}\right) \tag{2.21}
\end{align*}
$$

where $c_{1}=-\frac{3 c^{3}}{2\left(c^{2}+1\right)}$ is defined in (2.39),

$$
A_{0}=-5 B_{2}:=\frac{c^{2}\left(c^{2}-2\right)\left(3 c^{10}+228 c^{8}-540 c^{6}-180 c^{4}-13 c^{2}+42\right)}{12\left(c^{2}+1\right)^{6}}
$$

$$
\begin{gathered}
A_{1}:=-2\left(3 c+2 c_{1}\right)\left(c+c_{1}\right)=\frac{3 c^{2}\left(c^{2}-2\right)}{\left(c^{2}+1\right)^{2}} \\
A_{2}:=-\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}-15\right)}{3(c+\Omega)}\left(c+c_{1}\right)=-\frac{c^{2}\left(2-c^{2}\right)\left(c^{6}-7 c^{4}+5 c^{2}-5\right)}{\left(c^{2}+1\right)^{4}}, \\
A_{3}:=\frac{2 c^{2}}{3}+\frac{40 c c_{1}}{9}+\frac{4 c_{1}^{2}}{3}=\frac{-c^{2}\left(9 c^{4}+16 c^{2}-2\right)}{3\left(c^{2}+1\right)^{2}}, \quad A_{4}:=\frac{c^{2}}{3}+\frac{20 c c_{1}}{9}+\frac{8 c_{1}^{2}}{9}=\frac{-c^{2}\left(3 c^{4}+8 c^{2}-1\right)}{3\left(c^{2}+1\right)^{2}} .
\end{gathered}
$$

Furthermore, we have the followings from analyzing all orders

$$
\begin{gathered}
u_{00}=c \eta_{00} \\
u_{10}=c \eta_{10}-\left(c_{1}+c\right) \eta_{00}^{2} \\
u_{01}=c \eta_{01} \\
u_{11}=c \eta_{11}-2\left(c_{1}+c\right) \eta_{00} \eta_{01}+\left(\frac{c}{6}-\frac{2 c_{1}}{9}-\frac{c z^{2}}{2}\right) \eta_{00, \xi \xi}, \\
u_{20}=c \eta_{20}-2\left(c+c_{1}\right)\left(\eta_{00} \eta_{10}\right)-\frac{2 c_{1}-3 \Omega}{3(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{3}\right), \\
u_{30}=c \eta_{30}-2\left(c+c_{1}\right)\left(\eta_{00} \eta_{20}\right)-\left(c+c_{1}\right)\left(\eta_{10}^{2}\right)-\frac{2 c_{1}-3 \Omega}{\Omega+c}\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{10}\right) \\
-\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{24(c+\Omega)^{2}}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right)
\end{gathered}
$$

Then we obtain

$$
\begin{aligned}
& \eta_{00}=\frac{1}{c} u_{00}, \eta_{10}=\frac{1}{c} u_{10}+\gamma_{1} u_{00}^{2}, \eta_{01}=\frac{1}{c} u_{01}, \eta_{20}=\frac{1}{c} u_{20}+2 \gamma_{1} u_{00} u_{10}+\gamma_{2} u_{00}^{3} \\
& \eta_{30}=\frac{1}{c} u_{30}+\gamma_{1} u_{10}^{2}+2 \gamma_{1} u_{00} u_{20}+3 \gamma_{2} u_{00}^{2} u_{10}+\gamma_{3} u_{00}^{4}, \\
& \eta_{11}=\frac{1}{c} u_{11}+2 \gamma_{1} u_{00} u_{01}+\gamma_{4} u_{00, \xi \xi},
\end{aligned}
$$

where

$$
\begin{gathered}
\gamma_{1} \stackrel{\text { def }}{=} \frac{c_{1}+c}{c^{3}}, \\
\gamma_{2} \stackrel{\text { def }}{=} \frac{2\left(c+c_{1}\right)^{2}}{c^{5}}+\frac{\left(2 c_{1}-3 \Omega\right)\left(c+c_{1}\right)}{3 c^{4}(c+\Omega)}, \\
\gamma_{3} \stackrel{\text { def }}{=} \frac{5\left(c+c_{1}\right)^{3}}{c^{7}}+\frac{5\left(2 c_{1}-3 \Omega\right)\left(c+c_{1}\right)^{2}}{3 c^{6}(c+\Omega)}+\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{24 c^{5}(c+\Omega)^{2}}\left(c+c_{1}\right),
\end{gathered}
$$

$$
\gamma_{4} \stackrel{\text { def }}{=}-\left(\frac{1}{6 c}-\frac{2 c_{1}}{9 c^{2}}-\frac{z^{2}}{2 c}\right),
$$

or it is the same,

$$
\begin{align*}
& \gamma_{1}=\frac{2-c^{2}}{2 c^{2}\left(c^{2}+1\right)}, \quad \gamma_{2}=\frac{\left(c^{2}-1\right)\left(c^{2}-2\right)\left(2 c^{2}+1\right)}{2 c^{3}\left(c^{2}+1\right)^{3}}  \tag{2.22}\\
& \gamma_{3}=-\frac{\left(c^{2}-1\right)^{2}\left(c^{2}-2\right)\left(21 c^{4}+16 c^{2}+4\right)}{8 c^{4}\left(c^{2}+1\right)^{5}}, \quad \gamma_{4}=\frac{z^{2}}{2 c}-\frac{3 c^{2}+1}{6 c\left(c^{2}+1\right)}
\end{align*}
$$

Therefore, we can rewrite $\eta$ with respect to $u$,

$$
\begin{aligned}
\eta & =\eta_{00}+\varepsilon \eta_{10}+\varepsilon^{2} \eta_{20}+\mu \eta_{01}+\varepsilon^{3} \eta_{30}+\varepsilon \mu \eta_{11}+O\left(\varepsilon^{4}, \mu^{2}\right) \\
& =\frac{1}{c} u_{00}+\varepsilon\left(\frac{1}{c} u_{10}+\gamma_{1} u_{00}^{2}\right)+\varepsilon^{2}\left(\frac{1}{c} u_{20}+2 \gamma_{1} u_{00} u_{10}+\gamma_{2} u_{00}^{3}\right) \\
& +\mu \frac{1}{c} u_{01}+\varepsilon \mu\left(\frac{1}{c} u_{11}+2 \gamma_{1} u_{00} u_{01}+\gamma_{4} u_{00, \xi \xi}\right) \\
& +\varepsilon^{3}\left(\frac{1}{c} u_{30}+\gamma_{1} u_{10}^{2}+2 \gamma_{1} u_{00} u_{20}+3 \gamma_{2} u_{00}^{2} u_{10}+\gamma_{3} u_{00}^{4}\right)+O\left(\varepsilon^{4}, \mu^{2}\right) .
\end{aligned}
$$

and consider

$$
u=u_{00}+\varepsilon u_{10}+\varepsilon^{2} u_{20}+\mu u_{01}+\varepsilon^{3} u_{30}+\varepsilon \mu u_{11}+O\left(\varepsilon^{4}, \mu^{2}\right),
$$

so we obtain

$$
\begin{equation*}
\eta=\frac{1}{c} u+\gamma_{1} \varepsilon u^{2}+\gamma_{2} \varepsilon^{2} u^{3}+\gamma_{3} \varepsilon^{3} u^{4}+\gamma_{4} \varepsilon \mu u_{\xi \xi}+O\left(\varepsilon^{4}, \mu^{2}\right), \tag{2.23}
\end{equation*}
$$

where $\gamma_{i}(i=1,2,3,4)$ are defined in (2.22) and the parameter $z \in[0,1]$.
The equation (2.23) gives us a result that the free surface $\eta$ and the horizontal velocity $u$ do not have the relation with Coriolis effect. Also, it states that we can derive other water wave models, like the classical KdV equation, the BBM equation. and the (improved) Boussinesq equation, from relation (2.23) in the KdV regime $\varepsilon=O(\mu)$. Now, we find the each term of the equation (2.21) with respect to $u$ by using (2.23), so we have

$$
\begin{align*}
2(\Omega+c) \eta_{\tau}= & \frac{2(\Omega+c)}{c} u_{\tau}+\frac{2(\Omega+c)\left(c_{1}+c\right)}{c^{3}} \varepsilon\left(u^{2}\right)_{\tau}+2(\Omega+c) \gamma_{2} \varepsilon^{2}\left(u^{3}\right)_{\tau}  \tag{2.24}\\
& +2(\Omega+c) \gamma_{3} \varepsilon^{3}\left(u^{4}\right)_{\tau}+2(\Omega+c) \gamma_{4} \varepsilon \mu u_{\tau \xi \xi}+O\left(\varepsilon^{4}, \mu^{2}\right),
\end{align*}
$$

as $\varepsilon, \mu \rightarrow 0$

$$
\begin{gathered}
3 c^{2} \eta \eta_{\xi}=\frac{3 c^{2}}{2}\left(\left(\frac{1}{c} u+\frac{c_{1}+c}{c^{3}} \varepsilon u^{2}+\gamma_{2} \varepsilon^{2} u^{3}+\gamma_{3} \varepsilon^{3} u^{4}\right)^{2}+\gamma_{4} \varepsilon \mu u_{\xi \xi}\right)_{\xi}+O\left(\varepsilon^{4}, \mu^{2}\right) \\
=\frac{3 c^{2}}{2}\left(\frac{1}{c^{2}} u^{2}+\frac{2\left(c_{1}+c\right)}{c^{4}} \varepsilon u^{3}+\left(\frac{\left(c_{1}+c\right)^{2}}{c^{6}}+\frac{2}{c} \gamma_{2}\right) \varepsilon^{2} u^{4}+\frac{2}{c} \gamma_{4} \mu \varepsilon u u_{\xi \xi}\right. \\
\\
\left.+\left(\frac{2}{c} \gamma_{3}+\frac{2\left(c_{1}+c\right)}{c^{3}} \gamma_{2}\right) \varepsilon^{3} u^{5}\right)_{\xi}+O\left(\varepsilon^{4}, \mu^{2}\right) \\
\\
\frac{c^{2}}{3} \mu \eta_{\xi \xi \xi}=\frac{c^{2}}{3} \mu\left(\frac{1}{c} u+\frac{c_{1}+c}{c^{3}} \varepsilon u^{2}\right)_{\xi \xi \xi}+O\left(\varepsilon^{4}, \mu^{2}\right) \\
\varepsilon \mu\left(A_{3} \eta_{\xi} \eta_{\xi \xi}+A_{4} \eta \eta_{\xi \xi \xi}\right)=\varepsilon \mu\left(\frac{A_{3}}{c^{2}} u_{\xi} u_{\xi \xi}+\frac{A_{4}}{c^{2}} u u_{\xi \xi \xi}\right)+O\left(\varepsilon^{4}, \mu^{2}\right) \\
A_{1} \varepsilon \eta^{2} \eta_{\xi}= \\
\frac{A_{1}}{3} \varepsilon\left[\frac{1}{c^{3}} u^{3}+\frac{3\left(c_{1}+c\right)}{c^{5}} \varepsilon u^{4}+\left(\frac{3\left(c_{1}+c\right)^{2}}{c^{7}}+\frac{3}{c^{2}} \gamma_{2}\right) \varepsilon^{2} u^{5}\right]_{\xi}+O\left(\varepsilon^{4}, \mu^{2}\right), \\
A_{2} \varepsilon^{2} \eta^{3} \eta_{\xi}=\frac{A_{2}}{4 c^{4}} \varepsilon^{2}\left(u^{4}\right)_{\xi}+\frac{A_{2}\left(c_{1}+c\right)}{c^{6}} \varepsilon^{3}\left(u^{5}\right)_{\xi}+O\left(\varepsilon^{4}, \mu^{2}\right)
\end{gathered}
$$

and

$$
-5 B_{2} \varepsilon^{3} \eta^{4} \eta_{\xi}=-\frac{B_{2}}{c^{5}} \varepsilon^{3}\left(u^{5}\right)_{\xi}+O\left(\varepsilon^{4}, \mu^{2}\right)
$$

Therefore, the equation (2.21) turns out

$$
\begin{align*}
u_{\tau} & +\frac{2\left(c_{1}+c\right)}{c^{2}} \varepsilon u u_{\tau}+3 \gamma_{2} c \varepsilon^{2} u^{2} u_{\tau}+\gamma_{4} c \varepsilon \mu u_{\tau \xi \xi}+4 \gamma_{3} c \varepsilon^{3} u^{3} u_{\tau}+\frac{3 c}{2(\Omega+c)} u u_{\xi} \\
& +\frac{c A_{5}}{2(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi}+\frac{c A_{6}}{2(\Omega+c)} \varepsilon u^{2} u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u_{\xi \xi \xi}+\frac{c A_{7}}{2(\Omega+c)} \varepsilon^{3} u^{4} u_{\xi}  \tag{2.25}\\
& +\left(\frac{c A_{8}}{2(\Omega+c)} u_{\xi} u_{\xi \xi}+\frac{c A_{9}}{2(\Omega+c)} u u_{\xi \xi \xi}\right) \varepsilon \mu=O\left(\varepsilon^{4}, \varepsilon^{2} \mu, \mu^{2}\right),
\end{align*}
$$

where

$$
\begin{gathered}
A_{5}:=\frac{6\left(c_{1}+c\right)^{2}}{c^{4}}+12 c \gamma_{2}+\frac{4 A_{1}\left(c_{1}+c\right)}{c^{5}} \varepsilon^{2}+\frac{A_{2}}{c^{4}}, \\
A_{6}:=\frac{9\left(c_{1}+c\right)}{c^{2}}+\frac{A_{1}}{c^{3}}, \\
A_{8}:=3 c \gamma_{4}+\frac{2\left(c_{1}+c\right)}{c}-\frac{A_{3}}{c^{2}}, \\
A_{9}:=3 c \gamma_{4}+\frac{2\left(c_{1}+c\right)}{3 c}-\frac{A_{4}}{c^{2}},
\end{gathered}
$$

$$
A_{7}:=5\left[\frac{3}{2} c^{2}\left(\frac{2}{c} \gamma_{3}+\frac{2\left(c_{1}+c\right)}{c^{3}} \gamma_{2}\right)+\frac{A_{1}}{3}\left(\frac{3}{c^{7}}\left(c_{1}+c\right)^{2}+\frac{3}{c^{2}} \gamma_{2}\right)+\frac{A_{2}\left(c_{1}+c\right)}{c^{6}}-\frac{B_{2}}{c^{5}}\right] .
$$

Multiplying the equation (2.25) by $\varepsilon u$, we have

$$
\begin{gathered}
\varepsilon u u_{\tau}=-\varepsilon u\left(\frac{2\left(c_{1}+c\right)}{c^{2}} \varepsilon u u_{\tau}+3 \gamma_{2} c \varepsilon^{2} u^{2} u_{\tau}+\frac{3 c}{2(\Omega+c)} u u_{\xi}+\frac{c A_{5}}{2(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi}\right. \\
\left.+\frac{c A_{6}}{2(\Omega+c)} \varepsilon u^{2} u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u_{\xi \xi \xi}\right)+O\left(\varepsilon^{4}, \varepsilon^{2} \mu, \mu^{2}\right),
\end{gathered}
$$

which implies

$$
\begin{aligned}
\varepsilon u\left(1+\frac{2\left(c_{1}+c\right)}{c^{2}} \varepsilon u\right. & \left.+3 \gamma_{2} c \varepsilon^{2} u^{2}\right) u_{\tau}=-\varepsilon u\left(\frac{3 c}{2(\Omega+c)} u u_{\xi}+\frac{c A_{5}}{2(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi}\right. \\
& \left.+\frac{c A_{6}}{2(\Omega+c)} \varepsilon u^{2} u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u_{\xi \xi \xi}\right)+O\left(\varepsilon^{4}, \varepsilon^{2} \mu, \mu^{2}\right) .
\end{aligned}
$$

From the last equation, we obtain

$$
\begin{aligned}
\varepsilon u u_{\tau}= & -\varepsilon u\left[1-\left(\frac{2\left(c_{1}+c\right)}{c^{2}} \varepsilon u+3 \gamma_{2} c \varepsilon^{2} u^{2}\right)+\left(\frac{2\left(c_{1}+c\right)}{c^{2}} \varepsilon u\right)^{2}\right]\left[\frac{3 c}{2(\Omega+c)} u u_{\xi}\right. \\
& \left.+\frac{c A_{5}}{2(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi}+\frac{c A_{6}}{2(\Omega+c)} \varepsilon u^{2} u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u_{\xi \xi \xi}\right]+O\left(\varepsilon^{4}, \mu^{2}\right)
\end{aligned}
$$

and then

$$
\begin{align*}
\varepsilon u u_{\tau}= & -\varepsilon u\left[\frac{3 c}{2(\Omega+c)} u u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u_{\xi \xi \xi}+\frac{c^{2} A_{6}-6\left(c_{1}+c\right)}{2 c(\Omega+c)} \varepsilon u^{2} u_{\xi}\right. \\
& \left.+\frac{c^{2} A_{5}-2 A_{6}\left(c_{1}+c\right)+3 c^{2}\left(\frac{4\left(c_{1}+\right)^{2}}{c^{4}}-3 \gamma_{2} c\right)}{2 c(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi}\right]+O\left(\varepsilon^{4}, \mu^{2}\right),  \tag{2.26}\\
\varepsilon^{2} u^{2} u_{\tau}= & -\varepsilon^{2} u^{2}\left[\frac{3 c}{2(\Omega+c)} u u_{\xi}+\frac{c^{2} A_{6}-6\left(c_{1}+c\right)}{2 c(\Omega+c)} \varepsilon u^{2} u_{\xi}\right]+O\left(\varepsilon^{4}, \varepsilon^{2} \mu, \mu^{2}\right), \\
\varepsilon^{3} u^{3} u_{\tau}= & -\frac{3 c}{2(\Omega+c)} \varepsilon^{3} u^{4} u_{\xi}+O\left(\varepsilon^{4}, \mu^{2}\right), \quad \varepsilon \mu u_{\tau \xi \xi}=-\frac{3 c}{2(\Omega+c)} \varepsilon \mu\left(u u_{\xi}\right)_{\xi \xi}+O\left(\varepsilon^{4}, \mu^{2}\right) \tag{2.27}
\end{align*}
$$

In (2.27), we use the decompositon of the term $\varepsilon \mu u_{\tau \xi \xi}$, which is $\varepsilon \mu(1-\nu) u_{\tau \xi \xi}+\varepsilon \mu \nu u_{\tau \xi \xi}$ for some constant $\nu$ (to be determined later), as follows

$$
\begin{equation*}
\varepsilon \mu u_{\tau \xi \xi}=\varepsilon \mu(1-\nu) u_{\tau \xi \xi}-\frac{3 c \nu}{2(\Omega+c)} \varepsilon \mu\left(u u_{\xi}\right)_{\xi \xi}+O\left(\varepsilon^{4}, \mu^{2}\right) \tag{2.28}
\end{equation*}
$$

Considering (2.26)-(2.28), we can rewrite (2.25)

$$
\begin{aligned}
u_{\tau} & +c \gamma_{4}(1-\nu) \mu \varepsilon u_{\tau \xi \xi}+\frac{3 c}{2(\Omega+c)} u u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u_{\xi \xi \xi}-\frac{9 c^{2} \gamma_{2}}{2(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi} \\
& -\frac{3 c^{2} \gamma_{4} \nu}{2(\Omega+c)} \mu \varepsilon\left(u u_{\xi}\right)_{\xi \xi}+\frac{2\left(c_{1}+c\right)}{c^{2}} \varepsilon\left[\frac{3 c}{2(\Omega+c)} u^{2} u_{\xi}+\frac{c^{2}}{6(\Omega+c)} \mu u u_{\xi \xi \xi}\right. \\
& \left.+\frac{c^{2} A_{6}-6\left(c_{1}+c\right)}{2 c(\Omega+c)} \varepsilon u^{3} u_{\xi}\right]+\frac{c A_{5}}{2(\Omega+c)} \varepsilon^{2} u^{3} u_{\xi}+\frac{c A_{6}}{2(\Omega+c)} \varepsilon u^{2} u_{\xi} \\
& +\mu \varepsilon\left(\frac{c A_{8}}{2(\Omega+c)} u_{\xi} u_{\xi \xi}+\frac{c A_{9}}{2(\Omega+c)} u u_{\xi \xi \xi}\right)+A_{10} \varepsilon^{3} u^{4} u_{\xi}=O\left(\varepsilon^{4}, \mu^{2}\right),
\end{aligned}
$$

where

$$
\begin{aligned}
A_{10}:= & \frac{c A_{7}}{2(\Omega+c)}-\frac{\left(c_{1}+c\right)\left(c^{2} A_{5}-2 A_{6}\left(c_{1}+c\right)+3 c^{2}\left(\frac{4\left(c_{1}+c\right)^{2}}{c^{4}}-3 \gamma_{2} c\right)\right)}{c^{3}(\Omega+c)} \\
& -\frac{3 \gamma_{2}\left(c^{2} A_{6}-6\left(c_{1}+c\right)\right)+12 c^{2} \gamma_{3}}{2(\Omega+c)},
\end{aligned}
$$

which implies

$$
\begin{align*}
u_{\tau} & +\frac{3 c^{2}}{c^{2}+1} u u_{\xi}+\frac{c^{3}}{3\left(c^{2}+1\right)} \mu u_{\xi \xi \xi}+c \gamma_{4}(1-\nu) \mu \varepsilon u_{\tau \xi \xi}+A_{11} \varepsilon u^{2} u_{\xi}  \tag{2.29}\\
& +A_{12} \varepsilon^{2} u^{3} u_{\xi}+A_{10} \varepsilon^{3} u^{4} u_{\xi}+\mu \varepsilon\left[A_{13} u u_{\xi \xi \xi}+A_{14} u_{\xi} u_{\xi \xi}\right]=O\left(\varepsilon^{4}, \varepsilon^{2} \mu, \mu^{2}\right)
\end{align*}
$$

where

$$
\begin{gathered}
A_{11}:=\frac{c^{2} A_{6}-6\left(c_{1}+c\right)}{2 c(\Omega+c)}=\frac{-3 c\left(c^{2}-1\right)\left(c^{2}-2\right)}{2\left(c^{2}+1\right)^{3}}, \\
A_{12}:=\frac{c A_{5}}{2(\Omega+c)}-\frac{9 c^{2} \gamma_{2}}{2(\Omega+c)}-\frac{2\left(c_{1}+c\right)}{c^{2}} \frac{c^{2} A_{6}-6\left(c_{1}+c\right)}{2 c(\Omega+c)}=\frac{\left(c^{2}-1\right)^{2}\left(c^{2}-2\right)\left(8 c^{2}-1\right)}{2\left(c^{2}+1\right)^{5}}, \\
A_{13}:=\frac{c A_{9}}{2(\Omega+c)}-\frac{3 c^{2} \gamma_{4} \nu}{2(\Omega+c)}-\frac{c_{1}+c}{3(\Omega+c)}=\frac{3 c^{3} \gamma_{4}}{\left(c^{2}+1\right)}(1-\nu)+\frac{c^{2}\left(3 c^{4}+8 c^{2}-1\right)}{3\left(c^{2}+1\right)^{3}}, \\
A_{14}:=\frac{c A_{8}}{2(\Omega+c)}-\frac{9 c^{2} \gamma_{4} \nu}{2(\Omega+c)}=\frac{3 c^{3}}{\left(c^{2}+1\right)} \gamma_{4}(1-3 \nu)+\frac{c^{2}\left(6 c^{4}+19 c^{2}+4\right)}{3\left(c^{2}+1\right)^{3}} .
\end{gathered}
$$

Now, we need to use the transformation $x=\varepsilon^{-\frac{1}{2}} \xi+c \varepsilon^{-\frac{3}{2}} \tau, \quad t=\varepsilon^{-\frac{3}{2}} \tau$ to go back to the original variables, and have

$$
\frac{\partial}{\partial \xi}=\varepsilon^{-\frac{1}{2}} \partial_{x}, \quad \frac{\partial}{\partial \tau}=\varepsilon^{-\frac{3}{2}}\left(c \partial_{x}+\partial_{t}\right)
$$

Applying this transformation to the equation (2.29), we obtain

$$
\begin{aligned}
& u_{t}+c u_{x}+\frac{3 c^{2}}{c^{2}+1} \varepsilon u u_{x}+A_{11} \varepsilon^{2} u^{2} u_{x}+A_{12} \varepsilon^{3} u^{3} u_{x}+c \gamma_{4}(1-\nu) \mu u_{t x x} \\
& +\left(\frac{c^{3}}{3\left(c^{2}+1\right)}-c^{2} \gamma_{4}(1-\nu)\right) \mu u_{x x x}+\mu \varepsilon\left(A_{13} u u_{x x x}+A_{14} u_{x} u_{x x}\right)=O\left(\varepsilon^{4}, \mu^{2}\right)
\end{aligned}
$$

In order to get the rotational $b$-family of equations, we need

$$
A_{14}=b A_{13}=\frac{3 b c^{2}}{(b+1)\left(c^{2}+1\right)} c \gamma_{4}(1-\nu)
$$

which turns out

$$
\begin{equation*}
\frac{b c^{3}}{\left(c^{2}+1\right)} \gamma_{4}=\frac{-c^{2}\left(3 c^{4}+(8-b) c^{2}-(2 b+1)\right)}{6\left(c^{2}+1\right)^{3}} \tag{2.30}
\end{equation*}
$$

and then

$$
\frac{3 b c^{2}}{(b+1)\left(c^{2}+1\right)} c \gamma_{4}(1-\nu)=b A_{13}=A_{14}=\frac{-c^{2}\left(3 c^{4}+8 c^{2}-1\right)}{3\left(c^{2}+1\right)^{3}} .
$$

Therefore, it enables us to derive the rotational $b$-family of equations in the form

$$
\begin{aligned}
u_{t}-\beta \mu u_{x x t}+c u_{x}+(b+1) \alpha \varepsilon u u_{x}-\beta_{0} \mu u_{x x x} & +\omega_{1} \varepsilon^{2} u^{2} u_{x}+\omega_{2} \varepsilon^{3} u^{3} u_{x} \\
= & \alpha \beta \varepsilon \mu\left(b u_{x} u_{x x}+u u_{x x x}\right) .
\end{aligned}
$$

Combining (2.30) and (2.22), it is found that the height parameter $z$ in $\gamma_{4}$ may take the value

$$
\begin{equation*}
z_{0}=\left(\frac{b-1}{b}-\frac{2}{3} \frac{1}{\left(c^{2}+1\right)}+\frac{(b-2) c^{2}+2 b+4}{3 b\left(c^{2}+1\right)^{2}}\right)^{1 / 2} \tag{2.31}
\end{equation*}
$$

### 2.2 Details for Derivation of The Rotational $b$-Family of Equations

In this section the details for the asymptotic expansions of $u, W, p, \eta$ are given considering the vanishing orders $O\left(\varepsilon^{1}, \mu^{0}\right), O\left(\varepsilon^{0}, \mu^{1}\right), O\left(\varepsilon^{2}, \mu^{0}\right), O\left(\varepsilon^{1}, \mu^{1}\right), O\left(\varepsilon^{3}, \mu^{0}\right)$, $O\left(\varepsilon^{4}, \mu^{0}\right), O\left(\varepsilon^{2}, \mu^{1}\right)$ and $O\left(\varepsilon^{4}, \mu^{2}\right)$.

Considering the order $O\left(\varepsilon^{1}, \mu^{0}\right)$ terms of (2.3), we obtain from the second equation in (2.33) and the Taylor expansion

$$
\begin{equation*}
f(z)=f(1)+\sum_{n=1}^{\infty} \frac{(z-1)^{n}}{n!} f^{(n)}(1) \tag{2.32}
\end{equation*}
$$

that

$$
\begin{cases}-c u_{10, \xi}+u_{00, \tau}+u_{00} u_{00, \xi}+2 \Omega W_{10}=-p_{10, \xi} & \text { in } 0<z<1,  \tag{2.33}\\ 2 \Omega u_{10}=p_{10, z} & \text { in } 0<z<1, \\ u_{10, \xi}+W_{10, z}=0 & \text { in } 0<z<1, \\ u_{10, z}=0 & \text { in } 0<z<1, \\ p_{10}+p_{00, z} \eta_{00}=\eta_{10} & \text { on } z=1, \\ W_{10}+\eta_{00} W_{00, z}=-c \eta_{10, \xi}+\eta_{00, \tau}+u_{00} \eta_{00, \xi} & \text { on } z=1, \\ W_{10}=0 & \text { on } z=0 .\end{cases}
$$

Similar to the order $O\left(\varepsilon^{0}, \mu^{0}\right)$, we have $u_{10}$ as a function independent of $z$, which is $u_{10}=u_{10}(\tau, \xi)$. Also, from the third equation in (2.33) and the boundary conditions of $W$ on $z=0$ and $z=1$, we obtain

$$
\begin{equation*}
W_{10}=\left.W_{10}\right|_{z=0}+\int_{0}^{z} W_{10, z^{\prime}} d z^{\prime}=-z u_{10, \xi} \tag{2.34}
\end{equation*}
$$

and

$$
\begin{equation*}
\left.W_{10}\right|_{z=1}=-c \eta_{10, \xi}+\eta_{00, \tau}+\left(u_{00} \eta_{00}\right)_{\xi} . \tag{2.35}
\end{equation*}
$$

Considering the third equation in (2.5) and (2.8), (2.35) turns out

$$
\begin{equation*}
u_{10, \xi}=c \eta_{10, \xi}-\eta_{00, \tau}-\left(u_{00} \eta_{00}\right)_{\xi} \tag{2.36}
\end{equation*}
$$

and then

$$
W_{10}=z\left(\eta_{00, \tau}+2 c \eta_{00} \eta_{00, \xi}-c \eta_{10, \xi}\right)
$$

To solve the second equation in (2.33), we consider (2.8) and fourth equation in (2.33). Then, it follows that

$$
p_{10}=\left.p_{10}\right|_{z=1}+\int_{1}^{z} p_{10, z^{\prime}} d z^{\prime}=\eta_{10}-2 \Omega u_{00} \eta_{00}+2 \Omega(z-1) u_{10}
$$

and

$$
\begin{equation*}
p_{10, \xi}=\eta_{10, \xi}-2 \Omega\left(u_{00} \eta_{00}\right)_{\xi}+2 \Omega(z-1) u_{10, \xi} \tag{2.37}
\end{equation*}
$$

Now, from the first equation in (2.33), (2.34), and (2.8) we get

$$
-p_{10, \xi}=-c u_{10, \xi}+c \eta_{00, \tau}+c^{2} \eta_{00} \eta_{00, \xi}-2 \Omega z u_{10, \xi}
$$

and then applying (2.37) and (2.36) to the this equation, we have

$$
\begin{aligned}
0 & =-(c+2 \Omega) u_{10, \xi}+\eta_{10, \xi}+c \eta_{00, \tau}+c^{2} \eta_{00} \eta_{00, \xi}-2 \Omega\left(u_{00} \eta_{00}\right)_{\xi} \\
& =c\left(u_{00} \eta_{00}\right)_{\xi}-\left(c^{2}+2 \Omega c-1\right) \eta_{10, \xi}+2(c+\Omega) \eta_{00, \tau}+c^{2} \eta_{00} \eta_{00, \xi}
\end{aligned}
$$

Since we have (2.8) and (2.12), the equation (2.2) implies

$$
\begin{equation*}
2(\Omega+c) \eta_{00, \tau}+3 c^{2} \eta_{00} \eta_{00, \xi}=0 \tag{2.38}
\end{equation*}
$$

Let define

$$
\begin{equation*}
c_{1}:=-\frac{3 c^{2}}{4(\Omega+c)}=-\frac{3 c^{3}}{2\left(c^{2}+1\right)} . \tag{2.39}
\end{equation*}
$$

Then (2.38) can be written as

$$
\begin{equation*}
\eta_{00, \tau}=c_{1}\left(\eta_{00}^{2}\right)_{\xi} \tag{2.40}
\end{equation*}
$$

If we put (2.36) and (2.40) together, we have

$$
\begin{equation*}
u_{10, \xi}=\left(c \eta_{10}-\left(c+c_{1}\right) \eta_{00}^{2}\right)_{\xi} \tag{2.41}
\end{equation*}
$$

As a result, from the far field conditions $u_{10}, \eta_{00}, \eta_{10} \rightarrow 0$ as $|\xi| \rightarrow \infty$, the following equation is obtained

$$
\begin{equation*}
u_{10}=c \eta_{10}-\left(c+c_{1}\right) \eta_{00}^{2} . \tag{2.42}
\end{equation*}
$$

Also, thanks to (2.40), we have

$$
\begin{equation*}
u_{10, \tau}=c \eta_{10, \tau}-4\left(c+c_{1}\right) c_{1} \eta_{00}^{2} \eta_{00, \xi} \tag{2.43}
\end{equation*}
$$

For the order $O\left(\varepsilon^{0}, \mu^{1}\right)$, the terms of (2.3) are obtained from the second equation in (2.5) and the Taylor expansion (2.32) that

$$
\begin{cases}-c u_{01, \xi}+2 \Omega W_{01}=-p_{01, \xi} & \text { in } 0<z<1,  \tag{2.44}\\ 2 \Omega u_{01}=p_{01, z} & \text { in } 0<z<1, \\ u_{01, \xi}+W_{01, z}=0 & \text { in } 0<z<1, \\ u_{01, z}=0 & \text { in } 0<z<1, \\ p_{01}=\eta_{01} & \text { on } z=1, \\ W_{01}=-c \eta_{01, \xi} & \text { on } z=1, \\ W_{01}=0 & \text { on } z=0 .\end{cases}
$$

The following results may be easily obtained from the (2.44) using similar methods from the previous orders

$$
\begin{equation*}
u_{01}=c \eta_{01}=c \eta_{01}(\tau, \xi), W_{01}=-c z \eta_{01, \xi}, p_{01}=[2 \Omega c(z-1)+1] \eta_{01} \tag{2.45}
\end{equation*}
$$

For the order $O\left(\varepsilon^{2}, \mu^{0}\right)$, the terms of (2.3) are obtained from the Taylor expansion (2.32) that

$$
\begin{cases}-c u_{20, \xi}+u_{10, \tau}+\left(u_{00} u_{10}\right)_{\xi}+2 \Omega W_{20}=-p_{20, \xi} & \text { in } 0<z<1,  \tag{2.46}\\ -2 \Omega u_{20}=-p_{20, z} & \text { in } 0<z<1, \\ u_{20, \xi}+W_{20, z}=0 & \text { in } 0<z<1, \\ u_{20, z}=0 & \text { in } 0<z<1, \\ p_{20}+\eta_{00} p_{10, z}+\eta_{10} p_{00, z}=\eta_{20} & \text { on } z=1, \\ W_{20}+\eta_{00} W_{10, z}+\eta_{10} W_{00, z} & \text { on } z=1, \\ \quad=-c \eta_{20, \xi}+\eta_{10, \tau}+u_{00} \eta_{10, \xi}+u_{10} \eta_{00, \xi} \\ W_{20}=0 & \text { on } z=0 .\end{cases}
$$

Similar to the previous orders, $u_{20}$ is independent of $z$, which is $u_{20}=u_{20}(\tau, \xi)$. Also, from the third equation in (2.46) and the boundary condition of $W_{20}$ at $z=0$, we obtain

$$
\begin{equation*}
W_{20}=-z u_{20, \xi} . \tag{2.47}
\end{equation*}
$$

Considering the boundary condition of $W_{20}$ at $z=1$ and (2.47) with the equations of $W_{00, z}$ and $W_{10, z}$, we have

$$
u_{20, \xi}=c \eta_{20, \xi}-\eta_{10, \tau}-\left(u_{00} \eta_{10}+u_{10} \eta_{00}\right)_{\xi}
$$

which follows from (2.42) and (2.8)

$$
\begin{equation*}
u_{20, \xi}=c \eta_{20, \xi}-\eta_{10, \tau}-2 c\left(\eta_{00} \eta_{10}\right)_{\xi}+\left(c+c_{1}\right)\left(\eta_{00}^{3}\right)_{\xi} . \tag{2.48}
\end{equation*}
$$

Moreover, integrating the second equation in (2.46) and applying the boundary condition of $p_{20}$ at $z=1$, we obtain

$$
\begin{aligned}
p_{20}=\left.p_{20}\right|_{z=1}+\int_{1}^{z} p_{20, z^{\prime}} d z^{\prime} & =\eta_{20}-\left(\eta_{00} p_{10, z}+\eta_{10} p_{00, z}\right)+2 \Omega \int_{1}^{z} u_{20} d z^{\prime} \\
& =\eta_{20}-2 \Omega\left(\eta_{00} u_{10}+\eta_{10} u_{00}\right)+2 \Omega(z-1) u_{20}
\end{aligned}
$$

As taking derivative of the equation with respect to $\xi$, it follows that

$$
\begin{equation*}
p_{20, \xi}=\eta_{20, \xi}-2 \Omega\left(\eta_{00} u_{10}+\eta_{10} u_{00}\right)_{\xi}+2 \Omega(z-1) u_{20, \xi}, \tag{2.49}
\end{equation*}
$$

and combining with the first equation in (2.46), we have

$$
\eta_{20, \xi}-2 \Omega\left(\eta_{00} u_{10}+\eta_{10} u_{00}\right)_{\xi}-(c+2 \Omega) u_{20, \xi}+u_{10, \tau}+\left(u_{00} u_{10}\right)_{\xi}=0 .
$$

From (2.7), (2.42), and (2.43), we get

$$
\begin{equation*}
2(c+\Omega) \eta_{10, \tau}+3 c^{2}\left(\eta_{00} \eta_{10}\right)_{\xi}-\left(2 c+\frac{4}{3} c_{1}\right)\left(c+c_{1}\right)\left(\eta_{00}^{3}\right)_{\xi}=0 . \tag{2.50}
\end{equation*}
$$

It implies that

$$
\begin{equation*}
\eta_{10, \tau}=2 c_{1}\left(\eta_{00} \eta_{10}\right)_{\xi}+\frac{2 c_{1}+3 c}{3(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{3}\right)_{\xi} . \tag{2.51}
\end{equation*}
$$

Then, the equation (2.48) turns out

$$
u_{20, \xi}=c \eta_{20, \xi}-2\left(c+c_{1}\right)\left(\eta_{00} \eta_{10}\right)_{\xi}-\frac{2 c_{1}-3 \Omega}{3(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{3}\right)_{\xi} .
$$

As a result, from the far field conditions $\eta_{00}, \eta_{10}, \eta_{20} \rightarrow 0$ as $|\xi| \rightarrow \infty$, the following equation is obtained

$$
\begin{equation*}
u_{20}=c \eta_{20}-2\left(c+c_{1}\right) \eta_{00} \eta_{10}-\frac{2 c_{1}-3 \Omega}{3(c+\Omega)}\left(c+c_{1}\right) \eta_{00}^{3} \tag{2.52}
\end{equation*}
$$

Also, from (2.40) and (2.51), we have

$$
\begin{equation*}
u_{20, \tau}=c \eta_{20, \tau}-4\left(c+c_{1}\right) c_{1}\left(\eta_{00}^{2} \eta_{10}\right)_{\xi}-\frac{8 c c_{1}+4 c_{1}^{2}+\frac{21}{4} c^{2}}{2(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right)_{\xi} \tag{2.53}
\end{equation*}
$$

For the order $O\left(\varepsilon^{1}, \mu^{1}\right)$, the terms of (2.3) are obtained from the Taylor expansion (2.32) that

$$
\begin{cases}-c u_{11, \xi}+u_{01, \tau}+u_{00} u_{01, \xi}+u_{10} u_{00, \xi}+W_{00} u_{01, z} &  \tag{2.54}\\ \quad+W_{10} u_{00, z}+2 \Omega W_{11}=-p_{11, \xi} & \text { in } 0<z<1, \\ -c W_{00, \xi}-2 \Omega u_{11}=-p_{11, z} & \text { in } 0<z<1, \\ u_{11, \xi}+W_{11, z}=0 & \text { in } 0<z<1, \\ u_{11, z}-W_{00, \xi}=0 & \text { in } 0<z<1, \\ p_{11}=\eta_{11}-\left(\eta_{00} p_{01, z}+\eta_{01} p_{00, z}\right) & \text { on } z=1, \\ W_{11}+W_{00, z} \eta_{01}+W_{01, z} \eta_{00} & \text { on } z=1, \\ =-c \eta_{11, \xi}+\eta_{01, \tau}+u_{00} \eta_{01, \xi}+u_{01} \eta_{00, \xi} & \text { on } z=0 .\end{cases}
$$

From (2.8) and the fourth equation of (2.54), we get

$$
u_{11, z}=-c z \eta_{00, \xi \xi} .
$$

Then we have

$$
\begin{equation*}
u_{11}=-\frac{c}{2} z^{2} \eta_{00, \xi \xi}+\Phi_{11}(\tau, \xi) \tag{2.55}
\end{equation*}
$$

for any smooth function $\Phi_{11}(\tau, \xi)$ which is independent of $z$. Applying integration to the third equation in (2.54) with the boundary condition $\left.W_{11}\right|_{z=0}=0$, the following equation is obtained

$$
\begin{equation*}
W_{11}=\left.W_{11}\right|_{z=0}+\int_{0}^{z} W_{11, z^{\prime}} d z^{\prime}=\frac{c}{6} z^{3} \eta_{00, \xi \xi \xi}-z \partial_{\xi} \Phi_{11}(\tau, \xi) \tag{2.56}
\end{equation*}
$$

Thanks to the equations of $W_{00, z}$ and $W_{01, z}$, and the boundary condition of $W_{11}$ on $\{z=1\}$, we have

$$
\begin{equation*}
-\partial_{\xi} \Phi_{11}(\tau, \xi)=-\frac{c}{6} \eta_{00, \xi \xi \xi}+\left(u_{00} \eta_{01}+\eta_{00} u_{01}\right)_{\xi}-c \eta_{11, \xi}+\eta_{01, \tau} \tag{2.57}
\end{equation*}
$$

Taking account of (2.56) and (2.57), it follows

$$
\begin{equation*}
W_{11}=\frac{c}{6} z\left(z^{2}-1\right) \eta_{00, \xi \xi \xi}+z\left(-c \eta_{11, \xi}+\eta_{01, \tau}+\left(u_{00} \eta_{01}+\eta_{00} u_{01}\right)_{\xi}\right) \tag{2.58}
\end{equation*}
$$

Due to (2.8), (2.45), (2.55), and the boundary condition of $p_{11}$ in (2.54), we deduce from the second equation of (2.54) that

$$
\begin{aligned}
& p_{11}=\left.p_{11}\right|_{z=1}+\int_{1}^{z} p_{11, z^{\prime}} d z^{\prime}=\left.p_{11}\right|_{z=1}+\int_{1}^{z}\left(c W_{00, \xi}+2 \Omega u_{11}\right) d z^{\prime} \\
& =\eta_{11}-2 \Omega\left(u_{00} \eta_{01}+\eta_{00} u_{01}\right)-\left(\frac{c^{2}}{2}\left(z^{2}-1\right)+\frac{\Omega c}{3}\left(z^{3}-1\right)\right) \eta_{00, \xi \xi}+2 \Omega(z-1) \Phi_{11},
\end{aligned}
$$

which implies

$$
\begin{align*}
p_{11, \xi}=\eta_{11, \xi}-2 \Omega\left(u_{00} \eta_{01}+\eta_{00} u_{01}\right)_{\xi} & -\left(\frac{c^{2}}{2}\left(z^{2}-1\right)+\frac{\Omega c}{3}\left(z^{3}-1\right)\right) \eta_{00, \xi \xi \xi}  \tag{2.59}\\
& +2 \Omega(z-1) \partial_{\xi} \Phi_{11}
\end{align*}
$$

Combining (2.59) and the first equation in (2.54), it follows from (2.8), (2.45), and (2.55) that

$$
\begin{align*}
& -c u_{11, \xi}+c \eta_{01, \tau}+c^{2}\left(\eta_{00} \eta_{01}\right)_{\xi}+2 \Omega W_{11}+\eta_{11, \xi}-4 \Omega c\left(\eta_{00} \eta_{01}\right)_{\xi} \\
& -\left(\frac{c^{2}}{2}\left(z^{2}-1\right)+\frac{\Omega c}{3}\left(z^{3}-1\right)\right) \eta_{00, \xi \xi \xi}+2 \Omega(z-1) \partial_{\xi} \Phi_{11}=0 . \tag{2.60}
\end{align*}
$$

Substituting (2.55) and (2.57) into (2.60), we obtain

$$
\begin{equation*}
2(\Omega+c) \eta_{01, \tau}+3 c^{2}\left(\eta_{00} \eta_{01}\right)_{\xi}+\frac{c^{2}}{3} \eta_{00, \xi \xi \xi}=0 \tag{2.61}
\end{equation*}
$$

that is,

$$
\begin{equation*}
\eta_{01, \tau}=2 c_{1}\left(\eta_{00} \eta_{01}\right)_{\xi}+\frac{2 c_{1}}{9} \eta_{00, \xi \xi \xi} \tag{2.62}
\end{equation*}
$$

which, together with $(2.57),(2.58)$, and (2.55), leads to

$$
-\partial_{\xi} \Phi_{11}(\tau, \xi)=\left(\frac{2 c_{1}}{9}-\frac{c}{6}\right) \eta_{00, \xi \xi \xi}+2\left(c+c_{1}\right)\left(\eta_{00} \eta_{01}\right)_{\xi}-c \eta_{11, \xi}
$$

and then

$$
W_{11}=\left(\frac{2 c_{1}}{9}+\frac{c}{6}\left(z^{2}-1\right)\right) z \eta_{00, \xi \xi \xi}+2\left(c+c_{1}\right) z\left(\eta_{00} \eta_{01}\right)_{\xi}-c z \eta_{11, \xi}
$$

and

$$
\begin{equation*}
u_{11}=\left(\frac{c}{6}-\frac{2 c_{1}}{9}-\frac{c}{2} z^{2}\right) \eta_{00, \xi \xi}+c \eta_{11}-2\left(c+c_{1}\right) \eta_{00} \eta_{01} \tag{2.63}
\end{equation*}
$$

where use has been made by the far field conditions $u_{11}, \eta_{00, \xi \xi}, \eta_{00}, \eta_{01}, \eta_{11} \rightarrow 0$ as $|\xi| \rightarrow \infty$.

Thanks to (2.40) and (2.62), we obtain

$$
\begin{align*}
u_{11, \tau}= & c \eta_{11, \tau}+\left(\frac{c c_{1}}{6}-\frac{2 c_{1}^{2}}{9}-\frac{c c_{1}}{2} z^{2}\right)\left(\eta_{00}^{2}\right)_{\xi \xi \xi}  \tag{2.64}\\
& -2\left(c+c_{1}\right)\left(2 c_{1}\left(\eta_{00}^{2} \eta_{01}\right)_{\xi}+\frac{2 c_{1}}{9} \eta_{00} \eta_{00, \xi \xi \xi}\right) .
\end{align*}
$$

For the order $O\left(\varepsilon^{3}, \mu^{0}\right)$, the terms of (2.3) are obtained from the Taylor expansion (2.32) that

$$
\begin{cases}-c u_{30, \xi}+u_{20, \tau}+\left(u_{00} u_{20}+\frac{1}{2} u_{10}^{2}\right)_{\xi}+2 \Omega W_{30}=-p_{30, \xi} & \text { in } 0<z<1,  \tag{2.65}\\ -2 \Omega u_{30}=-p_{30, z} & \text { in } 0<z<1, \\ u_{30, \xi}+W_{30, z}=0 & \text { in } 0<z<1, \\ u_{30, z}=0 & \text { in } 0<z<1, \\ p_{30}+\eta_{00} p_{20, z}+\eta_{10} p_{10, z}+\eta_{20} p_{00, z}=\eta_{30} & \text { on } z=1, \\ W_{30}+\eta_{00} W_{20, z}+\eta_{10} W_{10, z}+\eta_{20} W_{00, z} & \text { on } z=1, \\ \quad=-c \eta_{30, \xi}+\eta_{20, \tau}+u_{00} \eta_{20, \xi}+u_{10} \eta_{10, \xi}+u_{20} \eta_{00, \xi} & \text { on } z=0 .\end{cases}
$$

Similarly, it follows that $u_{30}$ is independent of $z$, which is $u_{30}=u_{30}(\tau, \xi)$ from the fourth equation in (2.65). Also, from the third equation in (2.65) and the boundary condition of $W_{30}$ at $z=0$, we obtain

$$
\begin{equation*}
W_{30}=-z u_{30, \xi} . \tag{2.66}
\end{equation*}
$$

Considering the boundary condition of $W_{20}$ at $z=1$ and (2.47), we have

$$
\begin{equation*}
u_{30, \xi}=c \eta_{30, \xi}-\eta_{20, \tau}-\left(u_{00} \eta_{20}+u_{10} \eta_{10}+u_{20} \eta_{00}\right)_{\xi} . \tag{2.67}
\end{equation*}
$$

In addition, integrating the second equation in (2.65) and applying the boundary condition of $p_{30}$ at $z=1$, we have

$$
\begin{aligned}
& p_{30}=\left.p_{30}\right|_{z=1}+\int_{1}^{z} p_{30, z^{\prime}} d z^{\prime} \\
& =\eta_{30}-\left(\eta_{00} p_{20, z}+\eta_{10} p_{10, z}+\eta_{20} p_{00, z}\right)+2 \Omega \int_{1}^{z} u_{30} d z^{\prime} \\
& =\eta_{30}-2 \Omega\left(u_{00} \eta_{20}+u_{10} \eta_{10}+u_{20} \eta_{00}\right)+2 \Omega(z-1) u_{30} .
\end{aligned}
$$

As taking derivative of the equation with respect to $\xi$, it follows that

$$
\begin{equation*}
p_{30, \xi}=\eta_{30, \xi}-2 \Omega\left(u_{00} \eta_{20}+u_{10} \eta_{10}+u_{20} \eta_{00}\right)_{\xi}+2 \Omega(z-1) u_{30, \xi}, \tag{2.68}
\end{equation*}
$$

and from the first equation in (2.65), we get

$$
\begin{equation*}
-p_{30, \xi}=-c u_{30, \xi}+u_{20, \tau}+\left(u_{00} u_{20}+\frac{1}{2} u_{10}^{2}\right)_{\xi}-2 \Omega z u_{30, \xi} . \tag{2.69}
\end{equation*}
$$

Therefore, from (2.68) and (2.69), it follows that

$$
\begin{equation*}
0=\eta_{30, \xi}-2 \Omega\left(u_{00} \eta_{20}+u_{10} \eta_{10}+u_{20} \eta_{00}\right)_{\xi}-(c+2 \Omega) u_{30, \xi}+u_{20, \tau}+\left(u_{00} u_{20}+\frac{1}{2} u_{10}^{2}\right)_{\xi} \tag{2.70}
\end{equation*}
$$

Applying (2.67) and (2.53) to (2.70), we have

$$
\begin{align*}
2(c+\Omega) \eta_{20, \tau}+3 c^{2}\left(\eta_{00} \eta_{20}\right)_{\xi} & +\frac{3 c^{2}}{2}\left(\eta_{10}^{2}\right)_{\xi}-2\left(2 c_{1}+3 c\right)\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{10}\right)_{\xi} \\
& -\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}-15\right)}{12(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right)_{\xi}=0, \tag{2.71}
\end{align*}
$$

which is equivalent to

$$
\begin{align*}
\eta_{20, \tau}=2 c_{1}\left(\eta_{00} \eta_{20}\right)_{\xi}+c_{1}\left(\eta_{10}^{2}\right)_{\xi} & +\frac{2 c_{1}+3 c}{\Omega+c}\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{10}\right)_{\xi} \\
& +\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}-15\right)}{24(c+\Omega)^{2}}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right)_{\xi} \tag{2.72}
\end{align*}
$$

Then substituting (2.72) into (2.67), we obtain Thanks to (2.67) again, we have

$$
\begin{aligned}
u_{30, \xi}= & c \eta_{30, \xi}-2\left(c+c_{1}\right)\left(\eta_{00} \eta_{20}\right)_{\xi}-\left(c+c_{1}\right)\left(\eta_{10}^{2}\right)_{\xi}-\frac{2 c_{1}-3 \Omega}{\Omega+c}\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{10}\right)_{\xi} \\
& -\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{24(c+\Omega)^{2}}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right)_{\xi},
\end{aligned}
$$

which follows

$$
\begin{align*}
u_{30}= & c \eta_{30}-2\left(c+c_{1}\right)\left(\eta_{00} \eta_{20}\right)-\left(c+c_{1}\right)\left(\eta_{10}^{2}\right)-\frac{2 c_{1}-3 \Omega}{\Omega+c}\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{10}\right) \\
& -\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{24(c+\Omega)^{2}}\left(c+c_{1}\right)\left(\eta_{00}^{4}\right) \tag{2.73}
\end{align*}
$$

Taking account of (2.40), (2.51), and (2.72), it turns into

$$
\begin{align*}
u_{30, \tau}= & c \eta_{30, \tau}-\frac{2\left(3 c^{2}+5 c c_{1}+4 c_{1}^{2}-3 \Omega c_{1}\right)}{\Omega+c}\left(c+c_{1}\right)\left(\eta_{00}^{3} \eta_{10}\right)_{\xi}  \tag{2.74}\\
& -4 c_{1}\left(c+c_{1}\right)\left(\eta_{00} \eta_{10}^{2}\right)_{\xi}-4 c_{1}\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{20}\right)_{\xi}-B_{1} \eta_{00}^{4} \eta_{00, \xi},
\end{align*}
$$

where

$$
\begin{aligned}
B_{1} \stackrel{\text { def }}{=} & \frac{\left(c+c_{1}\right)^{2}\left(82 c c_{1}+36 c_{1}^{2}+45 c^{2}-18 \Omega c_{1}-27 \Omega c-15\right)}{3(\Omega+c)^{2}} \\
& +\frac{c_{1}\left(c+c_{1}\right)\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{3(\Omega+c)^{2}}
\end{aligned}
$$

For order $O\left(\varepsilon^{4}, \mu^{0}\right)$, the terms of (2.3) are obtained from the Taylor expansion (2.32) that

$$
\begin{cases}-c u_{40, \xi}+u_{30, \tau}+\left(u_{00} u_{30}+u_{10} u_{20}\right)_{\xi}+2 \Omega W_{40}=-p_{40, \xi} & \text { in } 0<z<1,  \tag{2.75}\\ -2 \Omega u_{40}=-p_{40, z} & \text { in } 0<z<1, \\ u_{40, \xi}+W_{40, z}=0 & \text { in } 0<z<1, \\ u_{40, z}=0 & \text { in } 0<z<1, \\ p_{40}+\eta_{00} p_{30, z}+\eta_{10} p_{20, z}+\eta_{20} p_{10, z}+\eta_{30} p_{00, z}=\eta_{40} & \text { on } z=1, \\ W_{40}+\eta_{00} W_{30, z}+\eta_{10} W_{20, z}+\eta_{20} W_{10, z}+\eta_{30} W_{00, z} & \text { on } z=1, \\ =-c \eta_{40, \xi}+\eta_{30, \tau}+u_{00} \eta_{30, \xi}+u_{10} \eta_{20, \xi}+u_{20} \eta_{10, \xi}+u_{30} \eta_{00, \xi} & \text { on } z=0 .\end{cases}
$$

Similarly to the previous orders, it follows that $u_{40}$ is independent of $z$, which is $u_{40}=u_{40}(\tau, \xi)$, from the fourth equation in (2.75). Also, from the third equation in (2.75) and the boundary condition of $W_{40}$ at $z=0$, we have

$$
\begin{equation*}
W_{40}=-z u_{40, \xi} . \tag{2.76}
\end{equation*}
$$

Taking account of the boundary condition of $W_{40}$ at $z=1,(2.76),(2.66),(2.47)$, (2.34) and (2.6), we obtain

$$
\begin{equation*}
u_{40, \xi}=c \eta_{40, \xi}-\eta_{30, \tau}-\left(u_{00} \eta_{30}+u_{10} \eta_{20}+u_{20} \eta_{10}+u_{30} \eta_{00}\right)_{\xi}, \tag{2.77}
\end{equation*}
$$

Moreover, integrating the second equation in (2.75) and considering the boundary condition of $p_{30}$ at $z=1$, we have

$$
\begin{aligned}
& p_{40}=\left.p_{40}\right|_{z=1}+\int_{1}^{z} p_{40, z^{\prime}} d z^{\prime} \\
& =\eta_{40}-\left(\eta_{00} p_{30, z}+\eta_{10} p_{20, z}+\eta_{20} p_{10, z}+\eta_{30} p_{00, z}\right)+2 \Omega \int_{1}^{z} u_{40} d z^{\prime} \\
& =\eta_{40}-2 \Omega\left(u_{00} \eta_{30}+u_{10} \eta_{20}+u_{20} \eta_{10}+u_{30} \eta_{00}\right)+2 \Omega(z-1) u_{40} .
\end{aligned}
$$

Taking derivative of the equation with respect to $\xi$, we get

$$
\begin{equation*}
p_{40, \xi}=-\eta_{40, \xi}-2 \Omega\left(u_{00} \eta_{30}+u_{10} \eta_{20}+u_{20} \eta_{10}+u_{30} \eta_{00}\right)_{\xi}+2 \Omega(z-1) u_{40, \xi} \tag{2.78}
\end{equation*}
$$

and then from the first equation in (2.75), we obtain

$$
-p_{40, \xi}=-c u_{40, \xi}+u_{30, \tau}+\left(u_{00} u_{30}+u_{10} u_{20}\right)_{\xi}+2 \Omega W_{40}
$$

Therefore, from (2.76) and (2.78), they turn into

$$
\begin{align*}
0= & -(c+2 \Omega) u_{40, \xi}+u_{30, \tau}+\left(u_{00} u_{30}+u_{10} u_{20}\right)_{\xi}  \tag{2.79}\\
& +\eta_{40, \xi}-2 \Omega\left(u_{00} \eta_{30}+u_{10} \eta_{20}+u_{20} \eta_{10}+u_{30} \eta_{00}\right)_{\xi} .
\end{align*}
$$

Applying (2.77) and (2.74) to (2.79), we have

$$
\begin{align*}
& 2(c+\Omega) \eta_{30, \tau}+3 c^{2}\left(\eta_{00} \eta_{30}+\eta_{10} \eta_{20}\right)_{\xi}-2\left(3 c+2 c_{1}\right)\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{20}+\eta_{00} \eta_{10}^{2}\right)_{\xi} \\
& \quad-\frac{\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}-15\right)}{3(c+\Omega)}\left(c+c_{1}\right)\left(\eta_{00}^{3} \eta_{10}\right)_{\xi}-B_{2}\left(\eta_{00}^{5}\right)_{\xi}=0, \tag{2.80}
\end{align*}
$$

where

$$
\begin{aligned}
B_{2} & \stackrel{\text { def }}{=} \frac{1}{5} B_{1}-\frac{\left(c+c_{1}\right)^{2}\left(2 c_{1}-3 \Omega\right)}{3(\Omega+c)}+\frac{2 c\left(c+c_{1}\right)\left(64 c c_{1}+24 c_{1}^{2}+45 c^{2}+24 \Omega^{2}-3\right)}{12(\Omega+c)^{2}} \\
& =\frac{c^{2}\left(2-c^{2}\right)\left(3 c^{10}+228 c^{8}-540 c^{6}-180 c^{4}-13 c^{2}+42\right)}{60\left(c^{2}+1\right)^{6}} .
\end{aligned}
$$

Similarly, For the order $O\left(\varepsilon^{2}, \mu^{1}\right)$ the terms in (2.3) are obtained as the following,

$$
\begin{cases}-c u_{21, \xi}+u_{11, \tau}+\left(u_{00} u_{11}+u_{10} u_{01}\right)_{\xi}+W_{00} u_{11, z}+2 \Omega W_{21}=-p_{21, \xi} & \text { in } 0<z<1,  \tag{2.81}\\ -c W_{10, \xi}+W_{00, \tau}+u_{00} W_{00, \xi}+W_{00} W_{00, z}-2 \Omega u_{21}=-p_{21, z} & \text { in } 0<z<1, \\ u_{21, \xi}+W_{21, z}=0 & \text { in } 0<z<1, \\ u_{21, z}-W_{10, \xi}=0 & \text { in } 0<z<1, \\ p_{21}+\eta_{10} p_{01, z}+\eta_{01} p_{10, z}+\eta_{00} p_{11, z}+\eta_{11} p_{00, z}=\eta_{21} & \text { on } z=1, \\ W_{21}+\eta_{10} W_{01, z}+\eta_{01} W_{10, z}+\eta_{00} W_{11, z}+\eta_{11} W_{00, z} & \text { on } z=1, \\ =-c \eta_{21, \xi}+\eta_{11, \tau}+u_{00} \eta_{11, \xi}+u_{11} \eta_{00, \xi}+u_{10} \eta_{01, \xi}+u_{01} \eta_{10, \xi} & \text { on } z=0 . \\ W_{21}=0 & \end{cases}
$$

Applying (2.34) and(2.41) to the fourth equation in (2.81), it follows

$$
u_{21, z}=W_{10, \xi}=z\left(2\left(c+c_{1}\right)\left(\eta_{00, \xi}^{2}+\eta_{00} \eta_{00, \xi \xi}\right)-c \eta_{10, \xi \xi}\right)
$$

which turns into

$$
u_{21}=\frac{z^{2}}{2}\left(2\left(c+c_{1}\right)\left(\eta_{00, \xi}^{2}+\eta_{00} \eta_{00, \xi \xi}\right)-c \eta_{10, \xi \xi}\right)+\Phi_{21}(\tau, \xi)=\frac{z^{2}}{2} H_{1}+\Phi_{21}(\tau, \xi)
$$

for any smooth function $\Phi_{21}(\tau, \xi)$ independent of $z$, where we describe

$$
H_{1} \stackrel{\text { def }}{=} 2\left(c+c_{1}\right)\left(\eta_{00, \xi}^{2}+\eta_{00} \eta_{00, \xi \xi}\right)-c \eta_{10, \xi \xi} .
$$

Then, we get

$$
u_{21, \xi}=\frac{z^{2}}{2} H_{1, \xi}+\partial_{\xi} \Phi_{21}(\tau, \xi)
$$

Also, from the third equation in (2.81) and the boundary condition of $W_{21}$ on $\{z=0\}$, we obtain

$$
W_{21}=\left.W_{21}\right|_{z=0}+\int_{0}^{z} W_{21, z^{\prime}} d z^{\prime}=-\int_{0}^{z} u_{21, \xi} d z^{\prime}=-\frac{z^{3}}{6} H_{1, \xi}-z \partial_{\xi} \Phi_{21}(\tau, \xi) .
$$

Considering the third equation of the orders $O\left(\varepsilon^{0}, \mu^{1}\right), O\left(\varepsilon^{1}, \mu^{0}\right), O\left(\varepsilon^{1}, \mu^{1}\right), O\left(\varepsilon^{0}, \mu^{0}\right)$ with the boundary condition of $W_{21}$ on $\{z=1\}$, it follows

$$
\begin{aligned}
-\frac{1}{6} H_{1, \xi}-\partial_{\xi} \Phi_{21}(\tau, \xi) & =-c \eta_{21, \xi}+\eta_{11, \tau}+\left.\left(u_{00} \eta_{11}+u_{11} \eta_{00}+u_{10} \eta_{01}+u_{01} \eta_{10}\right)_{\xi}\right|_{z=1} \\
& =-c \eta_{21, \xi}+\eta_{11, \tau}+\left.H_{2, \xi}\right|_{z=1}
\end{aligned}
$$

where $H_{2}$ is defined as

$$
H_{2} \stackrel{\text { def }}{=} u_{00} \eta_{11}+u_{11} \eta_{00}+u_{10} \eta_{01}+u_{01} \eta_{10} .
$$

Then we have

$$
\begin{equation*}
\partial_{\xi} \Phi_{21}(\tau, \xi)=c \eta_{21, \xi}-\eta_{11, \tau}-\frac{1}{6} H_{1, \xi}-\left.H_{2, \xi}\right|_{z=1} . \tag{2.82}
\end{equation*}
$$

This gives us

$$
\begin{equation*}
u_{21, \xi}=c \eta_{21, \xi}-\eta_{11, \tau}+\left(\frac{z^{2}}{2}-\frac{1}{6}\right) H_{1, \xi}-\left.H_{2, \xi}\right|_{z=1} \tag{2.83}
\end{equation*}
$$

and

$$
\begin{equation*}
W_{21}=\frac{z\left(1-z^{2}\right)}{6} H_{1, \xi}-c z \eta_{21, \xi}+z \eta_{11, \tau}+z\left(\left.H_{2, \xi}\right|_{z=1}\right) . \tag{2.84}
\end{equation*}
$$

Substituting the expressions of $W_{00, \tau}, u_{00}, W_{00, \xi}, W_{00}, W_{00, z}$, and $W_{10, \xi}$ into On the other hand, using the results from the previous orders, the second equation in (2.81) turns into

$$
\begin{equation*}
p_{21, z}=2 \Omega u_{21}-c^{2} z \eta_{10, \xi \xi}+c\left(c+4 c_{1}\right) z \eta_{00, \xi}^{2}+c\left(3 c+4 c_{1}\right) z \eta_{00} \eta_{00, \xi \xi} . \tag{2.85}
\end{equation*}
$$

From the boundary condition of $p_{21}$ on $z=1$, we get

$$
\left.p_{21}\right|_{z=1}=\eta_{21}+c^{2} \eta_{00} \eta_{00, \xi \xi}-\left.2 \Omega H_{2}\right|_{z=1}
$$

Also, integrating (2.85), it follows

$$
\begin{align*}
& p_{21}=\left.p_{21}\right|_{z=1}+\int_{1}^{z} p_{21, z^{\prime}} d z^{\prime} \\
& =\eta_{21}-\left.2 \Omega H_{2}\right|_{z=1}+2 \Omega \int_{1}^{z} u_{21} d z^{\prime}-\frac{c^{2}}{2}\left(z^{2}-1\right) \eta_{10, \xi \xi}  \tag{2.86}\\
& \quad+\frac{c\left(c+4 c_{1}\right)}{2}\left(z^{2}-1\right) \eta_{00, \xi}^{2}+\left(c^{2}+\frac{c\left(3 c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\right) \eta_{00} \eta_{00, \xi \xi}
\end{align*}
$$

Then we have

$$
\begin{align*}
& p_{21, \xi}=\eta_{21, \xi}-\left.2 \Omega H_{2, \xi}\right|_{z=1}+2 \Omega \int_{1}^{z} u_{21, \xi} d z^{\prime}-\frac{c^{2}}{2}\left(z^{2}-1\right) \eta_{10, \xi \xi \xi} \\
& \quad+\frac{c\left(c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\left(\eta_{00, \xi}^{2}\right)_{\xi}+\left(c^{2}+\frac{c\left(3 c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi} \\
& =-\left.2 \Omega z H_{2, \xi}\right|_{z=1}+2 \Omega(z-1)\left(c \eta_{21, \xi}-\eta_{11, \tau}\right)+\frac{z\left(z^{2}-1\right)}{6} H_{1, \xi}-\frac{c^{2}}{2}\left(z^{2}-1\right) \eta_{10, \xi \xi \xi} \\
& \quad+\eta_{21, \xi}+\frac{c\left(c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\left(\eta_{00, \xi}^{2}\right)_{\xi}+\left(c^{2}+\frac{c\left(3 c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi} . \tag{2.87}
\end{align*}
$$

From the first equation in (2.81), (2.84), and (2.8), we obtain

$$
\begin{align*}
-p_{21, \xi}= & -c u_{21, \xi}+u_{11, \tau}+\left(u_{00} u_{11}+u_{10} u_{01}\right)_{\xi}+c^{2} z^{2} \eta_{00, \xi} \eta_{00, \xi \xi} \\
& +\frac{\Omega}{3} z\left(1-z^{2}\right) H_{1, \xi}-2 \Omega c z \eta_{21, \xi}+2 \Omega z \eta_{11, \tau}+\left.2 \Omega z H_{2, \xi}\right|_{z=1} . \tag{2.88}
\end{align*}
$$

Taking account of (2.88) and (2.86), we have

$$
\begin{align*}
0 & =-c u_{21, \xi}+u_{11, \tau}+\left(u_{00} u_{11}+u_{10} u_{01}\right)_{\xi}+\left(\frac{c^{2}}{2} z^{2}+\frac{c\left(c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\right)\left(\eta_{00, \xi}^{2}\right)_{\xi} \\
& +\frac{\Omega}{3} z\left(1-z^{2}\right) H_{1, \xi}+(1-2 \Omega c) \eta_{21, \xi}+2 \Omega \eta_{11, \tau}+\frac{z\left(z^{2}-1\right)}{6} H_{1, \xi}-\frac{c^{2}}{2}\left(z^{2}-1\right) \eta_{10, \xi \xi \xi} \\
& +\left(c^{2}+\frac{c\left(3 c+4 c_{1}\right)}{2}\left(z^{2}-1\right)\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi} . \tag{2.89}
\end{align*}
$$

Notice that

$$
\begin{aligned}
& \left(u_{01} u_{10}+u_{00} u_{11}\right)_{\xi} \\
& =c^{2}\left(\eta_{01} \eta_{10}+\eta_{00} \eta_{11}\right)_{\xi}+\left(\frac{c^{2}}{6}-\frac{2 c c_{1}}{9}-\frac{c^{2} z^{2}}{2}\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi}-3 c\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{01}\right)_{\xi}
\end{aligned}
$$

and

$$
\left.c H_{2, \xi}\right|_{z=1}=2 c^{2}\left(\eta_{01} \eta_{10}+\eta_{00} \eta_{11}\right)_{\xi}-\left(\frac{c^{2}}{3}+\frac{2 c c_{1}}{9}\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi}-3 c\left(c+c_{1}\right)\left(\eta_{00}^{2} \eta_{01}\right)_{\xi}
$$

We substitute (2.83) and (2.64) into (2.89) to get

$$
\begin{align*}
& 2(\Omega+c) \eta_{11, \tau}+3 c^{2}\left(\eta_{00} \eta_{11}+\eta_{10} \eta_{01}\right)_{\xi}-2\left(c+c_{1}\right)\left(3 c+2 c_{1}\right)\left(\eta_{00}^{2} \eta_{01}\right)_{\xi}+\frac{c^{2}}{3} \eta_{10, \xi \xi \xi} \\
& -\left(\frac{c^{2}}{6}+\frac{10 c c_{1}}{9}+\frac{2 c_{1}^{2}}{9}\right)\left(\eta_{00, \xi}^{2}\right)_{\xi}-\left(\frac{c^{2}}{3}+\frac{20 c c_{1}}{9}+\frac{8 c_{1}^{2}}{9}\right)\left(\eta_{00} \eta_{00, \xi \xi}\right)_{\xi}=0 \tag{2.90}
\end{align*}
$$

## CHAPTER 3

## LOCAL WELL POSEDNESS OF ROTATIONAL $b$-FAMILY OF EQUATIONS

In this section, the local well-posedness for the R - $b$-family equations is investigated.

Consider the R-b-family equation (1.5) in terms of the evolution of $m$, namely, the equation (1.9). Applying the transformation $u_{\varepsilon, \mu}(t, x)=\alpha \varepsilon u(\sqrt{\beta \mu} t, \sqrt{\beta \mu} x)$ to (1.9), we know that $u_{\varepsilon, \mu}(t, x)$ solves

$$
\begin{equation*}
u_{t}-u_{x x t}+c u_{x}+(b+1) u u_{x}-\frac{\beta_{0}}{\beta} u_{x x x}+\frac{\omega_{1}}{\alpha^{2}} u^{2} u_{x}+\frac{\omega_{2}}{\alpha^{3}} u^{3} u_{x}=b u_{x} u_{x x}+u u_{x x x} \tag{3.1}
\end{equation*}
$$

and its corresponding one conserved quantity denoted by $I(u)$ is as follows

$$
I(u)=\int_{\mathbb{R}} u d x
$$

As a special case $b=2$, the equation has three conserved quantities as follows:

$$
I(u)=\int_{\mathbb{R}} u d x, \quad E(u)=\frac{1}{2} \int_{\mathbb{R}} u^{2}+u_{x}^{2} d x,
$$

and

$$
F(u)=\frac{1}{2} \int_{\mathbb{R}} c u^{2}+u^{3}+\frac{\beta_{0}}{\beta} u_{x}^{2}+\frac{\omega_{1}}{6 \alpha^{2}} u^{4}+\frac{\omega_{2}}{10 \alpha^{3}} u^{5}+u u_{x}^{2} d x .
$$

Also, we have two more forms of equations,

$$
\left\{\begin{array}{l}
m_{t}+u m_{x}+b u_{x} m+c u_{x}-\frac{\beta_{0}}{\beta} u_{x x x}+\frac{\omega_{1}}{\alpha^{2}} u^{2} u_{x}+\frac{\omega_{2}}{\alpha^{3}} u^{3} u_{x}=0  \tag{3.2}\\
m=u-u_{x x}
\end{array}\right.
$$

and
$u_{t}+u u_{x}+\frac{\beta_{0}}{\beta} u_{x}+p * \partial_{x}\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{(b-3)}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}=0$,
where $p=\frac{1}{2} e^{-|x|}$.

### 3.1 Preliminaries

In the Lebesgue space $L^{p}(\mathbb{R})$ the norm is defined as $\|f\|_{p}=\left(\int_{\mathbb{R}}|f(x)|^{p} d x\right)^{1 / p}$. For the space $L^{\infty}(\mathbb{R})$ consisting of all essentially bounded functions, the norm is given by $\|f\|_{\infty}=\inf _{\mu(E)=0} \sup _{x \in \mathbb{R} \backslash E}|f(x)|$, where $f$ is Lebesgue measurable functions. For a function $f$ in the classical Sobolev spaces $H^{s}(\mathbb{R})(s \geq 0)$ the norm is denoted by $\|f\|_{H^{s}}$. We denote $p(x)=\frac{1}{2} e^{-|x|}$ the fundamental solution of $1-\partial_{x}^{2}$ on $\mathbb{R}$, and define the two convolution operators $p_{+}, p_{-}$as

$$
\begin{aligned}
& p_{+} * f(x)=\frac{e^{-x}}{2} \int_{-\infty}^{x} e^{y} f(y) d y \\
& p_{-} * f(x)=\frac{e^{x}}{2} \int_{x}^{\infty} e^{-y} f(y) d y
\end{aligned}
$$

Then we have the relations $p=p_{+}+p_{-}, \quad p_{x}=p_{-}-p_{+}$.
Definition 3.1.1 ([15]). Let $s \in \mathbb{R}$. A tempered distribution $u$ belongs to $H^{s}\left(\mathbb{R}^{N}\right)$ if $\widehat{u} \in L_{l o c}^{2}\left(\mathbb{R}^{N}\right)$ and

$$
\|u\|_{H^{s}}:=\left(\int_{\mathbb{R}^{N}}\left(1+|\xi|^{2}\right)^{s}|\widehat{u}(\xi)|^{2} d \xi\right)^{\frac{1}{2}}<\infty
$$

It is classical that $H^{s}$ endowed with the norm $\|\cdot\|_{H^{s}}$ is a Banach space
Lemma 3.1.1 ([31]). [Commutator Estimate] Let $\Lambda$ be a operator defined as $\Lambda=$ $\left(1-\partial_{x}^{2}\right)^{1 / 2}$. If $f$ and $g$ are smooth enough, then

$$
\begin{equation*}
\left\|\left[\Lambda^{s}, f\right] g\right\|_{L^{2}} \leq C\left(\|f\|_{H^{s}}\|g\|_{L^{\infty}}+\left\|\partial_{x} f\right\|_{L^{\infty}}\|g\|_{H^{s-1}}\right) \tag{3.4}
\end{equation*}
$$

for all $s>\frac{3}{2}$ and $C>0$.
Lemma 3.1.2. [22] Let $f(t)$ and $g(t)$ be two positive function on $[0, T]$. If the differential inequality

$$
\frac{d}{d t}\left(h^{2}\right)(t) \leq f(t) h^{2}(t)+g(t) h(t)
$$

for a nonnegative absolutely continuous function $h(\cdot)$ on $[0, T]$ holds for almost everywhere on $[0, T]$, then

$$
\begin{equation*}
h(t) \leq e^{\frac{1}{2} \int_{0}^{t} f(\tau) d \tau}\left[h(0)+\frac{1}{2} \int_{0}^{t} g(\tau) d \tau\right] . \tag{3.5}
\end{equation*}
$$

### 3.2 Local Well Posedness

In this section, we introduce the local well-posedness result of the following Cauchy problem, which is similar to the approach in [14] with some adjustments

$$
\left\{\begin{array}{l}
u_{t}-u_{x x t}+c u_{x}+(b+1) u u_{x}-\frac{\beta_{0}}{\beta} u_{x x x}+\frac{\omega_{1}}{\alpha^{2}} u^{2} u_{x}+\frac{\omega_{2}}{\alpha^{3}} 3^{3} u_{x}=b u_{x} u_{x x}+u u_{x x x}  \tag{3.6}\\
\left.u\right|_{t=0}=u_{0}
\end{array}\right.
$$

Theorem 3.2.1 (Local well-posedness). Let $s>\frac{3}{2}, u_{0} \in H^{s}(\mathbb{R})$. Then, there exist the existence time $T>0$ and a unique solution $u \in C\left([0, T] ; H^{s}(\mathbb{R})\right) \cap C^{1}\left([0, T] ; H^{s-1}(\mathbb{R})\right)$ to the Cauchy problem.

Proof. Step 1: First of all, the Cauchy problem (3.6) can be written as

$$
\left\{\begin{aligned}
& \partial_{t} u+u \partial_{x} u+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x} u \\
&=-p * \partial_{x}\left\{\frac{3-b}{2} u_{x}^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\} \\
&\left.u\right|_{t=0}=u_{0}
\end{aligned}\right.
$$

Then, we establish a sequence of approximation solutions $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ as follow. Let $u^{(0)}=0, \forall n \in \mathbb{N}$

$$
\left\{\begin{align*}
& \partial_{t} u^{(n+1)}+u^{(n)} \partial_{x} u^{(n+1)}+\left[\frac{\beta_{0}}{\beta}+\right.\left.\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x} u^{(n+1)}  \tag{3.7}\\
&=-p * \partial_{x}\left\{\frac{3-b}{2}\left[u^{(n)}\right]_{x}^{2}+\frac{b}{2}\left[u^{(n)}\right]^{2}+\frac{\omega_{1}}{3 \alpha^{2}}\left[u^{(n)}\right]^{3}+\frac{\omega_{2}}{4 \alpha^{3}}\left[u^{(n)}\right]^{4}\right\}, \\
&\left.u^{(n+1)}\right|_{t=0}=S_{n+1} u_{0}
\end{align*}\right.
$$

where

$$
\widehat{S_{n+1} u_{0}}(\xi)=1_{|\xi|<2 \times 2^{n+2}}(\xi) \hat{u_{0}}(\xi) .
$$

By the theory of the linear evolution, there is a unique smooth solution $u^{(n+1)}$ of (3.7) and $\forall n \in \mathbb{N}, u^{(n+1)} \in C^{1}\left((R) ; H^{\infty}(\mathbb{R})\right)$. Next we will show that the sequence $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ converges. To prove the convergence, we will prove that $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ is uniformly bounded and a Cauchy sequence, since the Sobolev space is compact.

Step 2: In this step, we show that the sequence $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ is uniformly bounded in some Sobolev space. Consider the operator

$$
\Lambda^{s}=\left(1-\partial_{x}^{2}\right)^{s / 2}
$$

Applying the operator $\Lambda^{s}$ to the (3.7), we get

$$
\begin{aligned}
\partial_{t} \Lambda^{s} u^{(n+1)} & =-\Lambda^{s}\left(u^{(n)} \partial_{x} u^{(n+1)}\right)-\frac{\beta_{0}}{\beta} \Lambda^{s} \partial_{x} u^{(n+1)}-\left(c-\frac{\beta_{0}}{\beta}\right) \Lambda^{s}\left(p * \partial_{x} u^{(n+1)}\right) \\
- & \Lambda^{s}\left(p * \partial_{x}\left\{\frac{3-b}{2}\left[u^{(n)}\right]_{x}^{2}+\frac{b}{2}\left[u^{(n)}\right]^{2}+\frac{\omega_{1}}{3 \alpha^{2}}\left[u^{(n)}\right]^{3}+\frac{\omega_{2}}{4 \alpha^{3}}\left[u^{(n)}\right]^{4}\right\}\right) .
\end{aligned}
$$

Multiply the equation by $\Lambda^{s} u^{(n+1)}$ and integrate on $\mathbb{R}$, then we have

$$
\begin{align*}
\frac{1}{2} \frac{\partial}{\partial_{t}}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}}^{2}= & -\left\langle\Lambda^{s}\left(u^{(n)} \partial_{x} u^{(n+1)}\right), \Lambda^{s} u^{(n+1)}\right\rangle-\frac{\beta_{0}}{\beta}\left\langle\Lambda^{s} \partial_{x} u^{(n+1)}, \Lambda^{s} u^{(n+1)}\right\rangle \\
& -\left(c-\frac{\beta_{0}}{\beta}\right)\left\langle\Lambda^{(s-2)} \partial_{x} u^{(n+1)}, \Lambda^{s} u^{(n+1)}\right\rangle \\
& +\left(\frac{b-3}{2}\right)\left\langle\Lambda^{(s-2)} \partial_{x}\left(\left[u^{(n)}\right]_{x}^{2}\right), \Lambda^{s} u^{(n+1)}\right\rangle \\
& -\frac{b}{2}\left\langle\Lambda^{(s-2)} \partial_{x}\left[u^{(n)}\right]^{2}, \Lambda^{s} u^{(n+1)}\right\rangle-\frac{\omega_{1}}{3 \alpha^{2}}\left\langle\Lambda^{(s-2)} \partial_{x}\left[u^{(n)}\right]^{3}, \Lambda^{s} u^{(n+1)}\right\rangle \\
& -\frac{\omega_{2}}{4 \alpha^{3}}\left\langle\Lambda^{(s-2)} \partial_{x}\left[u^{(n)}\right]^{4}, \Lambda^{s} u^{(n+1)}\right\rangle \tag{3.8}
\end{align*}
$$

Since $\Lambda^{s}$ is a symmetric operator and $\partial_{x}$ is skew-symmetric, we have

$$
\left\langle\Lambda^{s} \partial_{x} u^{(n+1)}, \Lambda^{s} u^{(n+1)}\right\rangle=\left\langle\Lambda^{(s-2)} \partial_{x} u^{(n+1)}, \Lambda^{s} u^{(n+1)}\right\rangle=0,
$$

such that (3.8) becomes

$$
\begin{align*}
\frac{\partial}{\partial_{t}}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}}^{2}=- & 2
\end{aligned} \begin{aligned}
s & \left.\left.\Lambda^{(n)} \partial_{x} u^{(n+1)}\right), \Lambda^{s} u^{(n+1)}\right\rangle+(b-3)\left\langle\Lambda^{(s-2)} \partial_{x}\left(\left[u^{(n)}\right]_{x}^{2}\right), \Lambda^{s} u^{(n+1)}\right\rangle \\
& -b\left\langle\Lambda^{(s-2)} \partial_{x}\left[u^{(n)}\right]^{2}, \Lambda^{s} u^{(n+1)}\right\rangle-\frac{2 \omega_{1}}{3 \alpha^{2}}\left\langle\Lambda^{(s-2)} \partial_{x}\left[u^{(n)}\right]^{3}, \Lambda^{s} u^{(n+1)}\right\rangle \\
& -\frac{\omega_{2}}{2 \alpha^{3}}\left\langle\Lambda^{(s-2)} \partial_{x}\left[u^{(n)}\right]^{4}, \Lambda^{s} u^{(n+1)}\right\rangle \\
& =I_{1}+I_{2}+I_{3}+I_{4}+I_{5} \tag{3.9}
\end{align*}
$$

Part $I_{1}$ : For all constant coefficient skew-symmetric differential polynomial $P$, and $f, g$ smooth enough, a commutator process is

$$
\begin{equation*}
\Lambda^{s}(f P g)=f P \Lambda^{s} g+\left[\Lambda^{s}, f\right] P g \tag{3.10}
\end{equation*}
$$

Then, taking $P$ as $\partial_{x}, f=u^{(n)}$, and $g=u^{(n+1)}$, we have

$$
\Lambda^{s}\left(u^{(n)} \partial_{x} u^{(n+1)}\right)=u^{(n)} \partial_{x} \Lambda^{s} u^{(n+1)}+\left[\Lambda^{s}, u^{(n)}\right] \partial_{x} u^{(n+1)} .
$$

So, $I_{1}$ becomes

$$
\begin{aligned}
& I_{1}=\left\langle\Lambda^{s}\left(u^{(n)} \partial_{x} u^{(n+1)}\right), \Lambda^{s} u^{(n+1)}\right\rangle=\int_{\mathbb{R}} \Lambda^{s}\left(u^{(n)} \partial_{x} u^{(n+1)}\right) \Lambda^{s} u^{(n+1)} d x \\
& =\int_{\mathbb{R}} u^{(n)} \partial_{x} \Lambda^{s} u^{(n+1)} \Lambda^{s} u^{(n+1)} d x+\int_{\mathbb{R}}\left[\Lambda^{s}, u^{(n)}\right] \partial_{x} u^{(n+1)} \Lambda^{s} u^{(n+1)} d x \\
& =I_{11}+I_{12}
\end{aligned}
$$

Using integration by parts,

$$
\begin{aligned}
\left|I_{11}\right| & =\frac{1}{2}\left|\int_{\mathbb{R}} u^{(n)} \partial_{x}\left(\Lambda^{s} u^{(n+1)}\right)^{2} d x\right| \\
& =\frac{1}{2}\left|\int_{\mathbb{R}} u_{x}^{(n)}\left(\Lambda^{s} u^{(n+1)}\right)^{2} d x\right| \\
& \leq \frac{1}{2}\left\|u_{x}^{(n)}\right\|_{L^{\infty}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2} .
\end{aligned}
$$

Moreover, commutator estimate and Sobolev embedding theorem ( $H^{s-1}(R) \hookrightarrow$ $L^{\infty}(R)$ )give that for all $s>\frac{1}{2}$, and some $C>0$,

$$
\begin{align*}
\left|I_{11}\right| & \leq C\left\|u^{(n)}\right\|_{L^{\infty}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2}  \tag{3.11}\\
& \leq C\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2} .
\end{align*}
$$

On the other hand, by Hölder's inequality

$$
\begin{aligned}
\left|I_{12}\right| & =\left|\int_{\mathbb{R}}\left[\Lambda^{s}, u^{(n+1)}\right] \partial_{x} u^{(n+1)} \Lambda^{s} u^{(n+1)} d x\right| \\
& \leq\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}}\left\|\left[\Lambda^{s}, u^{(n)}\right] \partial_{x} u^{(n+1)}\right\|_{L^{2}} .
\end{aligned}
$$

Applying the commutator estimate on $\left\|\left[\Lambda^{s}, u^{(n)}\right] \partial_{x} u^{(n+1)}\right\|_{L^{2}}$ and by Sobolev embedding theorem, there exists $C^{\prime}>0$ such that

$$
\begin{align*}
\left|I_{12}\right| & \leq C^{\prime}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}}\left(\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}{ }_{x}\right\|_{L^{\infty}}+\left\|u_{x}^{(n)}\right\|_{L^{\infty}}\left\|u_{x}^{(n+1)}\right\|_{H^{s-1}}\right) \\
& \leq C^{\prime}\left\|u^{(n+1)}\right\|_{H^{s}}\left(\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}+\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}\right)  \tag{3.12}\\
& \leq C^{\prime}\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2}
\end{align*}
$$

Combining (3.11) and (3.12), there exists $C_{1}>0$

$$
\begin{equation*}
\left|I_{1}\right|=\left|I_{11}+I_{12}\right| \leq\left|I_{11}\right|+\left|I_{12}\right| \leq C_{1}\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2} \tag{3.13}
\end{equation*}
$$

Part $I_{2}$ : There exists $C_{2}>0$, such that

$$
\begin{align*}
\left|I_{2}\right| & =\left|(b-3)\left\langle\Lambda^{s-2} \partial_{x}\left(u^{(n)}{ }_{x}\right)^{2}, \Lambda^{s} u^{(n+1)}\right\rangle\right| \\
& \leq C_{2}|b-3|\left\|\Lambda^{s-2} \partial_{x}\left(u^{(n)}{ }_{x}\right)^{2}\right\|_{L^{2}}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}} \\
& \leq C_{2}|b-3|\left\|\partial_{x}\left(u^{(n)}{ }_{x}\right)^{2}\right\|_{H^{s-2}}\left\|u^{(n+1)}\right\|_{H^{s}}  \tag{3.14}\\
& \leq C_{2}|b-3|\left\|\left(u^{(n)}{ }_{x}\right)^{2}\right\|_{H^{s-1}}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \leq C_{2}|b-3|\left\|u^{(n)}{ }_{x}\right\|_{H^{s-1}}^{2}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \leq C_{2}|b-3|\left\|u^{(n)}\right\|_{H^{s}}^{2}\left\|u^{(n+1)}\right\|_{H^{s}} .
\end{align*}
$$

Part $I_{3}$ : There exists $C_{3}>0$ such that

$$
\begin{align*}
\left|I_{3}\right|= & \left|b\left\langle\Lambda^{s-2} \partial_{x}\left[u^{(n)}\right]^{2}, \Lambda^{s} u^{(n+1)}\right\rangle\right| \\
& \leq C_{3} b\left\|\Lambda^{s-2} \partial_{x}\left[u^{(n)}\right]^{2}\right\|_{L^{2}}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}} \\
& \leq C_{3} b\left\|\partial_{x}\left[u^{(n)}\right]^{2}\right\|_{H^{s-2}}\left\|u^{(n+1)}\right\|_{H^{s}}  \tag{3.15}\\
& \leq C_{3} b\left\|u^{(n)}\right\|_{H^{s-1}}^{2}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \leq C_{3} b\left\|u^{(n)}\right\|_{H^{s}}^{2}\left\|u^{(n+1)}\right\|_{H^{s}} .
\end{align*}
$$

Part $I_{4}$ : There exists $C_{4}>0$ such that

$$
\begin{align*}
\left|I_{4}\right| & =\left|\frac{2 \omega_{1}}{3 \alpha^{2}}\left\langle\Lambda^{s-2} \partial_{x}\left[u^{(n)}\right]^{3}, \Lambda^{s} u^{(n+1)}\right\rangle\right| \\
& \leq C_{4}^{\prime}\left\|\Lambda^{s-2} \partial_{x}\left[u^{(n)}\right]^{3}\right\|_{L^{2}}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}} \\
& \leq C_{4}^{\prime}\left\|\partial_{x}\left[u^{(n)}\right]^{3}\right\|_{H^{s-2}}\left\|u^{(n+1)}\right\|_{H^{s}}  \tag{3.16}\\
& \leq C_{4}^{\prime}\left\|u^{(n)}\right\|_{H^{s-1}}^{3}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \leq C_{4}^{\prime}\left\|u^{(n)}\right\|_{H^{s}}^{3}\left\|u^{(n+1)}\right\|_{H^{s}} .
\end{align*}
$$

Part $I_{5}$ : There exists $C_{5}>0$ such that

$$
\begin{align*}
\left|I_{5}\right| & =\left|\frac{\omega_{2}}{2 \alpha^{3}}\left\langle\Lambda^{s-2} \partial_{x}\left[u^{(n)}\right]^{4}, \Lambda^{s} u^{(n+1)}\right\rangle\right| \\
& \leq C_{5}^{\prime}\left\|\Lambda^{s-2} \partial_{x}\left[u^{(n)}\right]^{4}\right\|_{L^{2}}\left\|\Lambda^{s} u^{(n+1)}\right\|_{L^{2}} \\
& \leq C_{5}^{\prime}\left\|\partial_{x}\left[u^{(n)}\right]^{4}\right\|_{H^{s-2}}\left\|u^{(n+1)}\right\|_{H^{s}}  \tag{3.17}\\
& \leq C_{5}^{\prime}\left\|u^{(n)}\right\|_{H^{s-1}}^{4}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \leq C_{5}^{\prime}\left\|u^{(n)}\right\|_{H^{s}}^{4}\left\|u^{(n+1)}\right\|_{H^{s}} .
\end{align*}
$$

From (3.13) - (3.17), we get

$$
\begin{align*}
& \frac{\partial}{\partial_{t}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2} \leq C_{1}\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2}+C_{2}|b-3|\left\|u^{(n)}\right\|_{H^{s}}^{2}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \quad+C_{3} b\left\|u^{(n)}\right\|_{H^{s}}^{2}\left\|u^{(n+1)}\right\|_{H^{s}}+C_{4}^{\prime}\left\|u^{(n)}\right\|_{H^{s}}^{3}\left\|u^{(n+1)}\right\|_{H^{s}}+C_{5}^{\prime}\left\|u^{(n)}\right\|_{H^{s}}^{4}\left\|u^{(n+1)}\right\|_{H^{s}} \tag{3.18}
\end{align*}
$$

From (3.18) and by redefining $C_{1}$ and $C_{2}$, we have
$\frac{\partial}{\partial_{t}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2} \leq C_{1}\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2}+C_{2}\left(\left\|u^{(n)}\right\|_{H^{s}}^{2}+\left\|u^{(n)}\right\|_{H^{s}}^{3}+\left\|u^{(n)}\right\|_{H^{s}}^{4}\right)\left\|u^{(n+1)}\right\|_{H^{s}}$.
Then, taking $C_{0}=\max \left\{C_{1}, C_{2}\right\}$, we obtain
$\frac{\partial}{\partial_{t}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2} \leq C_{0}\left(\left\|u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}^{2}+\left(\left\|u^{(n)}\right\|_{H^{s}}^{2}+\left\|u^{(n)}\right\|_{H^{s}}^{3}+\left\|u^{(n)}\right\|_{H^{s}}^{4}\right)\left\|u^{(n+1)}\right\|_{H^{s}}\right)$.

Without lost of generality, suppose $C_{0} \geq 2$. Let's fix a $T>0$ so that

$$
\begin{equation*}
C_{0}^{5} \max \left\{\left\|u_{0}\right\|^{3}, 1\right\} T<1 \tag{3.20}
\end{equation*}
$$

Now, we claim that for all $n \in \mathbb{N}$

$$
\begin{equation*}
\left\|u^{(n)}\right\|_{H^{s}} \leq C_{0}\left\|u_{0}\right\|_{H^{s}}, \quad \forall t \in[0, T] \tag{3.21}
\end{equation*}
$$

We prove this claim by using an inductive argument. For $\mathrm{n}=0$, from (3.7) and $u^{(0)}=0$, we have

$$
\partial_{t} u^{(1)}+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x} u^{(1)}=0
$$

which implies that

$$
\frac{d}{d t}\left\|u^{(1)}\right\|_{H^{s}}^{2}=0
$$

by (3.7) and $\left.u^{(1)}\right|_{t=0}=S_{1} u_{0}$, then

$$
u^{(1)}=S_{1} u_{0}, \quad \forall t \in \mathbb{R} .
$$

Therefore, we get

$$
\left\|u^{(1)}\right\|_{H^{s}}=\left\|S_{1} u_{0}\right\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}} \leq C_{0}\left\|u_{0}\right\|_{H^{s}}
$$

For fixed $n \in \mathbb{N}$, we assume

$$
\begin{equation*}
\sup _{0 \leq t \leq T}\left\|u^{(n)}\right\|_{H^{s}} \leq C_{0}\left\|u_{0}\right\|_{H^{s}} \tag{3.22}
\end{equation*}
$$

Applying Gronwall's inequality to (3.19),

$$
\begin{equation*}
\left\|u^{(n+1)}\right\|_{H^{s}} \leq e^{\frac{1}{2} C_{0} \int_{0}^{t}\left\|u^{(n)}\right\|_{H^{s}} d \tau}\left(\left\|u^{(n+1)}(0)\right\|_{H^{s}}+\frac{1}{2} C_{0} \int_{0}^{t}\left(\left\|u^{(n)}\right\|_{H^{s}}^{2}+\left\|u^{(n)}\right\|_{H^{s}}^{3}+\left\|u^{(n)}\right\|_{H^{s}}^{4}\right) d \tau\right) \tag{3.23}
\end{equation*}
$$

Now, if we consider our assumption (3.22), then we have

$$
\left\|u^{(n+1)}\right\|_{H^{s}} \leq e^{\frac{C_{0}^{2}}{2}\left\|u_{0}\right\|_{H^{s} T}}\left(\left\|u^{(n+1)}(0)\right\|_{H^{s}}+\frac{1}{2}\left(C_{0}^{3}\left\|u_{0}\right\|_{H^{s}}^{2}+C_{0}^{4}\left\|u_{0}\right\|_{H^{s}}^{3}+C_{0}^{5}\left\|u_{0}\right\|_{H^{s}}^{4}\right) T\right)
$$

Also, we have

$$
\left\|u^{(n+1)}(0)\right\|_{H^{s}}=\left\|S_{n+1} u_{0}\right\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}}
$$

so

$$
\begin{gathered}
\left\|u^{(n+1)}\right\|_{H^{s}} \leq e^{\frac{C_{0}^{2}}{2}\left\|u_{0}\right\|_{H^{s} T}}\left(\left\|u_{0}\right\|_{H^{s}}+\frac{1}{2}\left(C_{0}^{3}\left\|u_{0}\right\|_{H^{s}}^{2}+C_{0}^{4}\left\|u_{0}\right\|_{H^{s}}^{3}+C_{0}^{5}\left\|u_{0}\right\|_{H^{s}}^{4}\right) T\right) \\
=e^{\frac{C_{0}^{2}}{2}\left\|u_{0}\right\|_{H^{s}} T}\left[1+\frac{T}{2}\left(C_{0}^{3}\left\|u_{0}\right\|_{H^{s}}+C_{0}^{4}\left\|u_{0}\right\|_{H^{s}}^{2}+C_{0}^{5}\left\|u_{0}\right\|_{H^{s}}^{3}\right)\right]\left\|u_{0}\right\|_{H^{s}} .
\end{gathered}
$$

From our assumption (3.20) and $C_{0} \geq 2$, we have

$$
\begin{aligned}
& C_{0}^{2}\left\|u_{0}\right\|_{H^{s}} T \leq C_{0}^{2} \max \left\{\left\|u_{0}\right\|^{3}, 1\right\} T<\frac{1}{C_{0}^{3}} \leq \frac{1}{8} \\
& C_{0}^{3}\left\|u_{0}\right\|_{H^{s}} T \leq C_{0}^{3} \max \left\{\left\|u_{0}\right\|^{3}, 1\right\} T<\frac{1}{C_{0}^{2}} \leq \frac{1}{4}
\end{aligned}
$$

and

$$
C_{0}^{4}\left\|u_{0}\right\|_{H^{s}}^{2} T \leq C_{0}^{4} \max \left\{\left\|u_{0}\right\|^{3}, 1\right\} T<\frac{1}{C_{0}} \leq \frac{1}{2}
$$

Thus, we obtain

$$
\left\|u^{(n+1)}\right\|_{H^{s}} \leq \frac{15}{8} e^{\frac{1}{16}}\left\|u_{0}\right\|_{H^{s}} \leq 2\left\|u_{0}\right\|_{H^{s}} \leq C_{0}\left\|u_{0}\right\|_{H^{s}}
$$

which implies that the claim (3.21) holds for $\forall n \in \mathbb{N}$. Therefore, the sequence $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ is uniformly bounded in $C\left([0, T] ; H^{s}\right)$ and then $\left\{u^{(n)} \partial_{x} u^{(n+1)}\right\}_{n \in \mathbb{N}}$ is uniformly bounded in $C\left([0, T] ; H^{s-1}\right)$ where we use the fact that

$$
\begin{aligned}
\left\|u^{(n)} \partial_{x} u^{(n+1)}\right\|_{H^{s-1}} & \leq C\left\|u^{(n)}\right\|_{H^{s-1}}\left\|\partial_{x} u^{(n+1)}\right\|_{H^{s-1}} \\
& \leq C\left\|u^{(n)}\right\|_{H^{s-1}}\left\|u^{(n+1)}\right\|_{H^{s}} \\
& \leq C C_{0}^{2}\left\|u_{0}\right\|_{H^{s}}^{2}
\end{aligned}
$$

Thanks to the equation (3.7) equation, we obtain that $\partial_{t} u^{(n+1)}$ is uniformly bounded in $C\left([0, T] ; H^{s-1}\right)$. Therefore we deduce that $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}^{\infty}$ is uniformly bounded in $C\left([0, T] ; H^{s}\right) \cap C^{1}\left([0, T] ; H^{s-1}\right)$.

Step 3: In this step, we will prove that $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ is a Cauchy sequence in $C\left([0, T] ; H^{s-1}\right)$. In fact, from (3.7), we have for all $n, m \in \mathbb{N}$

$$
\begin{align*}
& \partial_{t} u^{(m+n+1)}+u^{(m+n)} \partial_{x} u^{(m+n+1)}+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right)\left(1-\partial_{x}^{2}\right)^{-1}\right] \partial_{x} u^{(m+n+1)} \\
& =-\left(1-\partial_{x}^{2}\right)^{-1} \partial_{x}\left\{\left(\frac{3-b}{2}\right) \partial_{x}\left[u^{(m+n)}\right]^{2}+\frac{b}{2}\left[u^{(m+n)}\right]^{2}+\frac{\omega_{1}}{3 \alpha^{2}}\left[u^{(m+n)}\right]^{3}+\frac{\omega_{2}}{4 \alpha^{3}}\left[u^{(m+n)}\right]^{4}\right\}, \tag{3.24}
\end{align*}
$$

and

$$
\begin{align*}
& \partial_{t} u^{(n+1)}+u^{(n)} \partial_{x} u^{(n+1)}+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right)\left(1-\partial_{x}^{2}\right)^{-1}\right] \partial_{x} u^{(n+1)} \\
& =-\left(1-\partial_{x}^{2}\right)^{-1} \partial_{x}\left\{\left(\frac{3-b}{2}\right) \partial_{x}\left[u^{(n)}\right]^{2}+\frac{b}{2}\left[u^{(n)}\right]^{2}+\frac{\omega_{1}}{3 \alpha^{2}}\left[u^{(n)}\right]^{3}+\frac{\omega_{2}}{4 \alpha^{3}}\left[u^{(n)}\right]^{4}\right\} . \tag{3.25}
\end{align*}
$$

From (3.24) and (3.25), we get

$$
\begin{align*}
& \partial_{t}\left(u^{(m+n+1)}-u^{(n+1)}\right)+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right)\left(1-\partial_{x}^{2}\right)^{-1}\right] \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right) \\
& =-u^{(m+n)} \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right)-\left(u^{(m+n)}-u^{(n)}\right) \partial_{x} u^{(n+1)} \\
& -\left(1-\partial_{x}^{2}\right)^{-1} \partial_{x}\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x}\left[u^{(m+n)}\right]^{2}-\partial_{x}\left[u^{(n)}\right]^{2}\right)+\frac{b}{2}\left(\left[u^{(m+n)}\right]^{2}-\left[u^{(n)}\right]^{2}\right)\right. \\
& \left.+\frac{\omega_{1}}{3 \alpha^{2}}\left(\left[u^{(m+n)}\right]^{3}-\left[u^{(n)}\right]^{3}\right)+\frac{\omega_{2}}{4 \alpha^{3}}\left(\left[u^{(m+n)}\right]^{4}-\left[u^{(n)}\right]^{4}\right)\right\} . \tag{3.26}
\end{align*}
$$

Then, applying $\Lambda^{s-1}$ operator to the equation, we obtain the following

$$
\begin{align*}
& \partial_{t} \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right)+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right)\left(1-\partial_{x}^{2}\right)^{-1}\right] \partial_{x} \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) \\
& =-\Lambda^{s-1}\left\{u^{(m+n)} \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right)\right\}-\Lambda^{s-1}\left\{\left(u^{(m+n)}-u^{(n)}\right) \partial_{x} u^{(n+1)}\right\} \\
& -\Lambda^{s-1}\left(1-\partial_{x}^{2}\right)^{-1} \partial_{x}\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x} u^{(m+n)}-\partial_{x} u^{(n)}\right)\left(\partial_{x} u^{(m+n)}+\partial_{x} u^{(n)}\right)\right. \\
& +\frac{b}{2}\left(u^{(m+n)}-u^{(n)}\right)\left(u^{(m+n)}+u^{(n)}\right) \\
& +\frac{\omega_{1}}{3 \alpha^{2}}\left(u^{(m+n)}-u^{(n)}\right)\left(\left[u^{(m+n)}\right]^{2}+u^{(m+n)} u^{(n)}+\left[u^{(n)}\right]^{2}\right)+ \\
& \left.\frac{\omega_{2}}{4 \alpha^{3}}\left(u^{(m+n)}-u^{(n)}\right)\left(\left[u^{(m+n)}\right]^{3}+\left[u^{(m+n)}\right]^{2} u^{(n)}+u^{(m+n)}\left[u^{(n)}\right]^{2}+\left[u^{(n)}\right]^{3}\right)\right\} \tag{3.27}
\end{align*}
$$

If we multiply this equation by $\Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right)$ and take an integration on $\mathbb{R}$, we have

$$
\begin{align*}
& \frac{1}{2} \frac{d}{d t}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}^{2} \\
& =\int_{\mathbb{R}} \Lambda^{s-1}\left[u^{(m+n)} \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right)\right] \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& -\int_{\mathbb{R}} \Lambda^{s-1}\left[\left(u^{(m+n)}-u^{(n)}\right) \partial_{x} u^{(n+1)}\right] \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& -\int_{\mathbb{R}} \Lambda^{s-3} \partial_{x}\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x} u^{(m+n)}-\partial_{x} u^{(n)}\right)\left(\partial_{x} u^{(m+n)}+\partial_{x} u^{(n)}\right)\right. \\
& +\left(u^{(m+n)}-u^{(n)}\right)\left[\frac{b}{2}\left(u^{(m+n)}+u^{(n)}\right)+\frac{\omega_{1}}{3 \alpha^{2}}\left(\left[u^{(m+n)}\right]^{2}+u^{(m+n)} u^{(n)}+\left[u^{(n)}\right]^{2}\right)\right. \\
& \left.\left.+\frac{\omega_{2}}{4 \alpha^{3}}\left(\left[u^{(m+n)}\right]^{3}+\left[u^{(m+n)}\right]^{2} u^{(n)}+u^{(m+n)}\left[u^{(n)}\right]^{2}+\left[u^{(n)}\right]^{3}\right)\right]\right\} \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) \\
& =I_{1}+I_{2}+I_{3} . \tag{3.28}
\end{align*}
$$

Apply commutator estimate to $I_{1}$, we obtain

$$
\begin{aligned}
& \int_{\mathbb{R}} \Lambda^{s-1}\left[u^{(m+n)} \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right)\right] \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& =\int_{\mathbb{R}} u^{(m+n)} \partial_{x} \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right)\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& +\int_{\mathbb{R}}\left[\Lambda^{s-1}, u^{(m+n)}\right] \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right) \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& =I_{11}+I_{12}
\end{aligned}
$$

Then, we have

$$
\begin{align*}
& \left|I_{11}\right| \leq \frac{1}{2}\left\|u^{(m+n)}{ }_{x}\right\|_{L^{\infty}}\left\|\Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right)\right\|_{L^{2}}^{2} \\
& \leq \frac{1}{2}\left\|u^{(m+n)}{ }_{x}\right\|_{L^{\infty}}\left\|\left(u^{(m+n+1)}-u^{(n+1)}\right)\right\|_{H^{s-1}}^{2}  \tag{3.29}\\
& \leq C\left\|u^{(m+n)}\right\|_{H^{s}}\left\|\left(u^{(m+n+1)}-u^{(n+1)}\right)\right\|_{H^{s-1}}^{2}
\end{align*}
$$

and by Banach algebra

$$
\begin{align*}
& \left|I_{12}\right| \leq\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|\left[\Lambda^{s-1}, u^{(m+n)}\right] \partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right)\right\|_{L^{2}} \\
& \leq C\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}{ }_{x}\right\|_{H^{s-1}}\left\|\partial_{x}\left(u^{(m+n+1)}-u^{(n+1)}\right)\right\|_{H^{s-2}}  \tag{3.30}\\
& \leq C\left\|u^{(m+n)}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}^{2} .
\end{align*}
$$

From (3.29) and (3.30), we get

$$
\begin{aligned}
\left|I_{11}\right|+\left|I_{12}\right| & \leq C\left\|u^{(m+n)}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}^{2} \\
& \leq C\left\|u_{0}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}^{2}
\end{aligned}
$$

Also, applying commutator estimate to $I_{2}$, we have

$$
\begin{aligned}
& I_{2}=\int_{\mathbb{R}} \Lambda^{s-1}\left[\left(u^{(m+n)}-u^{(n)}\right) \partial_{x} u^{(n+1)}\right] \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& \quad=\int_{\mathbb{R}}\left(u^{(m+n)}-u^{(n)}\right) \partial_{x} \Lambda^{s-1} u^{(n+1)} \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& +\int_{\mathbb{R}}\left[\Lambda^{s-1}, u^{(m+n)}-u^{(n)}\right] \partial_{x} u^{(n+1)} \Lambda^{s-1}\left(u^{(m+n+1)}-u^{(n+1)}\right) d x \\
& = \\
& I_{21}+I_{22} .
\end{aligned}
$$

Therefore, we obtain

$$
\begin{align*}
\left|I_{21}\right| & \leq\left\|u^{(m+n)}-u^{(n)}\right\|_{L^{2}}\left\|\partial_{x} \Lambda^{s-1} u^{(n+1)}\right\|_{L^{\infty}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}  \tag{3.31}\\
& \leq\left\|u^{(m+n)}-u^{(n)}\right\|_{L^{2}}\left\|u^{(n+1)}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}
\end{align*}
$$

and

$$
\begin{align*}
\left|I_{22}\right| & \leq\left\|\left[\Lambda^{s-1}, u^{(m+n)}-u^{(n)}\right] \partial_{x} u^{(n+1)}\right\|_{L^{2}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \\
& \leq C^{\prime}\left\|\partial_{x}\left(u^{(m+n)}-u^{(n)}\right)\right\|_{H^{s-1}}\left\|\partial_{x} u^{(n+1)}\right\|_{H^{s-2}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}  \tag{3.32}\\
& \leq C^{\prime}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \\
& \leq C^{\prime}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\left\|u^{(n+1)}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}
\end{align*}
$$

From (3.31) and (3.32), we have

$$
\begin{aligned}
\left|I_{21}\right|+\left|I_{22}\right| & \leq C\left\|u^{(n+1)}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}} \\
& \leq\left\|u_{0}\right\|_{H^{s}}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s} .}
\end{aligned}
$$

In addition, we get

$$
\begin{align*}
\left|I_{3}\right| & \leq \| \partial_{x}\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x} u^{(m+n)}-\partial_{x} u^{(n)}\right)\left(\partial_{x} u^{(m+n)}+\partial_{x} u^{(n)}\right)\right. \\
& +\left(u^{(m+n)}-u^{(n)}\right)\left[\frac{b}{2}\left(u^{(m+n)}+u^{(n)}\right)+\frac{\omega_{1}}{3 \alpha^{2}}\left(\left[u^{(m+n)}\right]^{2}+u^{(m+n)} u^{(n)}+\left[u^{(n)}\right]^{2}\right)\right. \\
& +\frac{\omega_{2}}{4 \alpha^{3}}\left(\left[u^{(m+n)}\right]^{3}+\left[u^{(m+n)}\right]^{2} u^{(n)}+u^{(m+n)}\left[u^{(n)}\right]^{2}\right. \\
& \left.\left.\left.+\left[u^{(n)}\right]^{3}\right)\right]\right\}\left\|_{H^{s-3}}\right\| u^{(m+n+1)}-u^{(n+1)} \|_{H^{s-1}}, \tag{3.33}
\end{align*}
$$

so

$$
\begin{aligned}
\left|I_{3}\right| & \leq \|\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x} u^{(m+n)}-\partial_{x} u^{(n)}\right)\left(\partial_{x} u^{(m+n)}+\partial_{x} u^{(n)}\right)\right. \\
& +\left(u^{(m+n)}-u^{(n)}\right)\left[\frac{b}{2}\left(u^{(m+n)}+u^{(n)}\right)+\frac{\omega_{1}}{3 \alpha^{2}}\left(\left[u^{(m+n)}\right]^{2}+u^{(m+n)} u^{(n)}+\left[u^{(n)}\right]^{2}\right)\right. \\
& +\frac{\omega_{2}}{4 \alpha^{3}}\left(\left[u^{(m+n)}\right]^{3}+\left[u^{(m+n)}\right]^{2} u^{(n)}+u^{(m+n)}\left[u^{(n)}\right]^{2}\right. \\
& \left.\left.\left.+\left[u^{(n)}\right]^{3}\right)\right]\right\}\left\|_{H^{s-2}}\right\| u^{(m+n+1)}-u^{(n+1)} \|_{H^{s-1}} .
\end{aligned}
$$

To simplify this inequality, we obtain

$$
\begin{aligned}
& \|\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x} u^{(m+n)}-\partial_{x} u^{(n)}\right)\left(\partial_{x} u^{(m+n)}+\partial_{x} u^{(n)}\right)\right. \\
& +\left(u^{(m+n)}-u^{(n)}\right)\left[\frac{b}{2}\left(u^{(m+n)}+u^{(n)}\right)+\frac{\omega_{1}}{3 \alpha^{2}}\left(\left[u^{(m+n)}\right]^{2}+u^{(m+n)} u^{(n)}+\left[u^{(n)}\right]^{2}\right)\right. \\
& \left.\left.+\frac{\omega_{2}}{4 \alpha^{3}}\left(\left[u^{(m+n)}\right]^{3}+\left[u^{(m+n)}\right]^{2} u^{(n)}+u^{(m+n)}\left[u^{(n)}\right]^{2}+\left[u^{(n)}\right]^{3}\right)\right]\right\} \|_{H^{s-2}} \\
& \leq\left(\frac{3-b}{2}\right)\left\|\partial_{x}\left(u^{(m+n)}-u^{(n)}\right)\right\|_{H^{s-2}}\left\|\partial_{x}\left(u^{(m+n)}+u^{(n)}\right)\right\|_{H^{s-2}} \\
& +C\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s-2}}\left\{\left\|u^{(m+n)}\right\|_{H^{s-2}}+\left\|u^{(n)}\right\|_{H^{s-2}}+\left\|u^{(m+n)}\right\|_{H^{s-2}}^{2}+\left\|u^{(n)}\right\|_{H^{s-2}}^{2}\right. \\
& +\left\|u^{(m+n)}\right\|_{H^{s-2}}\left\|u^{(n)}\right\|_{H^{s-2}}+\left\|u^{(m+n)}\right\|_{H^{s-2}}^{3}+\left\|u^{(m+n)}\right\|_{H^{s-2}}^{2}\left\|u^{(n)}\right\|_{H^{s-2}} \\
& \left.+\left\|u^{(m+n)}\right\|_{H^{s-2}}\left\|u^{(n)}\right\|_{H^{s-2}}^{2}+\left\|u^{(n)}\right\|_{H^{s-2}}^{3}\right\} \\
& \leq\left(\frac{3-b}{2}\right)\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s-1}}\left\|u^{(m+n)}+u^{(n)}\right\|_{H^{s-1}} \\
& +C\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s-1}}\left\{\left\|u^{(m+n)}\right\|_{H^{s-1}}+\left\|u^{(n)}\right\|_{H^{s-1}}+\left\|u^{(m+n)}\right\|_{H^{s-1}}^{2}+\left\|u^{(n)}\right\|_{H^{s-1}}^{2}\right. \\
& +\left\|u^{(m+n)}\right\|_{H^{s-1}}\left\|u^{(n)}\right\|_{H^{s-1}}+\left\|u^{(m+n)}\right\|_{H^{s-1}}^{3}+\left\|u^{(m+n)}\right\|_{H^{s-1}}^{2}\left\|u^{(n)}\right\|_{H^{s-1}} \\
& \left.+\left\|u^{(m+n)}\right\|_{H^{s-1}}\left\|u^{(n)}\right\|_{H^{s-1}}^{2}+\left\|u^{(n)}\right\|_{H^{s-1}}^{3}\right\} \\
& \leq C^{\prime}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s-1}}\left\{\left\|u^{(m+n)}\right\|_{H^{s-1}}+\left\|u^{(n)}\right\|_{H^{s-1}}+\left\|u^{(m+n)}\right\|_{H^{s-1}}^{2}+\left\|u^{(n)}\right\|_{H^{s-1}}^{2}\right. \\
& +\left\|u^{(m+n)}\right\|_{H^{s-1}}\left\|u^{(n)}\right\|_{H^{s-1}}+\left\|u^{(m+n)}\right\|_{H^{s-1}}^{3}+\left\|u^{(m+n)}\right\|_{H^{s-1}}^{2}\left\|u^{(n)}\right\|_{H^{s-1}} \\
& \left.+\left\|u^{(m+n)}\right\|_{H^{s-1}}\left\|u^{(n)}\right\|_{H^{s-1}}^{2}+\left\|u^{(n)}\right\|_{H^{s-1}}^{3}\right\},
\end{aligned}
$$

where we use $\left\|\partial_{x} u^{(n)}\right\|_{H^{s-1}} \leq\left\|u^{(n)}\right\|_{H^{s-1}}$ and $\left\|\partial_{x} u^{(n)}\right\|_{H^{1 / 2} \cap L^{\infty}} \leq\left\|u^{(n)}\right\|_{H^{s}}$, for all $s>3 / 2$. By using $\left\|u^{(n)}\right\|_{H^{s-1}}<C\left\|u_{0}\right\|_{H^{s-1}}$,

$$
\begin{align*}
& \|\left\{\left(\frac{3-b}{2}\right)\left(\partial_{x} u^{(m+n)}-\partial_{x} u^{(n)}\right)\left(\partial_{x} u^{(m+n)}+\partial_{x} u^{(n)}\right)\right. \\
& \left(u^{(m+n)}-u^{(n)}\right)\left[\frac{b}{2}\left(u^{(m+n)}+u^{(n)}\right)+\frac{\omega_{1}}{3 \alpha^{2}}\left(\left[u^{(m+n)}\right]^{2}+u^{(m+n)} u^{(n)}+\left[u^{(n)}\right]^{2}\right)\right. \\
& \left.\left.+\frac{\omega_{2}}{4 \alpha^{3}}\left(\left[u^{(m+n)}\right]^{3}+\left[u^{(m+n)}\right]^{2} u^{(n)}+u^{(m+n)}\left[u^{(n)}\right]^{2}+\left[u^{(n)}\right]^{3}\right)\right]\right\} \|_{H^{s-2}} \\
& \leq C^{\prime}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s-1}}\left(\left\|u_{0}\right\|_{H^{s-1}}+\left\|u_{0}\right\|_{H^{s-1}}^{2}+\left\|u_{0}\right\|_{H^{s-1}}^{3}\right) . \tag{3.34}
\end{align*}
$$

Combining (3.33) and (3.34), we have

$$
\begin{align*}
& \left|I_{3}\right| \leq C^{\prime}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s-1}}\left(\left\|u_{0}\right\|_{H^{s-1}}+\left\|u_{0}\right\|_{H^{s-1}}^{2}+\left\|u_{0}\right\|_{H^{s-1}}^{3}\right) \\
& \quad \leq C^{\prime}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\left(\left\|u_{0}\right\|_{H^{s}}+\left\|u_{0}\right\|_{H^{s}}^{2}+\left\|u_{0}\right\|_{H^{s}}^{3}\right) \tag{3.35}
\end{align*}
$$

Considering all results from (3.28) to (3.35), we get

$$
\begin{aligned}
& \begin{array}{l}
\frac{1}{2} \frac{d}{d t}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}^{2} \leq C\left\{\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}^{2}\left\|u_{0}\right\|_{H^{s}}\right. \\
\quad+\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\left\|u_{0}\right\|_{H^{s}} \\
\left.+\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\left(\left\|u_{0}\right\|_{H^{s}}+\left\|u_{0}\right\|_{H^{s}}^{2}+\left\|u_{0}\right\|_{H^{s}}^{3}\right)\right\} \\
\leq C\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}\left[\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}+\right. \\
\left.\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\right]\left(\left\|u_{0}\right\|_{H^{s}}+\left\|u_{0}\right\|_{H^{s}}^{2}+\left\|u_{0}\right\|_{H^{s}}^{3}\right)
\end{array} .
\end{aligned}
$$

So, we can have

$$
\frac{d}{d t}\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \leq C^{\prime}\left(\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}}+\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}}\right)
$$

Thanks to the Gronwall's inequality, we have $\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \leq e^{\int_{0}^{t} C^{\prime} d s}\left(\left\|u_{0}{ }^{(m+n+1)}-u_{0}{ }^{(n+1)}\right\|_{H^{s-1}}+\int_{0}^{t}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}} d s\right)$.

Then, we obtain

$$
\begin{equation*}
\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \leq C_{T}\left(\left\|u_{0}^{(m+n+1)}-u_{0}^{(n+1)}\right\|_{H^{s-1}}+\int_{0}^{t}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}} d \tau\right) \tag{3.36}
\end{equation*}
$$

Notice that

$$
u_{0}^{(m+n+1)}-u_{0}^{(n+1)}=S_{m+n+1} u_{0}-S_{n+1} u_{0}=1_{2 \times 2^{n+2} \leq|\xi| \leq 2 \times 2^{m+n+2}}(\xi) u_{0}(\xi) .
$$

Then, we have

$$
\begin{aligned}
\left\|u_{0}^{(m+n+1)}-u_{0}^{(n+1)}\right\|_{H^{s-1}}^{2} & =\int_{\mathbb{R}}\left(1+|\xi|^{2}\right)^{s-1}\left|\widehat{u_{0}}(\xi)\right|^{2} 1_{2^{n+3} \leq|\xi| \leq 2^{m+n+3}} d \xi \\
& =\int_{2^{n+3} \leq|\xi| \leq 2^{m+n+3}}\left(1+|\xi|^{2}\right)^{s-1}\left|\widehat{u_{0}}(\xi)\right|^{2} d \xi \\
& \leq 2^{-2(n+3)} \int_{2^{n+3} \leq|\xi| \leq 2^{m+n+3}}|\xi|^{2}\left(1+|\xi|^{2}\right)^{s-1}\left|\widehat{u_{0}}(\xi)\right|^{2} d \xi \\
& \leq 2^{-2 n} \frac{1}{8^{2}}\left\|u_{0}\right\|_{H^{s}}^{2},
\end{aligned}
$$

which implies that

$$
\left\|u_{0}^{(m+n+1)}-u_{0}^{(n+1)}\right\|_{H^{s-1}} \leq \frac{1}{2} 2^{-n}\left\|u_{0}\right\|_{H^{s} .} .
$$

Therefore, from (3.36) we get that $\forall t \in[0, T]$

$$
\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \leq C_{T}\left(2^{-n}+\int_{0}^{t}\left\|u^{(m+n)}-u^{(n)}\right\|_{H^{s}} d \tau\right)
$$

Applying induction argument, we can prove that for all $t \in[0, T]$

$$
\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} \leq \sum_{k=0}^{n} C_{T} 2^{-(n-k)}\left(C_{T} t\right)^{k} \frac{1}{k!}+C_{T}^{n+1} \frac{T^{n+1}}{(n+1)!}\left\|u^{(m)}-u^{(0)}\right\|_{L_{T}^{\infty} H^{s-1}}
$$

Due to the fact that $\left\|u^{(m+n)}-u^{(0)}\right\|_{L_{T}^{\infty} H^{s}} \leq\left\|u^{(m)}\right\|_{L_{T}^{\infty} H^{s}} \leq C_{0}\left\|u_{0}\right\|_{H^{s}}$ implies that for all $t \in[0, T], \forall m, n \in \mathbb{N}$ and some $C_{T}^{\prime}$ (independent of $\mathrm{m}, \mathrm{n}$ )

$$
\begin{align*}
\left\|u^{(m+n+1)}-u^{(n+1)}\right\|_{H^{s-1}} & \leq 2^{-n}\left(\sum_{k=0}^{n} C_{T} \frac{\left(2 C_{T} T\right)^{k}}{k!}+C_{0}\left\|u_{0}\right\|_{H^{s}} \frac{\left(2 C_{T} T\right)^{n+1}}{(n+1)!}\right)  \tag{3.37}\\
& \leq C_{T}^{\prime} 2^{-n}
\end{align*}
$$

Therefore, $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ is a Cauchy sequence in $C\left([0, T] ; H^{s-1}\right)$. Since the Banach space is complete $C\left([0, T] ; H^{s-1}\right)$, we get a limit $u \in C\left([0, T] ; H^{s-1}\right)$ such that $u^{(n)} \rightarrow u$ in $C\left([0, T] ; H^{s-1}\right)$.

Step 4: (Passing to the limit) Since $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}^{\infty}$ is uniformly bounded in $C\left([0, T] ; H^{s}\right) \cap C^{1}\left([0, T] ; H^{s-1}\right)$ and using interpolation inequality, we have for any $s^{\prime} \in\left(s-1, s^{\prime}\right)$,

$$
\begin{aligned}
\left\|u^{(m+n+1)}-u^{(n)}\right\|_{H^{s^{\prime}}} & \leq C\left\|u^{(m+n+1)}-u^{(n)}\right\|_{H^{s-1}}^{s-s^{\prime}}\left\|u^{(m+n+1)}-u^{(n)}\right\|_{H^{s}}^{1+s^{\prime}-s} \\
& \leq C\left\|u^{(m+n+1)}-u^{(n)}\right\|_{H^{s-1}}^{s-s^{\prime}}\left(\left\|u^{(m+n+1)}\right\|_{H^{s}}+\left\|u^{(n)}\right\|_{H^{s}}\right)^{1+s^{\prime}-s} \\
& \leq C\left\|u^{(m+n+1)}-u^{(n)}\right\|_{H^{s-1}}^{s-s^{\prime}}\left\|u_{0}\right\|_{H^{s}}^{1+s^{\prime}-s} \\
& \leq C^{\prime}\left\|u^{(m+n+1)}-u^{(n)}\right\|_{H^{s-1}}^{s-s^{\prime}}
\end{aligned}
$$

which along with (3.37) implies that $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ is a Cauchy sequence in $C\left([0, T] ; H^{s^{\prime}}\right)$, for all $s^{\prime} \in(s-1, s)$. By the uniqueness of the limit, we deduce that $u^{(n)} \rightarrow u$ in $C\left([0, T] ; H^{s^{\prime}}\right)$ which yields that $u^{(n)} \partial_{x} u^{(n+1)} \rightarrow u \partial_{x} u$ in $C\left([0, T] ; H^{s^{\prime}-1}\right)$, where $s^{\prime}>\frac{3}{2}$. Therefore, we have

$$
\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x} u^{(n+1)} \rightarrow\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x} u
$$

in $C\left([0, T] ; H^{s^{\prime}-1}\right)$, where $s^{\prime}>\frac{3}{2}$;

$$
p * \partial_{x}\left\{\frac{b}{2}\left[u^{(n)}\right]^{2}+\frac{\omega_{1}}{3 \alpha^{2}}\left[u^{(n)}\right]^{3}+\frac{\omega_{2}}{4 \alpha^{3}}\left[u^{(n)}\right]^{4}\right\} \rightarrow p * \partial_{x}\left\{\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}
$$

in $C\left([0, T] ; H^{s^{\prime}+1}\right)$ where $s^{\prime}>-\frac{1}{2}$; and

$$
p * \partial_{x}\left(\frac{3-b}{2}\right) \partial_{x}\left[u^{(n)}\right]^{2} \rightarrow p * \partial_{x}\left(\frac{3-b}{2}\right) \partial_{x} u^{2}
$$

in $C\left([0, T] ; H^{s^{\prime}}\right)$ where $s^{\prime}>\frac{1}{2}$. Then, from (3.25), it implies that $\left\{\partial_{t} u^{(n)}\right\}_{n \in \mathbb{N}}$ is a Cauchy sequence in $C\left([0, T] ; H^{s^{\prime}-1}\right)$, for any $s^{\prime}>\frac{3}{2}$ and there exists a $v \in$ $C\left([0, T] ; H^{s^{\prime}-1}\right)$ such that $\partial_{t} u^{(n)} \rightarrow v$ in $C\left([0, T] ; H^{s^{\prime}-1}\right)$, for all $\frac{3}{2}<s^{\prime}<s$.

On the other hand, since $u^{(n)}$ converges to $u$ in $C\left([0, T] ; H^{s^{\prime}}\right)$ and $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}^{\infty}$ is uniformly bounded in $C\left([0, T] ; H^{s}\right) \cap C^{1}\left([0, T] ; H^{s-1}\right)$ we conclude that $\partial_{t} u^{(n)} \rightarrow$ $\partial_{t} u$ in the sense of distribution which along with $\exists v$ in $C\left([0, T] ; H^{s^{\prime}-1}\right)$ such that $\partial_{t} u^{(n)} \rightarrow v$ in $C\left([0, T] ; H^{s^{\prime}-1}\right), \forall s^{\prime} \in\left(\frac{3}{2}, s\right)$ implies that $v=\partial_{t} u$ and $\partial_{t} u^{(n)} \rightarrow \partial_{t} u$ in $C\left([0, T] ; H^{s^{\prime}-1}\right)$ for any $s^{\prime} \in\left(\frac{3}{2}, s\right)$.

Therefore, it follows that

$$
\left\{\begin{align*}
\partial_{t} u+u \partial_{x} u+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] & \partial_{x} u  \tag{3.38}\\
& =-p * \partial_{x}\left\{\left(\frac{3-b}{2}\right) u_{x}^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}
\end{align*} \quad \begin{array}{rl}
\left.u\right|_{t=0}=u_{0}
\end{array}\right.
$$

in $C\left([0, T] ; H^{s^{\prime}-1}\right), \forall \frac{3}{2}<s^{\prime}<s$ and $u \in C\left([0, T] ; H^{s^{\prime}}\right) \cap C^{1}\left([0, T] ; H^{s^{\prime}-1}\right)$.
Moreover, by the Banach-Alaoglu Theorem, since $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}^{\infty}$ is uniformly bounded in $C\left([0, T] ; H^{s}\right) \cap C^{1}\left([0, T] ; H^{s-1}\right)$, there is a subsequence $\left\{u^{\left(n_{j}\right)}\right\}_{j \in \mathbb{N}}$ of $\left\{u^{(n)}\right\}_{n \in \mathbb{N}}$ such that

$$
u^{\left(n_{j}\right)} \stackrel{\text { weakly }}{\longrightarrow} u^{*}
$$

in $L^{2}\left([0, T] ; H^{s}\right)$. Also, thanks to $L^{\infty}\left([0, T] ; H^{s}\right) \hookrightarrow L^{2}\left([0, T] ; H^{s}\right)$, for $\forall t \in[0, T] ;$

$$
\begin{align*}
& u^{\left(n_{j}\right)} \stackrel{\text { weakly }}{\sim} u^{*} \quad \text { in } \quad H^{s}  \tag{3.39}\\
& \partial_{t} u^{\left(n_{j}\right)} \stackrel{\text { weakly }}{\longrightarrow} \partial_{t} u \quad \text { in } \quad H^{s-1}
\end{align*}
$$

which along with the fact that $u^{(n)} \rightarrow u$ in $C\left([0, T] ; H^{s^{\prime}}\right)$ implies

$$
u=u^{*} \in L^{\infty}\left([0, T] ; H^{s}\right) \cap \operatorname{Lip}\left([0, T] ; H^{s-1}\right) .
$$

Next, we claim that $u \in C_{w}\left([0, T] ; H^{s}\right)$ that means for all $\varphi \in H^{-s},[u(t), \varphi]_{H^{s} \times H^{-s}}$ is continuous with respect to $t \in[0, T]$.

In fact, $\forall \varphi \in H^{-s^{\prime}}$, for $s^{\prime} \in\left(\frac{3}{2}, s\right), u^{\left(n_{j}\right)} \rightharpoonup u$ in $C\left([0, T] ; H^{s^{\prime}}\right)$. Then,

$$
\left[u^{n_{j}}(t), \varphi\right]_{H^{s^{\prime}} \times H^{-s^{\prime}}} \rightarrow[u(t), \varphi]_{H^{s^{\prime}} \times H^{-s^{\prime}}} \quad \text { uniformly on } \quad[0, T] .
$$

While $H^{-s^{\prime}}$ is dense in $H^{-s}$, for $s^{\prime}<s$, we get from last result that for all $\psi \in H^{-s}$, $\left[u^{n_{j}}(t), \psi\right]_{H^{s} \times H^{-s}} \rightarrow[u(t), \psi]_{H^{s} \times H^{-s}}$ is uniformly on $[0, \mathrm{~T}]$. Since $\left[u^{n_{j}}(t), \psi\right]_{H^{s} \times H^{-s}} \in$ $C([0, T])$ and it is uniformly continuous, we obtain that its limit $[u(t), \psi]_{H^{s} \times H^{-s}}$ is uniformly continuous on $[0, T]$ which implies that the claim holds.

Up to a subsequence, we get from (3.39) that for fixed $t \in[0, T]$,

$$
\limsup _{n \rightarrow \infty}\left\|u^{(n)}(t)\right\|_{H^{s}} \geq\|u(t)\|_{H^{s}}
$$

which implies that

$$
\begin{gather*}
\|u(t)\|_{H^{s}} \leq \lim _{n \rightarrow \infty} \sup \left\|u^{(n)}(t)\right\|_{H^{s}}  \tag{3.40}\\
\leq\left\|u_{0}\right\|_{H^{s}} e^{\frac{C_{0}}{2}}\left\|u_{0}\right\|_{H^{s}}
\end{gather*}
$$

Hence, we have

$$
\begin{equation*}
\limsup _{t \rightarrow 0^{+}}\left\|u^{(n)}(t)\right\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}} \tag{3.41}
\end{equation*}
$$

On the other hand, from $u \in C_{w}\left([0, T] ; H^{s}\right)$ and the fact that

$$
\|f\|_{H^{s}}=\sup _{\|\psi\|_{H^{-s}} \leq 1}\left|\langle f, \psi\rangle_{H^{s} \times H^{-s}}\right|
$$

we get

$$
\liminf _{t \rightarrow 0^{+}}\|u(t)\|_{H^{s}} \geq\left\|u_{0}\right\|_{H^{s}}
$$

which along with (3.41) implies that

$$
\lim _{t \rightarrow 0^{+}}\|u(t)\|_{H^{s}}=\left\|u_{0}\right\|_{H^{s}}
$$

i.e. $\|u(t)\|_{H^{s}}$ is strongly right continuous at $\mathrm{t}=0$. Similarly we may get that $\|u(t)\|_{H^{s}}$ is strongly left continuous at $\mathrm{t}=0$ and so $\|u(t)\|_{H^{s}}$ is strongly continuous at $\mathrm{t}=0$. It remains to prove continuity of $\|u(t)\|_{H^{s}}$ at times other than the initial time.For this, let's first prove the uniqueness of the solution $u \in C\left([0, T] ; H^{s^{\prime}}\right) \cap L^{\infty}\left([0, T] ; H^{s}\right)$ with $\partial_{t} u \in C\left([0, T] ; H^{s^{\prime}-1}\right) \cap L^{\infty}\left([0, T] ; H^{s}\right)$

## Step 5: (Uniqueness)

It is easy to prove the uniqueness of the solution of the Rotational $b$-family of equations which process is similar to the one in Step 3.

In fact, assume that $u, v \in C\left([0, T] ; H^{s^{\prime}}\right) \cap L^{\infty}\left([0, T] ; H^{s}\right)$ with $\partial_{t} u, \partial_{t} v \in$ $C\left([0, T] ; H^{s^{\prime}-1}\right) \cap L^{\infty}\left([0, T] ; H^{s-1}\right)$ where $\frac{3}{2}<s^{\prime}<s$ and $\left.u\right|_{t=0}=\left.v\right|_{t=0}$ solve the Rotational $b$-family of equations, then we have

$$
\begin{aligned}
& \partial_{t}(u-v)+u \partial_{x}(u-v)+(u-v) \partial_{x} v+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x}(u-v) \\
& =-p * \partial_{x}\left\{\left(\frac{3-b}{2}\right)\left(\left(u_{x}^{2}-v_{x}^{2}\right)+\frac{b}{2}\left(u^{2}-v^{2}\right)+\frac{\omega_{1}}{3 \alpha^{2}}\left(u^{3}-v^{3}\right)+\frac{\omega_{2}}{4 \alpha^{3}}\left(u^{4}-v^{4}\right)\right\}\right.
\end{aligned}
$$

Let $u-v=r$, so

$$
\begin{aligned}
& \partial_{t} r+u \partial_{x} r+r \partial_{x} v+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right] \partial_{x} r \\
& =-p * \partial_{x}\left\{\left(\frac{3-b}{2}\right) r_{x}\left(u_{x}+v_{x}\right)+\frac{b}{2} r(u+v)+\frac{\omega_{1}}{3 \alpha^{2}} r\left(u^{2}+u v+v^{2}\right)+\frac{\omega_{2}}{4 \alpha^{3}} r(u+v)\left(u^{2}+v^{2}\right)\right\}
\end{aligned}
$$

Apply $\Lambda^{s^{\prime}-1}$ operator, multiply by $\Lambda^{s^{\prime}-1} r$ and integrate over $\mathbb{R}$, we have

$$
\begin{aligned}
& \left\langle\Lambda^{s^{\prime}-1} \partial_{t} r, \Lambda^{s^{\prime}-1} r\right\rangle+\left\langle\Lambda^{s^{\prime}-1} u \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle+\left\langle\Lambda^{s^{\prime}-1} r \partial_{x} v, \Lambda^{s^{\prime}-1} r\right\rangle+\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right]\left\langle\Lambda^{s^{\prime}-1} \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle \\
& =-\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left\{\left(\frac{3-b}{2}\right) r_{x}\left(u_{x}+v_{x}\right)+\frac{b}{2} r(u+v)+\frac{\omega_{1}}{3 \alpha^{2}} r\left(u^{2}+u v+v^{2}\right)+\frac{\omega_{2}}{4 \alpha^{3}} r(u+v)\left(u^{2}+v^{2}\right)\right\}, \Lambda^{s^{\prime}-1} r\right\rangle
\end{aligned}
$$

which is equivalent to

$$
\begin{aligned}
& \frac{d}{d t}\left\|\Lambda^{s^{\prime}-1} r\right\|_{L^{2}}^{2}=-2\left\langle\Lambda^{s^{\prime}-1} u \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle-2\left\langle\Lambda^{s^{\prime}-1} r \partial_{x} v, \Lambda^{s^{\prime}-1} r\right\rangle-2\left[\frac{\beta_{0}}{\beta}+\left(c-\frac{\beta_{0}}{\beta}\right) p *\right]\left\langle\Lambda^{s^{\prime}-1} \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle \\
& +(b-3)\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left(r_{x}\left(u_{x}+v_{x}\right)\right), \Lambda^{s^{\prime}-1} r\right\rangle-b\left\langle\Lambda^{s^{\prime}-3} \partial_{x}(r(u+v)), \Lambda^{s^{\prime}-1} r\right\rangle \\
& -\frac{2 \omega_{1}}{3 \alpha^{2}}\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left(r\left(u^{2}+u v+v^{2}\right)\right), \Lambda^{s^{\prime}-1} r\right\rangle-\frac{\omega_{2}}{2 \alpha^{3}}\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left(r(u+v)\left(u^{2}+v^{2}\right)\right), \Lambda^{s^{\prime}-1} r\right\rangle \\
& =-2\left\langle\Lambda^{s^{\prime}-1} u \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle-2\left\langle\Lambda^{s^{\prime}-1} r \partial_{x} v, \Lambda^{s^{\prime}-1} r\right\rangle \\
& +(b-3)\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left(r_{x}\left(u_{x}+v_{x}\right)\right), \Lambda^{s^{\prime}-1} r\right\rangle-b\left\langle\Lambda^{s^{\prime}-3} \partial_{x}(r(u+v)), \Lambda^{s^{\prime}-1} r\right\rangle \\
& -\frac{2 \omega_{1}}{3 \alpha^{2}}\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left(r\left(u^{2}+u v+v^{2}\right)\right), \Lambda^{s^{\prime}-1} r\right\rangle-\frac{\omega_{2}}{2 \alpha^{3}}\left\langle\Lambda^{s^{\prime}-3} \partial_{x}\left(r(u+v)\left(u^{2}+v^{2}\right)\right), \Lambda^{s^{\prime}-1} r\right\rangle \\
& =I_{1}+I_{2}+I_{3}+I_{4}+I_{5}+I_{6} .
\end{aligned}
$$

Then, we obtain each part such as

$$
\begin{aligned}
& \left|I_{1}\right|=\left|-2\left\langle\Lambda^{s^{\prime}-1} u \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle\right| \\
& =\left|2\left\langle u \partial_{x} \Lambda^{s^{\prime}-1} r+\left[\Lambda^{s^{\prime}-1}, u\right] \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle\right| \\
& =\left|2\left\langle u \partial_{x} \Lambda^{s^{\prime}-1} r, \Lambda^{s^{\prime}-1} r\right\rangle+\left\langle\left[\Lambda^{s^{\prime}-1}, u\right] \partial_{x} r, \Lambda^{s^{\prime}-1} r\right\rangle\right| \\
& =\left|2 \int_{\mathbb{R}} u \partial_{x} \Lambda^{s^{\prime}-1} r \Lambda^{s^{\prime}-1} r d x+\int_{\mathbb{R}}\left[\Lambda^{s^{\prime}-1}, u\right] \partial_{x} r \Lambda^{s^{\prime}-1} r d x\right| \\
& \leq 2\left\|u_{x}\right\|_{L^{\infty}}\|r\|_{H^{s^{\prime}-1}}^{2}+\left\|\Lambda^{s^{\prime}-1} r\right\|_{L^{2}}\left\|\left[\Lambda^{s^{\prime}-1}, u\right] \partial_{x} r\right\|_{L^{2}} \\
& \leq C\|u\|_{H^{s^{\prime}}}\|r\|_{H^{s^{\prime}-1}}^{2}+\|r\|_{H^{s^{\prime}-1}}\left(\|u\|_{H^{s^{\prime}-1}}\left\|r_{x}\right\|_{L^{\infty}}+\left\|u_{x}\right\|_{L^{\infty}}\|r\|_{H^{s^{\prime}-1}}\right) \\
& \leq C\|u\|_{H^{s^{\prime}}}\|r\|_{H^{s^{\prime}-1}}^{2}+\|r\|_{H^{s^{\prime}-1}}\left(\|u\|_{H^{s^{\prime}-1}}\|r\|_{H^{s^{\prime}-1}}+\|u\|_{H^{s^{\prime}}}\|r\|_{H^{s^{\prime}-1}}\right) \\
& \leq C\|u\|_{H^{s^{\prime}}}\|r\|_{H^{s^{\prime}-1}}^{2}, \\
& \left|I_{2}\right|=2\left|\int_{\mathbb{R}} \Lambda^{s^{\prime}-1}\left(r \partial_{x} v\right) \Lambda^{s^{\prime}-1} r d x\right| \\
& \leq C\left\|r \partial_{x} v\right\|_{H^{s^{\prime}-1}}\|r\|_{H^{s^{\prime}-1}} \\
& \leq C\|v\|_{H^{s^{\prime}}}\|r\|_{H^{s^{\prime}-1}}^{2}, \\
& \left|I_{3}\right|=|b-3| \mid \int_{\mathbb{R}} \Lambda^{s^{\prime}-3} \partial_{x}\left(r_{x}\left(u_{x}+v_{x}\right)\right) \Lambda^{s^{\prime}-1} r d x \\
& \leq C|b-3|\left\|\partial_{x}\left(r_{x}\left(u_{x}+v_{x}\right)\right)\right\|_{H^{s^{\prime}-3}}\|r\|_{H^{s^{\prime}-1}} \\
& \leq C\left\|\left(r_{x}\left(u_{x}+v_{x}\right)\right)\right\|_{H^{s^{\prime}-2}}\|r\|_{H^{s^{\prime}-1}} \\
& \leq C\|u+v\|_{H^{s^{\prime}-1}}\|r\|_{H^{s^{\prime}-1}}^{2}, \\
& \left|I_{4}\right|=b\left|\int_{\mathbb{R}} \Lambda^{s^{\prime}-3} \partial_{x}(r(u+v)) \Lambda^{s^{\prime}-1} r d x\right| \\
& \leq C\|(r(u+v))\|_{H^{s^{\prime}-2}}\|r\|_{H^{s^{\prime}-1}} \\
& \leq C\|u+v\|_{H^{s^{\prime}-1}}\|r\|_{H^{s^{\prime}-1}}^{2}, \\
& \left|I_{5}\right|=\left|\frac{2 \omega_{1}}{3 \alpha^{2}} \int_{\mathbb{R}} \Lambda^{s^{\prime}-3} \partial_{x}\left(r\left(u^{2}+u v+v^{2}\right)\right) \Lambda^{s^{\prime}-1} r d x\right| \\
& \leq C\|r\|_{H^{s^{\prime}-2}}\left\|u^{2}+u v+v^{2}\right\|_{H^{s^{\prime}-2}}\|r\|_{H^{s^{\prime}-1}} \\
& \leq C\|r\|_{H^{s^{\prime}-1}}^{2}\left(\|u\|_{H^{s^{\prime}-1}}^{2}+\|u\|_{H^{s^{\prime}-1}}\|v\|_{H^{s^{\prime}-1}}+\|v\|_{H^{s^{\prime}-1}}^{2}\right),
\end{aligned}
$$

$$
\begin{aligned}
\left|I_{6}\right| & =\left|\frac{\omega_{2}}{2 \alpha^{3}} \int_{\mathbb{R}} \Lambda^{s^{\prime}-3} \partial_{x}\left(r(u+v)\left(u^{2}+v^{2}\right)\right) \Lambda^{s^{\prime}-1} r d x\right| \\
& \leq C\|r\|_{H^{s^{\prime}-1}}^{2}\|u+v\|_{H^{s^{\prime}-2}}\left\|u^{2}+v^{2}\right\|_{H^{s^{\prime}-2}} \\
& \leq C\|r\|_{H^{s^{\prime}-1}}^{2}\left(\|u\|_{H^{s^{\prime}-1}}^{3}+\|u\|_{H^{s^{\prime}-1}}\|v\|_{H^{s^{\prime}-1}}{ }^{2}+\|u\|_{H^{s^{\prime}-1}}^{2}\|v\|_{H^{s^{\prime}-1}}+\|v\|_{H^{s^{\prime}-1}}^{3}\right) .
\end{aligned}
$$

From $I_{1}$ to $I_{6}$, we get

$$
\begin{aligned}
& \frac{d}{d t}\|r\|_{H^{s^{\prime}-1}}^{2} \leq C\|r\|_{H^{s^{\prime}-1}}^{2}\left[\|u\|_{H^{s^{\prime}}}+\|v\|_{H^{s^{\prime}}}+\|u\|_{H^{s^{\prime}}}^{2}+\|u\|_{H^{s^{\prime}}}\|v\|_{H^{s^{\prime}}}\right. \\
& \quad+\|v\|_{H^{s^{\prime}}}^{2}+\|u\|_{H^{s^{\prime}}}^{3}+\|v\|_{H^{s^{\prime}}}^{3}+\|u\|_{H^{s^{\prime}}}\|v\|_{H^{s^{\prime}}}^{2}+\|u\|_{H^{s^{\prime}}}^{2}\|v\|_{H^{s^{\prime}}} .
\end{aligned}
$$

When Gronwall's inequality is applied to the last inequality, we conclude that $\forall t \in$ $[0, t]$,

$$
\|u-v\|_{H^{s^{\prime}-1}} \equiv 0
$$

which implies $u \equiv v, \forall x \in \mathbb{R}, t \in[0, T]$.
Step 6: (Conclusion)
Back to the proof of the fact that $u \in C\left([0, T] ; H^{s}\right)$, we have known that $\|u(t)\|_{H^{s}}$ is continuous at $\mathrm{t}=0$. Then $\forall T_{0} \in[0, T]$ and the solution $u\left(\cdot, T_{0}\right)$ at the fixed time $T_{0} u\left(\cdot, T_{0}\right) \equiv u_{0}{ }^{T_{0}} \in H^{s}((R)$ and from (3.40) we obtain

$$
\left\|u_{0}^{T_{0}}\right\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}} e^{\frac{C_{0}{ }^{2}}{2}\left\|u_{0}\right\|_{H^{s}} T_{0}}
$$

So we take $u_{0}{ }^{T}$ as initial data and construct a forward and backward -in-time solution by solving (3.7). We obtain approximation solution $u_{T_{0}}^{(n)}(t)$ and its limit $u_{T_{0}}(t)$ which solves the Rotational $b$-family of equations with initial data

$$
\left.u_{T_{0}}(t)\right|_{t=0}=u_{0}{ }^{T_{0}} \equiv u\left(\cdot, T_{0}\right)
$$

with $u_{T_{0}} \in C\left(\left[0, T_{1}\right] ; H^{s^{\prime}}\right) \cap L^{\infty}\left(\left[0, T_{1}\right] ; H^{s}\right)$ and $\partial_{t} u_{T_{0}} \in C\left(\left[0, T_{1}\right] ; H^{s^{\prime}-1}\right) \cap L^{\infty}\left(\left[0, T_{1}\right] ; H^{s}\right)$ for some positive time $T_{1}>0$ and then $\left\|u_{T_{0}}(t)\right\|_{H^{s}}$ is continuous at $\mathrm{t}=0$.

By the uniqueness, we get that

$$
u_{T_{0}}(t)=u\left(t+T_{0}\right) \quad \text { on } \quad\left[T_{0}-T_{1}, T_{0}+T_{1}\right]
$$

which implies that $u(t)$ is continuous at $t=T_{0}$. Therefore, we obtain that $u \in$ $C\left([0, T] ; H^{s}\right)$.

Return to the original Rotational b-family of equations and define

$$
\|f\|_{X_{\mu}^{s+1}}^{2} \stackrel{\text { def }}{=}\|f\|_{H^{s}}^{2}+\mu \beta\left\|f_{x}\right\|_{H^{s}}^{2}
$$

where $\mu>0, \beta>0$. For some $\mu_{0}>0$ and $M>0$, we define the Camassa-Holm regime $\mathcal{P}_{\mu_{0}, M}:=\left\{(\varepsilon, \mu): 0<\mu \leq \mu_{0}, 0<\varepsilon \leq M \sqrt{\mu}\right\}$. Then, we have the following corollary.

Corollary 3.2.2. Let $u_{0} \in H^{s+1}(\mathbb{R}), \mu_{0}>0$ and $M>0, s>\frac{3}{2}$. Then, there exist $T>0$ and a unique family of solutions $\left.\left(u_{\varepsilon, \mu}\right)\right|_{(\varepsilon, \mu)} \in \mathcal{P}_{\mu_{0}, M}$ in $C\left(\left[0, \frac{T}{\varepsilon}\right] ; X^{s+1}(\mathbb{R})\right) \cap$ $C^{1}\left(\left[0, \frac{T}{\varepsilon} ; X^{s}(\mathbb{R})\right]\right)$ to the Cauchy problem

$$
\left\{\begin{array}{l}
\partial_{t} u-\beta \mu \partial_{t} u_{x x}+c u_{x}+(b+1) \alpha \varepsilon u u_{x}-\beta_{0} \mu u_{x x x}+ \\
\\
\omega_{1} \varepsilon^{2} u^{2} u_{x}+\omega_{2} \varepsilon^{3} u^{3} u_{x}=\alpha \beta \varepsilon \mu\left(b u_{x} u_{x x}+u u_{x x x}\right), \\
\left.u\right|_{t=0}=u_{0} .
\end{array}\right.
$$

## CHAPTER 4

## WAVE BREAKING PHENOMENA

In this chapter, we investigated effects of higher nonlinear terms on finite-time blow-up solutions and wave breaking phenomena is investigated for the rotational $b$-family of equations. Since the rotational $b$-family of equations does not have enough conserved quantities, it leads us to find different approach.

### 4.1 Finite-time blow-up solutions

The blow-up criterion and wave-breaking criterion for the rotational $b$-family of equations have been proved in this section.

Theorem 4.1.1. (Blow-up criterion) Let $u_{0} \in H^{s}, s>\frac{3}{2}$ and $u$ be the corresponding solution to (3.6). Assume $T_{u_{0}}^{*}$ is the maximal time of existence. Then,

$$
\begin{equation*}
T_{u_{0}}^{*}<\infty \Longrightarrow \int_{0}^{T_{u_{0}}^{*}}\left\|\partial_{x} u(\tau)\right\|_{L^{\infty}} d \tau=+\infty \tag{4.1}
\end{equation*}
$$

Approach of the proof for this theorem base on the following propositions.
Proposition 4.1.2. [14, 15] Support that $s>-\frac{d}{2}$. Let $v$ be a vector field such that $\nabla v$ belongs to $L^{1}\left([0,1] ; H^{s-1}\right)$ if $s>1+\frac{d}{2}$ or to $L^{1}\left([0,1] ; H^{\frac{d}{2}} \cap L^{\infty}\right)$ otherwise. Suppose also that $f_{0} \in H^{s}, F \in L^{1}\left([0,1] ; H^{s}\right)$ and that $f \in L^{\infty}\left([0, T] ; S^{\prime}\right)$ solves the $d$ - dimensional linear transport equations

$$
\left\{\begin{array}{l}
\partial_{t} f+v \cdot \nabla f=F \\
\left.f\right|_{t=0}=f_{0}
\end{array}\right.
$$

Then $f \in C\left([0, T] ; H^{s}\right)$. More precisely, there exists a constant $C$ depending only on $s, p$ and $d$, and such that the following statements hold:
(1) If $s \neq 1+\frac{d}{2}$

$$
\begin{equation*}
\|f\|_{H^{s}} \leq\left\|f_{0}\right\|_{H^{s}}+\int_{0}^{t}\|F(\tau)\|_{H^{s}} d \tau+C \int_{0}^{t} V^{\prime}(\tau)\|f(\tau)\|_{H^{s}} d \tau \tag{4.2}
\end{equation*}
$$

or hence,

$$
\begin{equation*}
\|f\|_{H^{s}} \leq e^{C V(t)}\left(\left\|f_{0}\right\|_{H^{s}}+\int_{0}^{t} e^{-C V(\tau)}\|F(\tau)\|_{H^{s}} d \tau\right) \tag{4.3}
\end{equation*}
$$

with $V(t)=\int_{0}^{t}\|\nabla v(\tau)\|_{H^{\frac{d}{2}} \cap L^{\infty}} d \tau$ if $s<1+\frac{d}{2}$ and $V(t)=\int_{0}^{t}\|\nabla v(\tau)\|_{H^{s-1}} d \tau$ else.
(2) If $f=v$, then for all $s>0$ the estimates (4.2) and (4.3) hold with $V(t)=\int_{0}^{t}\left\|\partial_{x} v(\tau)\right\|_{L^{\infty}} d \tau$.
Proposition 4.1.3. [6] ( $1-D$ Moser-type estimates). The following estimates hold, (1) For $s \geq 0$,

$$
\begin{equation*}
\|f g\|_{H^{s}(\mathbb{R})} \leq C\left(\|f\|_{H^{s}(\mathbb{R})}\|g\|_{L^{\infty}(\mathbb{R})}+\|f\|_{L^{\infty}(\mathbb{R})}\|g\|_{H^{s}(\mathbb{R})}\right), \tag{4.4}
\end{equation*}
$$

(2) For $s>0$,

$$
\left\|f \partial_{x} g\right\|_{H^{s}(\mathbb{R})} \leq C\left(\|f\|_{H^{s+1}(\mathbb{R})}\|g\|_{L^{\infty}(\mathbb{R})}+\|f\|_{L^{\infty}(\mathbb{R})}\left\|\partial_{x} g\right\|_{H^{s}(\mathbb{R})}\right)
$$

(3) For $s_{1} \leq \frac{1}{2}, s_{2}>\frac{1}{2}$ and $s_{1}+s_{2}>0$,

$$
\|f g\|_{H^{s_{1}(\mathbb{R})}} \leq C\|f\|_{H^{s_{1}(\mathbb{R})}}\|g\|_{H^{s_{2}}(\mathbb{R})}
$$

Proposition 4.1.4. ([14, 15]) Let $m \in \mathbb{R}$ and $f$ be an $S^{m}$-multiplier (that is, $f$ : $\mathbb{R}^{d} \rightarrow \mathbb{R}$ is smooth and satisfies that for all multi-index $\alpha$, there exists a constant $C_{\alpha}$ such that $\left.\forall \xi \in \mathbb{R}^{d},\left|\partial^{\alpha} f(\xi)\right| \leq C_{\alpha}(1+|\xi|)^{m-|\alpha|}\right)$. Then for all $s \in \mathbb{R}$ and $1 \leq p, r \leq \infty$, the operator $f(D)$ is continuous from $B_{p, r}^{s}$ to $B_{p, r}^{s-m}$.

The proof of the Theorem 4.1.1 will be given as follows.
Proof. First of all, we apply the translation $u(t, x) \rightarrow u\left(t, x-\frac{\beta_{0}}{\beta} t\right)$ to the equation

$$
\begin{equation*}
u_{t}+u u_{x}=-p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\} . \tag{3.3}
\end{equation*}
$$

Also applying $\Lambda^{s}$ operator to the equation

$$
\begin{equation*}
\partial_{t} \Lambda^{s} u+\Lambda^{s}\left(u \partial_{x} u\right)=\Lambda^{s}\left(-p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}\right) \tag{4.6}
\end{equation*}
$$

Then taking the inner product between the equation (4.6) and $\Lambda^{s} u$ in $L^{2}$, we have

$$
\begin{aligned}
& \frac{1}{2} \frac{d}{d t}\left\|\Lambda^{s} u\right\|_{L^{2}}^{2}=-\frac{1}{2} \int_{\mathbb{R}} u \partial_{x}\left(\Lambda^{s} u\right)^{2} d x-\int_{\mathbb{R}}\left[\Lambda^{s}, u\right] \partial_{x} u \Lambda^{s} u d x \\
& +\int_{\mathbb{R}} \Lambda^{s}\left(-p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}\right) \Lambda^{s} u d x \\
& \leq\left\|\partial_{x} u\right\|_{L^{\infty}}\left\|\Lambda^{s} u\right\|_{L^{2}}^{2}+\left\|\left[\Lambda^{s}, u\right] \partial_{x} u\right\|_{L^{2}}\left\|\Lambda^{s} u\right\|_{L^{2}} \\
& +\left\|\Lambda^{s}\left(-p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}\right)\right\|_{L^{2}}\left\|\Lambda^{s} u\right\|_{L^{s}}
\end{aligned}
$$

By using commutator estimate for $s>0$, we get

$$
\left\|\left[\Lambda^{s}, u\right] \partial_{x} u\right\|_{L^{2}} \leq C\left(\|u\|_{H^{s}}\left\|\partial_{x} u\right\|_{L^{\infty}}+\left\|\partial_{x} u\right\|_{L^{\infty}}\left\|\partial_{x} u\right\|_{H^{s-1}}\right) \leq C\left\|\partial_{x} u\right\|_{L^{\infty}}\|u\|_{H^{s}}
$$

Then, we have

$$
\begin{align*}
& \frac{1}{2} \frac{d}{d t}\|u\|_{H^{s}}^{2} \leq C\left(\left\|\partial_{x} u\right\|_{L^{\infty}}\|u\|_{H^{s}}\right. \\
& \left.+\left\|p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}\right\|_{H^{s}}\right)\|u\|_{H^{s}} \tag{4.7}
\end{align*}
$$

Taking integration of the equation from 0 to $t$, we get

$$
\begin{aligned}
\|u\|_{H^{s}} \leq & \left\|u_{0}\right\|_{H^{s}}+\int_{0}^{t}\|u(\tau)\|_{H^{s}}\left\|\partial_{x} u(\tau)\right\|_{L^{\infty}} d \tau+C \int_{0}^{t} \| p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u\right. \\
& \left.-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\} \|_{H^{s}} d \tau
\end{aligned}
$$

Thanks to the Moser-type estimate (4.4) and the Proposition 4.1.4, we obtain

$$
\begin{aligned}
& \left\|p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}\right\|_{H^{s}} \leq C\left(\|u\|_{H^{s-1}}\right. \\
& \left.+\left\|u_{x}\right\|_{H^{s-1}}\left\|u_{x}\right\|_{L^{\infty}}+\|u\|_{H^{s-1}}\|u\|_{L^{\infty}}+\|u\|_{H^{s-1}}\|u\|_{L^{\infty}}^{2}+\|u\|_{H^{s-1}}\|u\|_{L^{\infty}}^{3}\right)
\end{aligned}
$$

Therefore, we have

$$
\|u\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}}+C \int_{0}^{t}\|u(\tau)\|_{H^{s}}\left(1+\|u\|_{L^{\infty}}+\left\|u_{x}\right\|_{L^{\infty}}+\|u\|_{L^{\infty}}^{2}+\|u\|_{L^{\infty}}^{3}\right) d \tau .
$$

Applying the Gronwall's inequality, it follows

$$
\begin{equation*}
\|u\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}} e^{C \int_{0}^{t}\left(1+\|u\|_{L^{\infty}}+\left\|u_{x}\right\|_{L^{\infty}}+\|u\|_{L^{\infty}}^{2}+\|u\|_{L}^{3}\right) d \tau} . \tag{4.8}
\end{equation*}
$$

On the other hand, if we multiply (3.2) by $u$ and integrate over $\mathbb{R}$, we obtain

$$
\begin{gathered}
\int_{\mathbb{R}} u m_{t} d x+\int_{\mathbb{R}} u^{2} m_{x} d x+b \int_{\mathbb{R}} u u_{x} m d x+c \int_{\mathbb{R}} c u u_{x} d x-\frac{\beta_{0}}{\beta} \int_{\mathbb{R}} u u_{x x x} d x \\
+\frac{\omega_{1}}{\alpha^{2}} \int_{\mathbb{R}} u^{3} u_{x} d x+\frac{\omega_{2}}{\alpha^{3}} \int_{\mathbb{R}} u^{4} u_{x} d x=0 .
\end{gathered}
$$

Then, we have

$$
\int_{\mathbb{R}} u m_{t} d x+\int_{\mathbb{R}} u^{2} m_{x} d x+b \int_{\mathbb{R}} u u_{x} m d x=0
$$

Note that, by applying integration by parts, we have

$$
\int_{\mathbb{R}} u m_{t} d x=\frac{1}{2} \frac{d}{d t} \int_{\mathbb{R}}\left(u^{2}+\left(u_{x}\right)^{2}\right) d x
$$

and

$$
\begin{array}{rl}
\int_{\mathbb{R}} u^{2} m_{x} d x+b \int_{\mathbb{R}} & u u_{x} m d x=\int_{\mathbb{R}}\left(u^{2} u_{x}-u^{2} u_{x x x}\right) d x+b \int_{\mathbb{R}}\left(u^{2} u_{x}-u u_{x} u_{x x}\right) d x \\
& =(2-b) \int_{\mathbb{R}} u u_{x} u_{x x} d x \\
& =\frac{b-2}{2} \int_{\mathbb{R}}\left(u_{x}\right)^{3} d x .
\end{array}
$$

Therefore, we have

$$
\begin{equation*}
\frac{1}{2} \frac{d}{d t} \int_{\mathbb{R}}\left(u^{2}+\left(u_{x}\right)^{2}\right) d x=\frac{2-b}{2} \int_{\mathbb{R}} u_{x}\left(u_{x}\right)^{2} d x \tag{4.9}
\end{equation*}
$$

Then, it follows

$$
\begin{gathered}
\frac{d}{d t} \int_{\mathbb{R}}\left(u^{2}+\left(u_{x}\right)^{2}\right) d x \leq|b-2|\left\|u_{x}\right\|_{L^{\infty}} \int_{\mathbb{R}}\left(u_{x}\right)^{2} d x \\
\leq|b-2|\left\|u_{x}\right\|_{L^{\infty}} \int_{\mathbb{R}}\left(u^{2}+\left(u_{x}\right)^{2}\right) d x
\end{gathered}
$$

$$
\frac{d}{d t}\|u\|_{H^{1}}^{2} \leq|b-2|\left\|u_{x}\right\|_{L^{\infty}}\|u\|_{H^{1}}^{2}
$$

Thanks to the Gronwall inequality, we get

$$
\|u\|_{H^{1}} \leq e^{|b-2| \int_{0}^{T}\left\|u_{x}\right\|_{L^{\infty}} d t}\left\|u_{0}\right\|_{H^{1}}
$$

By using the following Sobolev inequality

$$
\|u\|_{L^{\infty}} \leq \frac{\|u\|_{H_{1}}}{\sqrt{2}}
$$

we have

$$
\begin{equation*}
\|u\|_{L^{\infty}} \leq \frac{1}{\sqrt{2}}\|u\|_{H^{1}} \leq \frac{1}{\sqrt{2}} e^{|b-2| \int_{0}^{T}\left\|u_{x}\right\|_{L^{\infty}} d t}\left\|u_{0}\right\|_{H^{1}} . \tag{4.10}
\end{equation*}
$$

Then if the maximal existence time $T_{u_{0}}^{*}<\infty$ satisfies $\int_{0}^{T_{u_{0}}^{*}}\left\|\partial_{x} u(\tau)\right\|_{L^{\infty}} d \tau<\infty$, we have from (4.8) and (4.10)

$$
\|u\|_{L^{\infty}} \leq C(T)
$$

and

$$
\|u\|_{H^{s}} \leq\left\|u_{0}\right\|_{H^{s}} e^{C \int_{0}^{t} C(T) d \tau}<\infty
$$

where $C(T)$ is any positive constant depending on $T$. Therefore, we conclude

$$
\lim \sup _{t \rightarrow T_{u_{0}}^{*}}\|u(t)\|_{H^{s}}<\infty
$$

which contradicts the assumption on the maximal existence time $T_{u_{0}}^{*}<\infty$. This completes the proof of Theorem 4.1.1

Theorem 4.1.5. (Wave-breaking criterion) Assume $u_{0} \in H^{s}(\mathbb{R})$ with $s>\frac{3}{2}$. Let $T_{u_{0}}^{*}>0$ be maximal existence time of the corresponding solution $u(t, x)$ to the system (3.6). Then the corresponding solution breaks down in the following two cases,
(i) for $b \geq 2$, if and only if

$$
\begin{equation*}
\liminf _{t \uparrow T_{u_{0}}^{*}} \inf _{x \in \mathbb{R}} u_{x}(t, x)=-\infty \tag{4.11}
\end{equation*}
$$

(ii) for $b \leq 1$, if and only if

$$
\begin{equation*}
\limsup _{t \uparrow T_{u_{0}}^{*}} \sup _{x \in \mathbb{R}} u_{x}(t, x)=+\infty . \tag{4.12}
\end{equation*}
$$

Proof. We have the following equation by differentiating the equation (4.5) with respect to $x$ and considering $\partial_{x}^{2} p * f=p * f-f$

$$
\begin{align*}
u_{t x}+u_{x}^{2}+u u_{x x} & =\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4} \\
& -p *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\} . \tag{4.13}
\end{align*}
$$

Also, we consider the following differential equation

$$
\begin{cases}\frac{\partial q}{\partial t}=u(t, q), & 0<t<T  \tag{4.14}\\ q(0, x)=x, & x \in \mathbb{R}\end{cases}
$$

where $u \in C^{1}\left([0, T), H^{s-1}\right)$ is the solution to equation (4.5) with initial data $u_{0} \in H^{s}$, for $s>3 / 2$. Differentiating (4.14) with respect to $x$, we obtain

$$
\begin{align*}
\frac{d}{d t} u_{x}(t, q(t, x)) & =u_{x t}(t, q(t, x))+u_{x x}(t, q(t, x)) q_{t}(t, x)  \tag{4.15}\\
& =\left(u_{t x}+u u_{x x}\right)(t, q(t, x))
\end{align*}
$$

From (4.13) and (4.15), we achieve

$$
\begin{align*}
\frac{d}{d t} u_{x}(t, q(t, x)) & =-\frac{b-1}{2} u_{x}^{2}+\left(c-\frac{\beta_{0}}{\beta}\right) u+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4} \\
& -p *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u-\frac{b-3}{2}\left(u_{x}\right)^{2}+\frac{b}{2} u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\} . \tag{4.16}
\end{align*}
$$

Considering (4.9) and if $b \geq 2$, we have

$$
\begin{aligned}
& \frac{d}{d t}\|u\|_{H^{1}}^{2} \leq(b-2)\left(-\inf _{x \in \mathbb{R}} u_{x}\right) \int_{\mathbb{R}} u_{x}^{2} d x \\
& \leq(b-2)\left(-\inf _{x \in \mathbb{R}} u_{x}\right) \int_{\mathbb{R}} u^{2}+u_{x}^{2} d x \\
&=(b-2)\left(-\inf _{x \in \mathbb{R}} u_{x}\right)\|u\|_{H^{1}}^{2} .
\end{aligned}
$$

Applying Gronwall inequality, we get

$$
\|u\|_{H^{1}} \leq e^{(b-2) \int_{0}^{T}\left(-\inf _{x \in \mathbb{R}} u_{x}\right) d t}\left\|u_{0}\right\|_{H^{1}}
$$

Assume that $T_{u_{0}}^{*}<\infty$ and there exists $M>0$ such that

$$
u_{x}(t, x) \geq-M, \forall(t, x) \in\left[0, T_{u_{0}}^{*}\right) \times \mathbb{R}
$$

Then, we have

$$
\|u\|_{L^{\infty}} \leq \frac{1}{\sqrt{2}}\|u\|_{H^{1}} \leq e^{(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}} \leq C
$$

where $C>0$ is a constant, depending on $T_{u_{0}}^{*}, M$ and $\left\|u_{0}\right\|_{H^{1}}$. Considering the following inequalities such as $u_{x}{ }^{2} \geq 0, p * u^{2} \geq 0,\|p * u\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{\infty}} \leq e^{(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}$,

$$
\begin{aligned}
& \left\|p * u_{x}^{2}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\left\|u_{x}\right\|_{L^{2}}^{2} \leq e^{2(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{2}, \\
& \left\|p * u^{2}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{2}}^{2} \leq e^{2(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{2}, \\
& \left\|p * u^{3}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{\infty}}^{3} \leq e^{3(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{3}, \\
& \left\|p * u^{4}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{2}}^{4} \leq e^{4(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{4},
\end{aligned}
$$

then from (4.16), we have

$$
\begin{aligned}
\frac{d}{d t} u_{x}(t, q(t, x)) & \leq C^{\prime}\left(e^{(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}+e^{2(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{2}\right. \\
& \left.+e^{3(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{3}+e^{4(b-2) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{4}\right) \\
& \leq C\left(T_{u_{0}}^{*}, M\right) .
\end{aligned}
$$

where $C^{\prime}$ is a positive constant. Integrating the inequality with respect to $t<T_{u_{0}}^{*}$ on $[0, t]$, we obtain

$$
u_{x}(t, q(t, x)) \leq u_{x}(0)+C\left(T_{u_{0}}^{*}, M\right) t
$$

Then, for $\forall t \in\left[0, T_{u_{0}}^{*}\right)$, we have

$$
\sup _{x \in \mathbb{R}} u_{x}(t, x) \leq\left\|\partial_{x} u_{0}\right\|_{L^{\infty}}+C t \leq\left\|u_{0}\right\|_{H^{1}}+C t
$$

which together with

$$
\inf _{x \in \mathbb{R}} u_{x} \geq-M, \quad \forall(t, x) \in\left[0, T_{u_{0}}^{*}\right) \times \mathbb{R}
$$

and $T_{u_{0}}^{*}<\infty$ implies that

$$
\int_{0}^{T_{u_{0}}^{*}}\left\|\partial u_{x}(\tau)\right\|_{L^{\infty}} d \tau<\infty
$$

This contradicts Theorem 4.1.1. On the contrary, by using the Sobolev embedding theorem, $H^{s}(\mathbb{R}) \hookrightarrow L^{\infty}(\mathbb{R})$, (with $s>\frac{1}{2}$ ) if

$$
\liminf _{t \uparrow T_{u_{0}}} \inf _{x \in \mathbb{R}} u_{x}(t, x)=-\infty
$$

holds, then the solution $u$ blows up in finite time,
On the other hand, suppose that $\sup _{x \in \mathbb{R}} u_{x} \geq 0$ if $b \leq 1$, we have

$$
\begin{aligned}
\frac{d}{d t}\|u\|_{H^{1}}^{2} \leq(2 & -b)\left(\sup _{x \in \mathbb{R}} u_{x}\right) \int_{\mathbb{R}} u_{x}{ }^{2} d x \\
& \leq(2-b)\left(\sup _{x \in \mathbb{R}} u_{x}\right) \int_{\mathbb{R}} u^{2}+u_{x}{ }^{2} d x \\
& \leq(2-b)\left(\sup _{x \in \mathbb{R}} u_{x}\right)\|u\|_{H^{1}}^{2} .
\end{aligned}
$$

Applying Gronwall inequality, it follows

$$
\|u\|_{H^{1}} \leq e^{(2-b) \int_{0}^{T}\left(\sup _{x \in \mathbb{R}} u_{x}\right) d t}\left\|u_{0}\right\|_{H^{1}}
$$

Assume that $T_{u_{0}}^{*}<\infty$ and there exists $M>0$ such that

$$
u_{x}(t, x)<M, \forall(t, x) \in\left[0, T_{u_{0}}^{*}\right) \times \mathbb{R}
$$

Then, we have

$$
\|u\|_{L^{\infty}} \leq \frac{1}{\sqrt{2}}\|u\|_{H^{1}} \leq e^{(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}} \leq C
$$

where $C>0$ is a constant, depending on $T_{u_{0}}^{*}, M$ and $\left\|u_{0}\right\|_{H^{1}}$. Considering the following inequalities such as $u_{x}{ }^{2} \geq 0, p * u^{2} \geq 0,\|p * u\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{\infty}} \leq e^{(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}$,

$$
\begin{aligned}
& \left\|p * u_{x}^{2}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\left\|u_{x}\right\|_{L^{2}}^{2} \leq e^{2(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{2} \\
& \left\|p * u^{2}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{2}}^{2} \leq e^{2(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{2} \\
& \left\|p * u^{3}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{\infty}}^{3} \leq e^{3(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{3}, \\
& \left\|p * u^{4}\right\|_{L^{\infty}} \leq\|p\|_{L^{1}}\|u\|_{L^{2}}^{4} \leq e^{4(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{4},
\end{aligned}
$$

then from (4.16), we have

$$
\begin{aligned}
\frac{d}{d t} u_{x}(t, q(t, x)) & \geq-C^{\prime}\left(e^{(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}+e^{2(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{2}\right. \\
& \left.+e^{3(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{3}+e^{4(2-b) M T_{u_{0}}^{*}}\left\|u_{0}\right\|_{H^{1}}^{4}\right) \\
& \geq-C\left(T_{u_{0}}^{*}, M\right)
\end{aligned}
$$

where $C^{\prime}$ is a positive constant. Applying absolute value to this inequality, we get

$$
-C\left(T_{u_{0}}^{*}, M\right) \leq \frac{d}{d t} u_{x}(t, q(t, x)) \leq C\left(T_{u_{0}}^{*}, M\right)
$$

So, by integrating the inequality with respect to $t<T_{u_{0}}^{*}$ on $[0, t]$, we obtain

$$
-C\left(T_{u_{0}}^{*}, M\right) t+u_{x}(0) \leq \inf _{x \in \mathbb{R}} u_{x}(t, q(t, x))
$$

Then, $\forall t \in\left[0, T_{u_{0}}^{*}\right)$, we have

$$
-C\left(T_{u_{0}}^{*}, M\right) t-\left\|u_{0}\right\|_{H^{1}} \leq \inf _{x \in \mathbb{R}} u_{x}(t, q(t, x))
$$

which together with

$$
\sup _{x \in \mathbb{R}} u_{x} \leq M, \quad \forall(t, x) \in\left[0, T_{u_{0}}^{*}\right) \times \mathbb{R},
$$

and $T_{u_{0}}^{*}<\infty$ implies that

$$
\int_{0}^{T_{u_{0}}^{*}}\left\|\partial u_{x}(\tau)\right\|_{L^{\infty}} d \tau<\infty
$$

This contradicts Theorem 4.1.1. On the contrary, by using the Sobolev embedding theorem, $H^{s}(\mathbb{R}) \hookrightarrow L^{\infty}(\mathbb{R})$, (with $s>\frac{1}{2}$ ) if

$$
\limsup _{t \uparrow T_{u_{0}}^{*}} \sup _{x \in \mathbb{R}} u_{x}(t, x)=+\infty
$$

holds, then the solution $u$ blows up in finite time, which completes proof.

### 4.2 Wave-Breaking Phenomena

In this section, we consider the special case of the rotational $b$-family of equations which is called Rotational Camassa-Holm equation $(b=2)$

$$
\begin{equation*}
u_{t}+u u_{x}+\frac{\beta_{0}}{\beta} u_{x}+p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u+\frac{1}{2}\left(u_{x}\right)^{2}+u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}=0 \tag{4.17}
\end{equation*}
$$

where $p=\frac{1}{2} e^{-|x|}$. We state the following theorem to find the wave breaking data for the $\mathrm{R}-\mathrm{CH}$.

Theorem 4.2.1. Assume that $u \in C\left([0, T) ; H^{s}(\mathbb{R})\right) \cap C^{1}\left([0, T) ; H^{s-1}(\mathbb{R})\right)$, $s>3 / 2$ is a solution of the Cauchy problem (3.6) for $b=2$ with the initial value $u_{0} \in H^{s}(\mathbb{R})$ ). Suppose that

$$
m_{0}<\min \{A, B\}
$$

where

$$
\begin{aligned}
& A=-2\left(\frac{|\gamma|}{2}+\sqrt{\frac{\gamma^{2}}{4}+\frac{1}{8}\left\|u_{0}\right\|_{H^{1}}^{2}+\frac{1}{2} C_{0}^{2}}\right) \\
& B=-\left(\frac{|\gamma|}{2}+\sqrt{\frac{\gamma^{2}}{4}+2\left(\frac{|\gamma|}{2} M_{0}+\frac{1}{4}\left\|u_{0}\right\|_{H^{1}}^{2}+C_{0}^{2}\right)}\right) \\
& \gamma=c-\frac{\beta_{0}}{\beta}
\end{aligned}
$$

and $C_{0}>0$ is defined by

$$
C_{0}^{2}=\frac{2\left|\omega_{1}\right|}{3 \alpha^{2}} E_{0}^{3 / 2}+\frac{\left|\omega_{2}\right|}{2 \alpha^{3}} E_{0}^{2}
$$

and

$$
E_{0}=\frac{1}{2} \int_{\mathbb{R}}\left(u_{0}^{2}+\left(\partial_{x} u_{0}\right)^{2}\right) d x
$$

The corresponding solution of (3.6) for $b=2$ breaks down in finite time $T_{0}$ with an estimate as

$$
T_{0} \leq t^{*}=\frac{4 m_{0}^{2}+4|\gamma| m_{0}-2\left\|u_{0}\right\|_{H^{1}}^{2}-2 C_{0}^{2}}{\left(m_{0}^{2}+2|\gamma| m_{0}-\frac{1}{2}\left\|u_{0}\right\|_{H^{1}}^{2}-2 C_{0}^{2}\right) \sqrt{m_{0}^{2}-|\gamma|\left(M_{0}-m_{0}\right)-\frac{1}{2}\left\|u_{0}\right\|_{H^{1}}^{2}-2 C_{0}^{2}}} .
$$

We will give the proof for this theorem by using the following proposition and lemma.

Proposition 4.2.2. ([35]) Let $m(t)=\inf _{x \in \mathbb{R}}\left\{u_{x}(t, x)\right\}$ and $M(t)=\sup _{x \in \mathbb{R}}\left\{u_{x}(t, x)\right\}$ be two continuous and almost everywhere differentiable defined in $t \in[0, T) T \leq \infty$ with satisfying

$$
\left\{\begin{array}{l}
\frac{d m}{d t} \leq-a^{\prime} m^{2}(t)+b^{\prime}[M(t)-m(t)]+c^{\prime}, \\
\frac{d M}{d t} \leq-a^{\prime} M^{2}(t)+b^{\prime}[M(t)-m(t)]+c^{\prime},
\end{array} ; \text { a.e. in } \quad t \in[0, T)\right.
$$

where $a^{\prime}$ is a positive constant, $b^{\prime}$ and $c^{\prime}$ are non-negative constants, and $M(t)$ is a nonnegative function of $t$. Suppose that the initial data $m_{0}=m(0)$ and $M_{0}=M(0)$ satisfy

$$
m_{0}<\min \left\{-\frac{1}{a^{\prime}}\left(b^{\prime}+\sqrt{\left(b^{\prime}\right)^{2}+a^{\prime} c^{\prime}}\right),-\frac{1}{2 a^{\prime}}\left(b^{\prime}+\sqrt{\left.\left(b^{\prime}\right)^{2}+4 a^{\prime}\left(b^{\prime} M_{0}+c^{\prime}\right)\right)}\right\} .\right.
$$

Then $m(t)$ is monotonically decreasing and breaks down in the finite time $T_{0}$ with

$$
T_{0} \leq t^{*}=\frac{a^{\prime} m_{0}^{2}+b^{\prime} m_{0}-c}{a^{\prime}\left(a^{\prime} m_{0}^{2}+2 b^{\prime} m_{0}-c^{\prime}\right) \sqrt{a^{\prime}\left(a^{\prime} m_{0}^{2}-b^{\prime}\left(M_{0}-m_{0}\right)-c^{\prime}\right)}}
$$

in the sense that

$$
\lim \inf _{t \rightarrow T_{0}^{-}} m(t)=-\infty
$$

In the case of $T_{0}=t^{*}$, the wave-breaking rate can be estimated by

$$
m(t) \leq \frac{a^{\prime} m_{0}^{2}+b^{\prime} m_{0}-c^{\prime}}{a^{\prime} m_{0}^{2}+2 b^{\prime} m_{0}-c^{\prime}} \frac{1}{t-t^{*}}
$$

Furthermore, if $m(t)$ is bounded below by some negative constant $m_{l}$, i.e. $m(t) \geq m_{l}$, then $M(t)$ is bounded by

$$
M(t) \leq \max \left\{M_{0}, \frac{b^{\prime}+\sqrt{b^{\prime 2}+4 a^{\prime}\left(c^{\prime}-b^{\prime} m_{l}\right)}}{2 a^{\prime}}\right\}
$$

Lemma 4.2.3. ([8] Let $T>0$ and $v \in C^{1}\left([0, T) ; H^{2}(\mathbb{R})\right)$. Then for every $t \in[0, T)$ there exists at least one point $\xi \in \mathbb{R}$ with

$$
\begin{equation*}
m(t):=\inf _{x \in \mathbb{R}}\left[v_{x}(t, x)\right]=v_{x}(t, \xi(t)), \tag{4.18}
\end{equation*}
$$

and the function $m$ is almost everywhere differentiable on $(0, T)$ with

$$
\begin{equation*}
\frac{d m}{d t}(t)=v_{t x}(t, \xi(t)) \quad \text { a.e. } \quad \text { on } \quad(0, T) \tag{4.19}
\end{equation*}
$$

Now, we can prove the Theorem 4.2.1.
Proof. First, we apply the translation $u(t, x) \rightarrow u\left(t, x-\frac{\beta_{0}}{\beta} t\right)$ to the equation (4.17). Then we get the equation in the form, as follows

$$
\begin{equation*}
u_{t}+u u_{x}+p_{x} *\left\{\left(c-\frac{\beta_{0}}{\beta}\right) u+\frac{1}{2}\left(u_{x}\right)^{2}+u^{2}+\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right\}=0 \tag{4.20}
\end{equation*}
$$

Differentiating the equation with respect to $x$, it yields

$$
\begin{aligned}
& u_{x t}+u u_{x x}+u_{x}^{2}+\gamma \partial_{x} p *\left(u_{x}\right)+p *\left(u^{2}+\frac{1}{2}\left(u_{x}\right)^{2}\right)-\left(u^{2}+\frac{1}{2}\left(u_{x}\right)^{2}\right) \\
& +p *\left(\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right)-\left(\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right)=0
\end{aligned}
$$

where $\gamma=\left(c-\frac{\beta_{0}}{\beta}\right)$. Using the Lemma 3.1 in [2], we get the following

$$
\begin{gathered}
p *\left(u^{2}+\frac{1}{2} u_{x}^{2}\right) \geq \frac{1}{2} u^{2} . \\
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\end{gathered}
$$

On the other hand, we have

$$
\begin{aligned}
\left\lvert\,\left(\frac{\omega_{1}}{3 \alpha^{2}} u^{3}+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right)-p *\left(\frac{\omega_{1}}{3 \alpha^{2}} u^{3}\right.\right. & \left.+\frac{\omega_{2}}{4 \alpha^{3}} u^{4}\right) \left\lvert\, \leq \frac{\left|\omega_{1}\right|}{3 \alpha^{2}}\|u\|_{L^{\infty}}^{3}+\frac{\left|\omega_{2}\right|}{4 \alpha^{3}}\|u\|_{L^{\infty}}^{4}\right. \\
& +\frac{1}{2} \frac{\left|\omega_{1}\right|}{3 \alpha^{2}}\|u\|_{L^{\infty}}\|u\|_{L^{2}}^{2}+\frac{1}{2} \frac{\left|\omega_{2}\right|}{4 \alpha^{3}}\|u\|_{L^{\infty}}^{2}\|u\|_{L^{2}}^{2} \\
& \leq \frac{2\left|\omega_{1}\right|}{3 \alpha^{2}} E_{0}^{\frac{3}{2}}+\frac{\left|\omega_{2}\right|}{2 \alpha^{3}} E_{0}^{2}=C_{0}^{2}>0 .
\end{aligned}
$$

Then, we obtain

$$
\begin{align*}
u_{x t} & +u u_{x x} \leq-\frac{1}{2} u_{x}^{2}-\gamma \partial_{x} p *\left(u_{x}\right)+\frac{1}{2} u^{2}+C_{0}^{2}  \tag{4.21}\\
& \leq-\frac{1}{2} u_{x}^{2}-\gamma \partial_{x} p *\left(u_{x}\right)+\frac{1}{4}\left\|u_{0}\right\|_{H^{1}}^{2}+C_{0}^{2}
\end{align*}
$$

Now, let estimate $\partial_{x} p *\left(u_{x}\right)$, first of all define for $t \in[0, T)$,

$$
\begin{align*}
m(t) & :=\inf _{\mathbb{R}}\left[u_{x}(t, x)\right]=u_{x}\left(t, \xi_{1}(t)\right)  \tag{4.22}\\
M(t) & :=\sup _{\mathbb{R}}\left[u_{x}(t, x)\right]=u_{x}\left(t, \xi_{2}(t)\right)
\end{align*}
$$

where $\xi_{i}, i=1,2$, are some points in $\mathbb{R}$; see Theorem 2.1 in [8] for the existence of $\xi_{1}(t), \xi_{2}(t)$. By Lebesgue's dominated convergence theorem, we have that

$$
\begin{equation*}
\lim _{n \rightarrow \infty} \int_{-n}^{n} p(\eta) u_{x x}(t, \xi(t)-\eta) d \eta=\int_{\mathbb{R}} p(\eta) u_{x x}(t, \xi(t)-\eta) d \eta \tag{4.23}
\end{equation*}
$$

If $[a, b] \subset \mathbb{R}$ is an interval where p is monotone and $f:[a, b] \rightarrow \mathbb{R}$ is continuous, by the second mean-value theorem there is some $c \in[a, b]$ such that

$$
\begin{equation*}
\int_{a}^{b} p(x) f(x) d x=p(a) \int_{a}^{c} f(x) d x+p(b) \int_{c}^{b} f(x) d x \tag{4.24}
\end{equation*}
$$

Therefore, we get for $n \geq 1$ and $\mathrm{i}=1,2$, points $c \in[-n, 0], d \in[0, n]$ such that

$$
\begin{align*}
\int_{-n}^{0} \frac{1}{2} e^{\eta} u_{x x}\left(t, \xi_{i}(t)-\eta\right) d \eta= & -\frac{1}{2} e^{-n}\left[u_{x}\left(t, \xi_{i}(t)-c\right)-u_{x}\left(t, \xi_{i}(t)+n\right)\right]  \tag{4.25}\\
& -\frac{1}{2}\left[u_{x}\left(t, \xi_{i}(t)\right)-u_{x}\left(t, \xi_{i}(t)-c\right)\right] \\
\int_{0}^{n} \frac{1}{2} e^{-\eta} u_{x x}\left(t, \xi_{i}(t)-\eta\right) d \eta= & -\frac{1}{2} e^{-n}\left[u_{x}\left(t, \xi_{i}(t)-n\right)-u_{x}\left(t, \xi_{i}(t)-d\right)\right]  \tag{4.26}\\
& -\frac{1}{2}\left[u_{x}\left(t, \xi_{i}(t)-d\right)-u_{x}\left(t, \xi_{i}(t)\right)\right]
\end{align*}
$$

respectively. Recalling definition of $m(t)$ and $M(t)$, we deduce by adding (4.25) and (4.26) that for $n \geq 1$

$$
\begin{gathered}
\int_{-n}^{n} \frac{1}{2} e^{\eta} u_{x x}\left(t, \xi_{i}(t)-\eta\right) d \eta=-\frac{1}{2} e^{-n}\left[u_{x}\left(t, \xi_{i}(t)-c\right)-u_{x}\left(t, \xi_{i}(t)+n\right)+u_{x}\left(t, \xi_{i}(t)-n\right)\right. \\
\left.-u_{x}\left(t, \xi_{i}(t)-d\right)\right]-\frac{1}{2}\left[u_{x}\left(t, \xi_{i}(t)-d\right)-u_{x}\left(t, \xi_{i}(t)-c\right)\right] \\
\leq \frac{M(t)-m(t)}{2}+e^{-n}[M(t)-m(t)]
\end{gathered}
$$

Thus, on account of (4.23), we obtain the estimate

$$
\begin{aligned}
& \partial_{x} p *\left(u_{x}\right)=p *\left(u_{x x}\right)=\frac{1}{2} \int_{\mathbb{R}} e^{-|x-\xi|} u_{x x}(t, \xi) d \xi \\
& =\int_{\mathbb{R}} \frac{1}{2} e^{\eta} u_{x x}\left(t, \xi_{i}(t)-\eta\right) d \eta \leq \frac{M(t)-m(t)}{2}, t \in(0, T), \text { for } i=1,2
\end{aligned}
$$

Then it follows from (4.21) and by Lemma (4.2.3), $u_{x x}(t, \xi(t))=0$ since $m(t)$ is minimum for $u_{x}(t, \cdot) \in C^{2}$ that $m(t)$ and $M(t)$ defined by (4.22) satisfy

$$
\begin{aligned}
\frac{d m}{d t} & \leq-\frac{1}{2} m^{2}+\frac{|\gamma|}{2}(M-m)+\frac{1}{4}\left\|u_{0}\right\|_{H^{1}}^{2}+C_{0}^{2} \\
\frac{d M}{d t} & \leq-\frac{1}{2} M^{2}+\frac{|\gamma|}{2}(M-m)+\frac{1}{4}\left\|u_{0}\right\|_{H^{1}}^{2}+C_{0}^{2}
\end{aligned}
$$

Applying Proposition 4.2.2, we complete the proof.

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